

Changing-look AGN in the BAT AGN Spectroscopic Survey

[arXiv:2211.04478](https://arxiv.org/abs/2211.04478)

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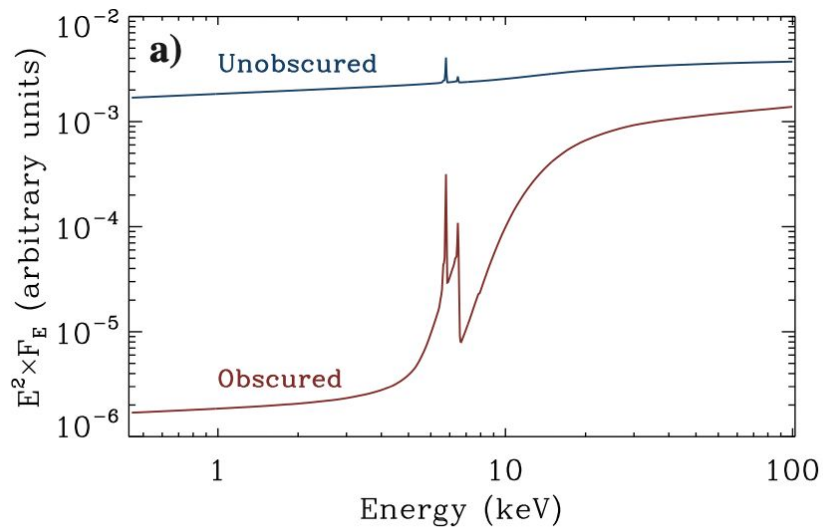
And the BASS collaboration:

Claudio Ricci, Mike Koss, Benny Trakhtenbrot, Franz Bauer, Alejandra Rojas etc.

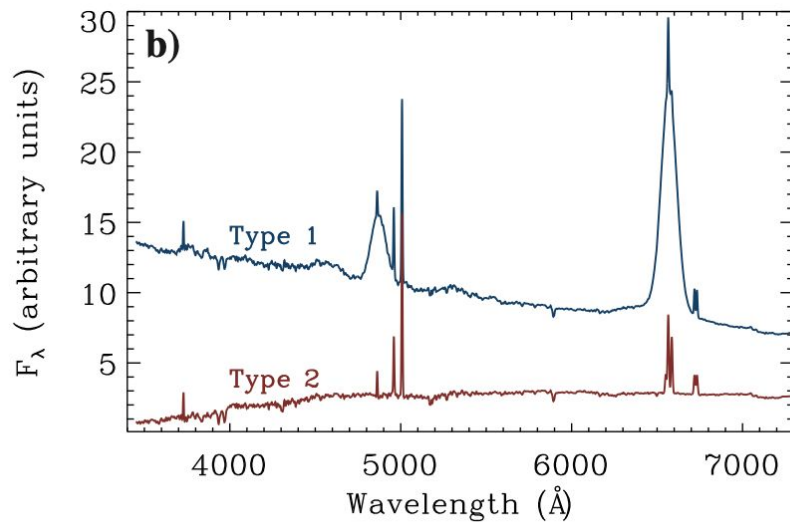
Changing-look AGN

Spectral classes identified in:

X-rays



UV/optical



Changing-look AGN: some open questions

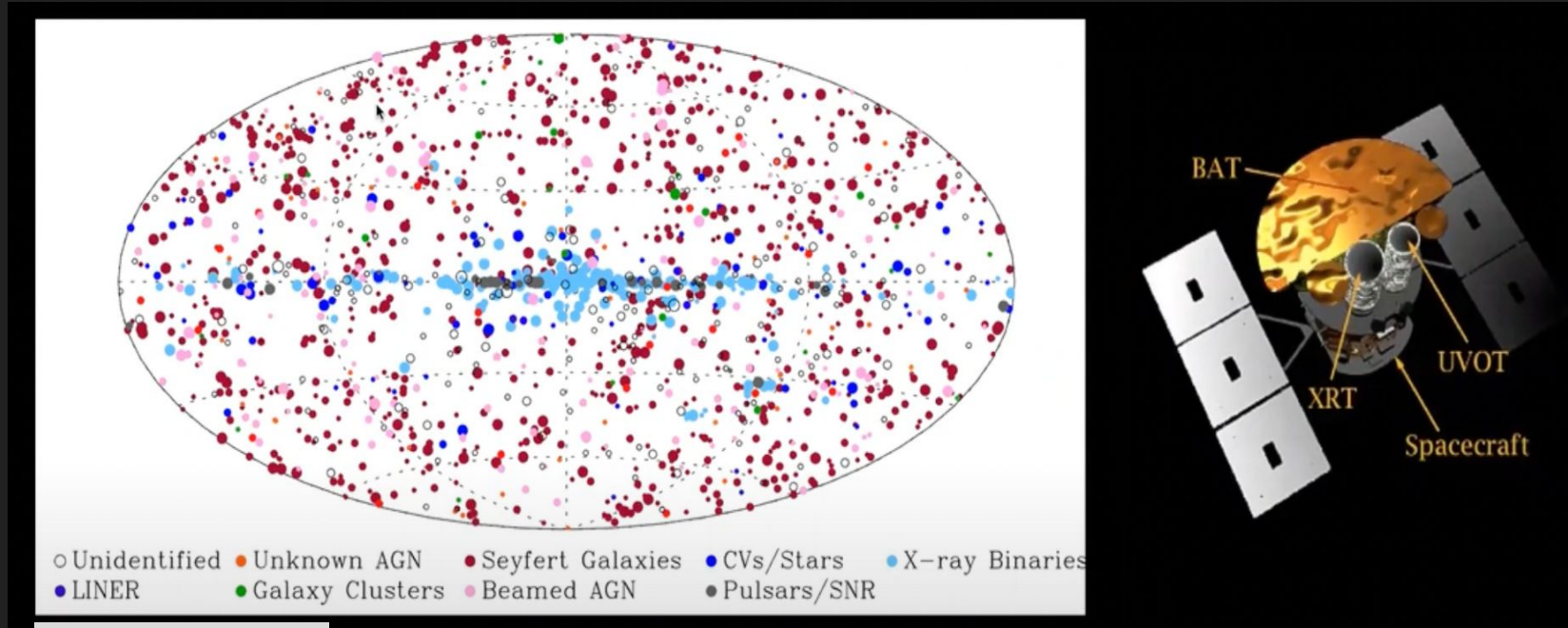
Changes in obscuration, or in accretion flow, or both (or neither)?

