



Multi-Wavelength Reverberation Mapping Survey Synergies Between Future Facilities



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