How common are CL events?

What are the timescales of changing state events?

How does the BLR link to the accretion flow? What drives CS events?

858 AGN from 70 months of Swift-BAT operations

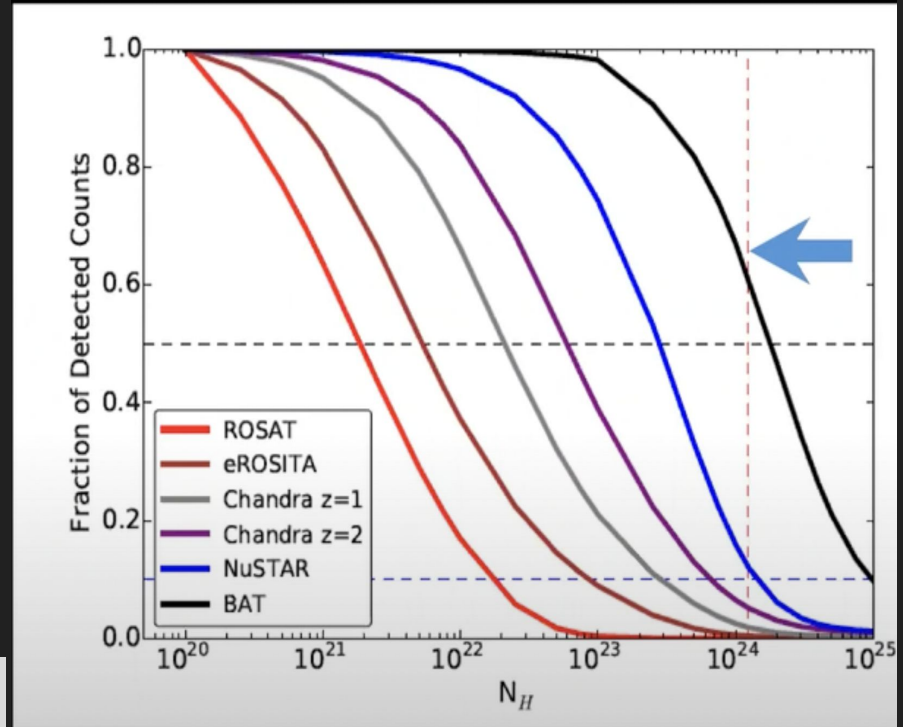


Oh et al. 2018

858 AGN from 70 months of Swift-BAT operations

Hard (14-195 keV) X-ray selection from Swift-BAT is sensitive to AGN up to $N_H \sim 10^{24}$

Measure of coronal X-ray emission gives clean probe of accretion power

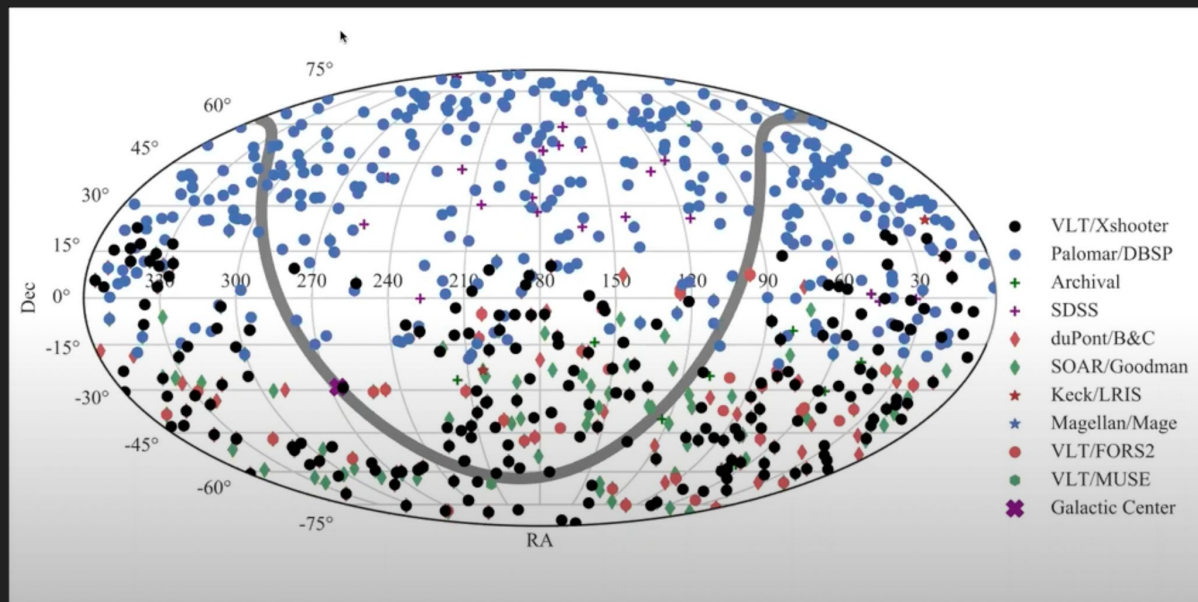


Koss et al. 2016b

BAT AGN Spectroscopic Survey (BASS)

Multi-wavelength follow-up of
858 BAT detected sources


- DR2: 1449 spectra
- M_{BH} , L/L_{Edd} , N_H etc.



Koss et al. 2022a

Searching for CL AGN in BASS

412 BAT AGN with
>1 optical spectrum

Change in Broad
Balmer lines? 

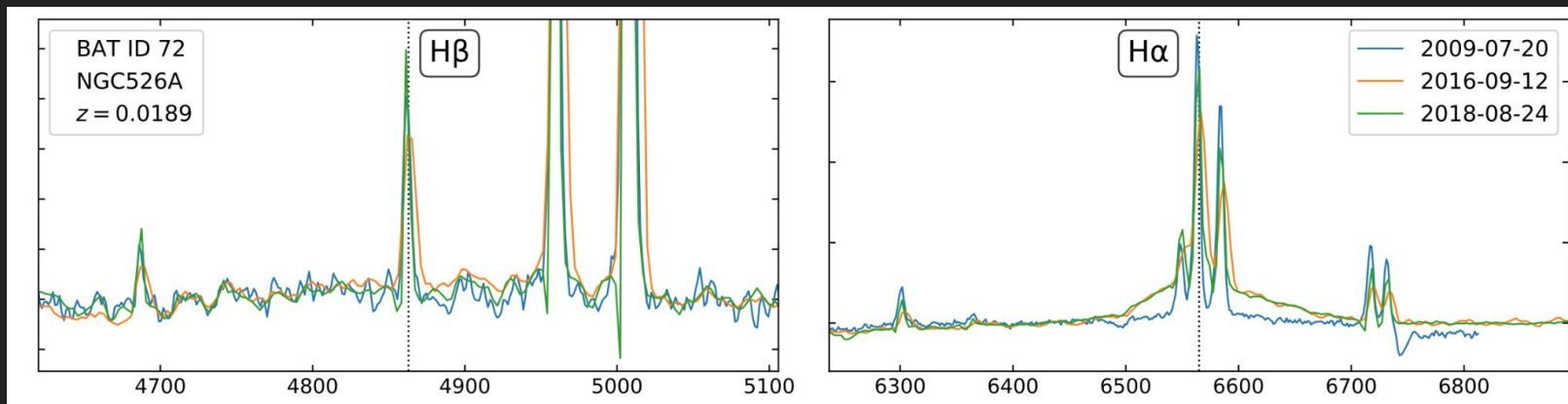
$$z < 0.5$$

$$L_{\text{Bol}} \sim 10^{42-47} \text{ erg/s}$$

$$M_{\text{BH}} \sim 10^{6-9.5} M_{\text{sol}}$$

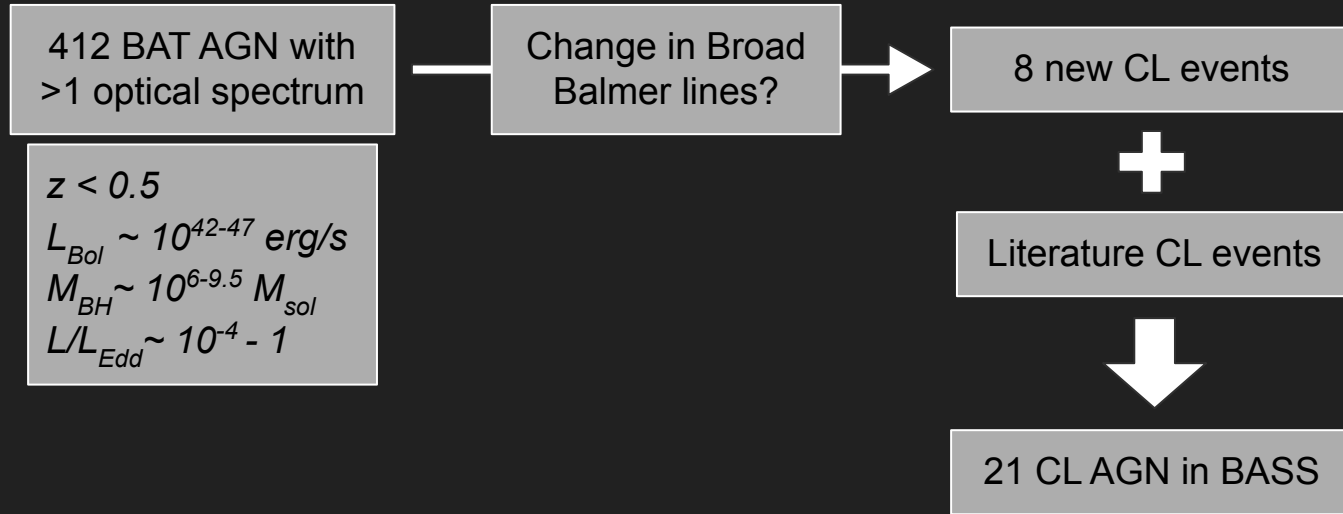
$$L/L_{\text{Edd}} \sim 10^{-4} - 1$$

Searching for CL AGN in BASS



Temple et al. 2023a

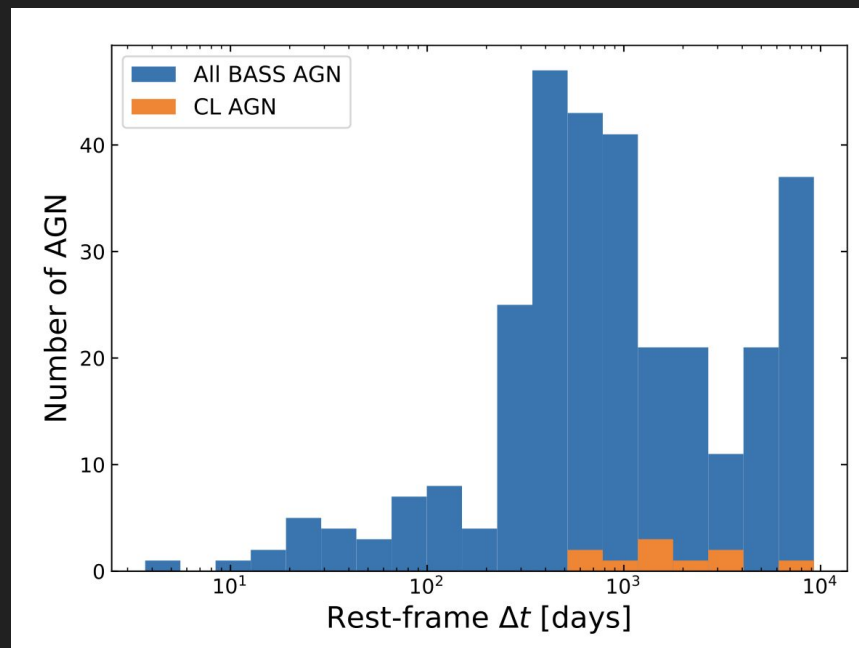
Searching for CL AGN in BASS



CL AGN timescales and rates

We find $\sim 3\%$ (0.7-6.2%) of local AGN displaying complete BL-NL (or NL-BL) changes on $\sim 15\text{yr}$ timescales

CL events are temporally unresolved - potentially missing fast events which turn both off+on within $\sim 1\text{yr}$



Temple et al. 2023a

CL AGN timescales and rates

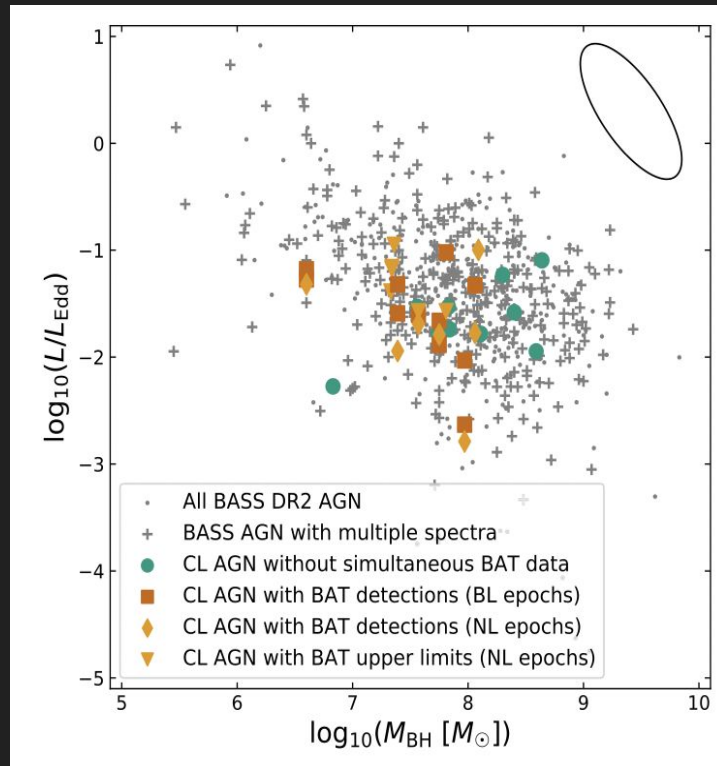
We find $\sim 3\%$ (0.7-6.2%) of local AGN displaying complete BL-NL (or NL-BL) changes on ~ 15 yr timescales

=> a relatively common phenomenon

CL rates much higher than in SDSS quasars

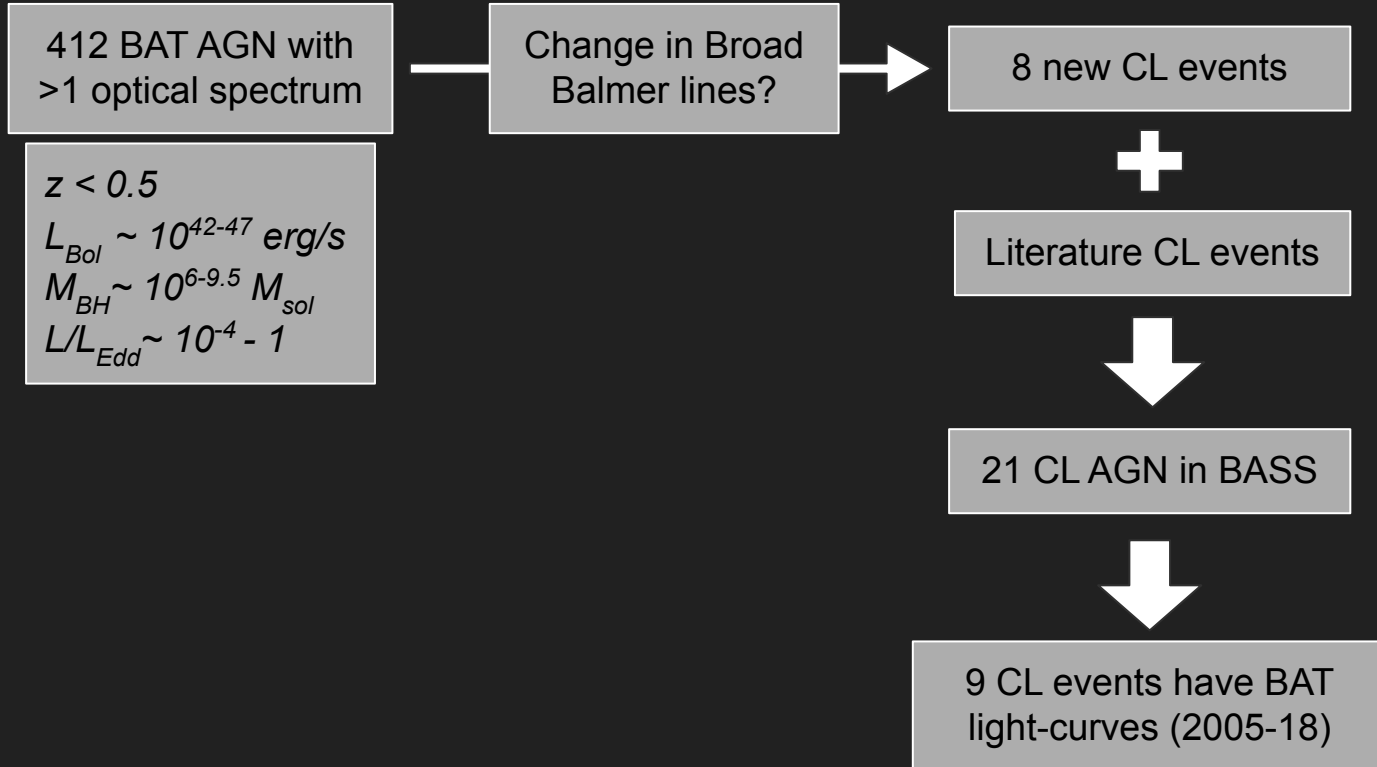
=> consistent with most CL events

occurring at lower L/L_{Edd}



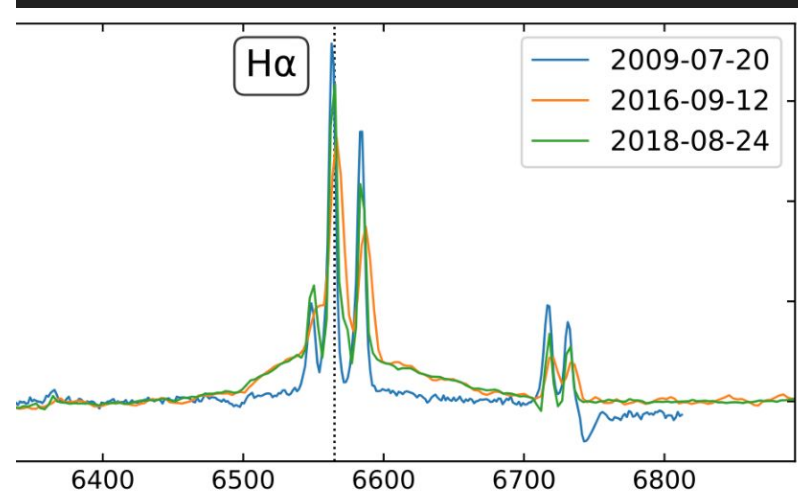
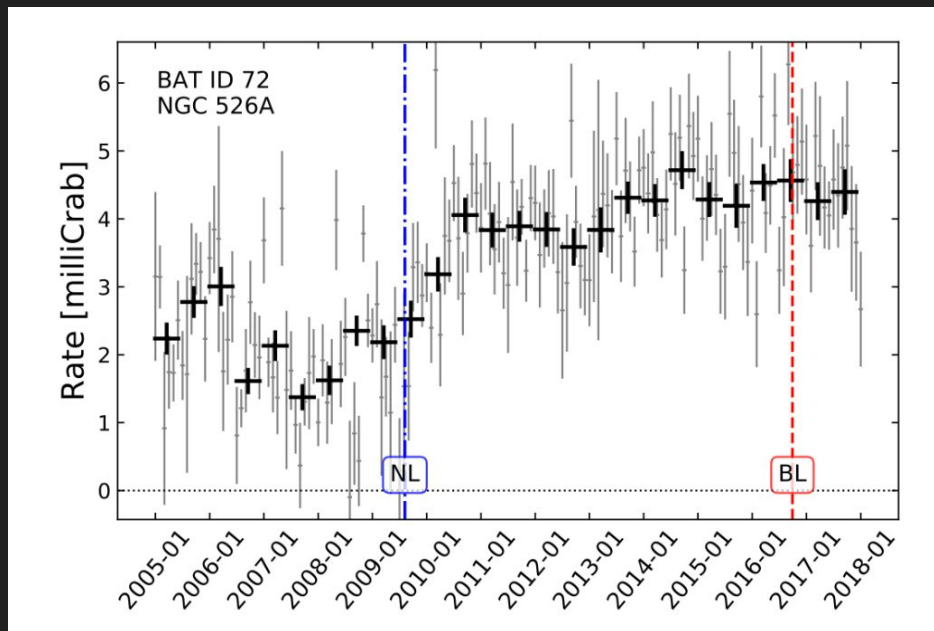
Temple et al. 2023a

Understanding CL AGN in BASS



Swift-BAT provides 14-195 keV light-curves for 9 CL events

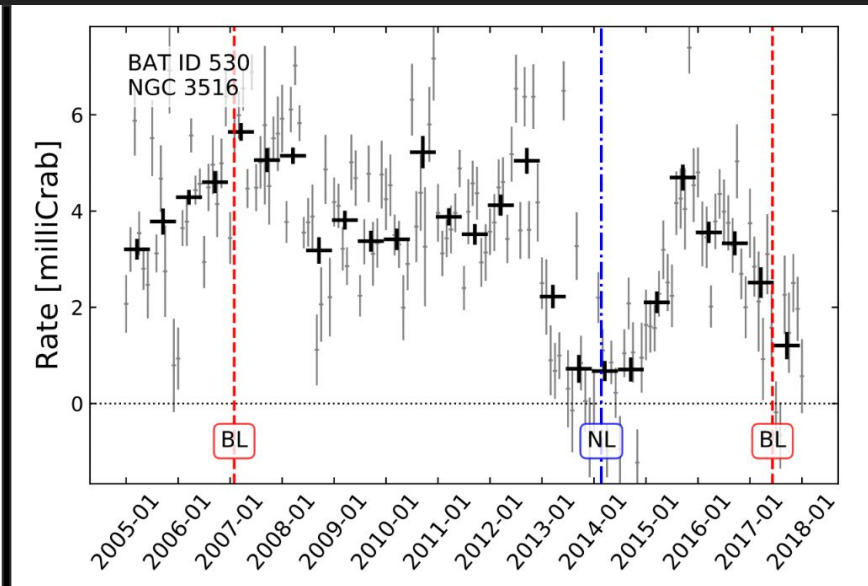
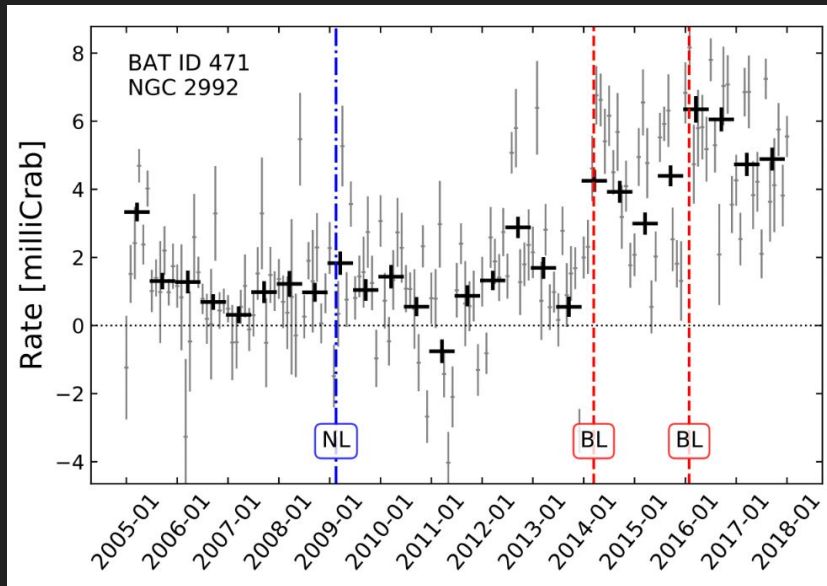
- Six of these CL events show significant changes in their X-ray emission



Temple et al. 2023a

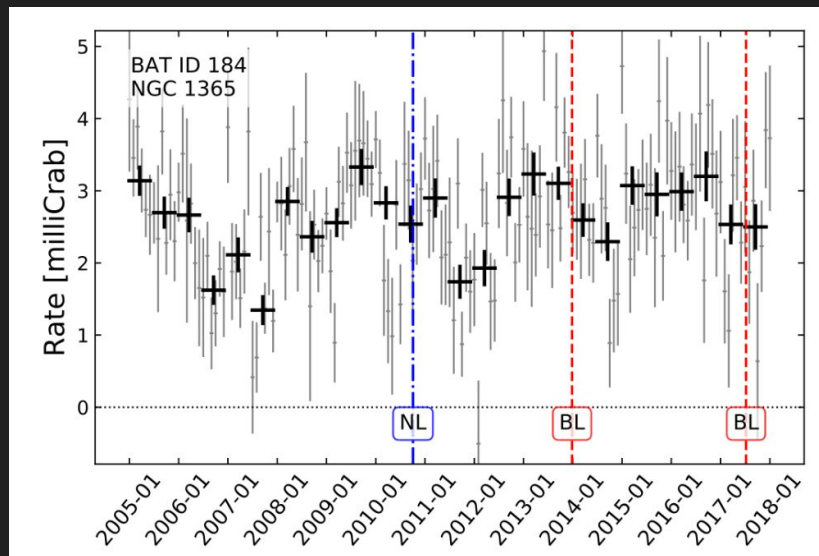
Swift-BAT provides 14-195 keV light-curves for 9 CL events

- Six of these CL events show significant changes in their X-ray emission
- (One AGN, NGC 3516, has two CL events)



Swift-BAT provides 14-195 keV light-curves for 9 CL events

- Three of these CL events could be due to variable obscuration
- For example, NGC 1365 is a well-known CO AGN:



Temple et al. 2023a

Summary

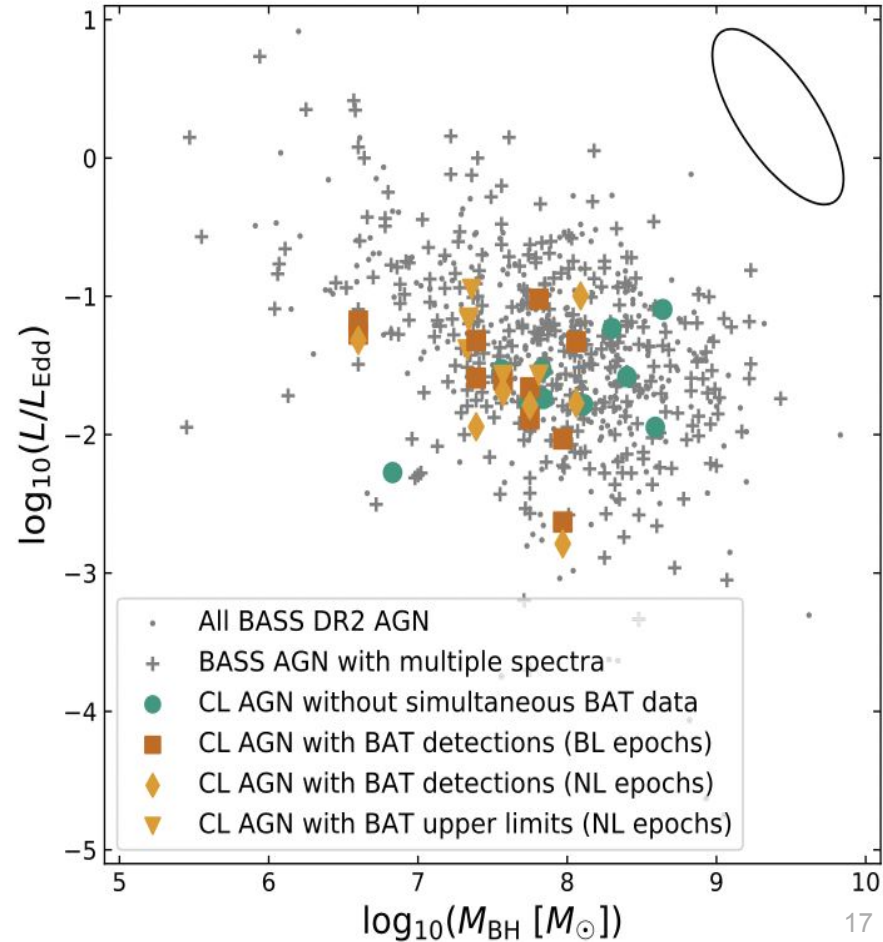
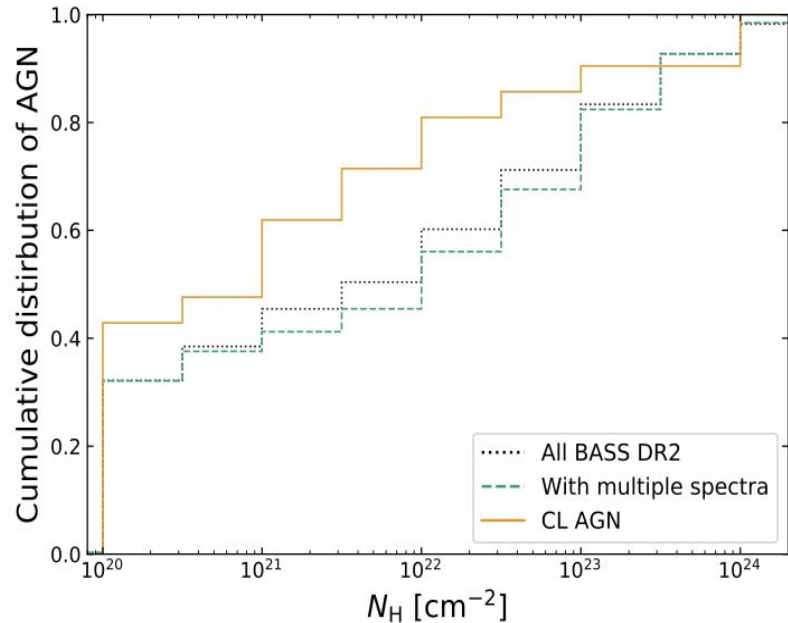
The BASS sample is an ideal test-bed to better understand the physics of highly variable and changing-look AGN, with rich multiwavelength data available

- see also poster by Priscilla Behar

~3% of Swift-BAT AGN show complete BL-NL transitions on ~15yr timescales

Over half of these changing-look events show simultaneous changes in their X-ray emission, which are unlikely to be due to changes in obscuration: changes in the accretion flow must be driving many CS events

CL AGN are typically less obscured (X-ray N_H) and show lower accretion rates ($L/L_{\text{Edd}} \sim < 0.1$) compared to the wider BASS sample



BASS XXXIX: *Swift*-BAT AGN with changing-look optical spectra

Matthew J. Temple¹,¹★ Claudio Ricci^{1,2}, Michael J. Koss^{3,4}, Benny Trakhtenbrot⁵, Franz E. Bauer,^{4,6,7} Richard Mushotzky,⁸ Alejandra F. Rojas^{1,9}, Turgay Caglar,¹⁰ Fiona Harrison,¹¹ Kyuseok Oh,^{12,13} Estefania Padilla Gonzalez,^{14,15} Meredith C. Powell,¹⁶ Federica Ricci,^{17,18} Rogério Riffel¹⁹, Daniel Stern²⁰ and C. Megan Urry²¹

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ABSTRACT

Changing-look (CL) AGN are unique probes of accretion onto supermassive black holes (SMBHs), especially when simultaneous observations in complementary wavebands allow investigations into the properties of their accretion flows. We present the results of a search for CL behaviour in 412 *Swift*-BAT detected AGN with multiple epochs of optical spectroscopy from the BAT AGN Spectroscopic Survey (BASS). 125 of these AGN also have 14–195 keV ultra-hard X-ray light curves from *Swift*-BAT which are contemporaneous with the epochs of optical spectroscopy. Eight CL events are presented for the first time, where the appearance or disappearance of broad Balmer line emission leads to a change in the observed Seyfert type classification. Combining with known events from the literature, 21 AGN from BASS are now known to display CL behaviour. Nine CL events have 14–195 keV data available, and five of these CL events can be associated with significant changes in their 14–195 keV flux from BAT. The ultra-hard X-ray flux is less affected by obscuration and so these changes in the 14–195 keV band suggest that the majority of our CL events are not due to changes in line-of-sight obscuration. We derive a CL rate of 0.7–6.2 per cent on 10–25 yr time-scales, and show that many transitions happen within at most a few years. Our results motivate further multiwavelength observations with higher cadence to better understand the variability physics of accretion onto SMBHs.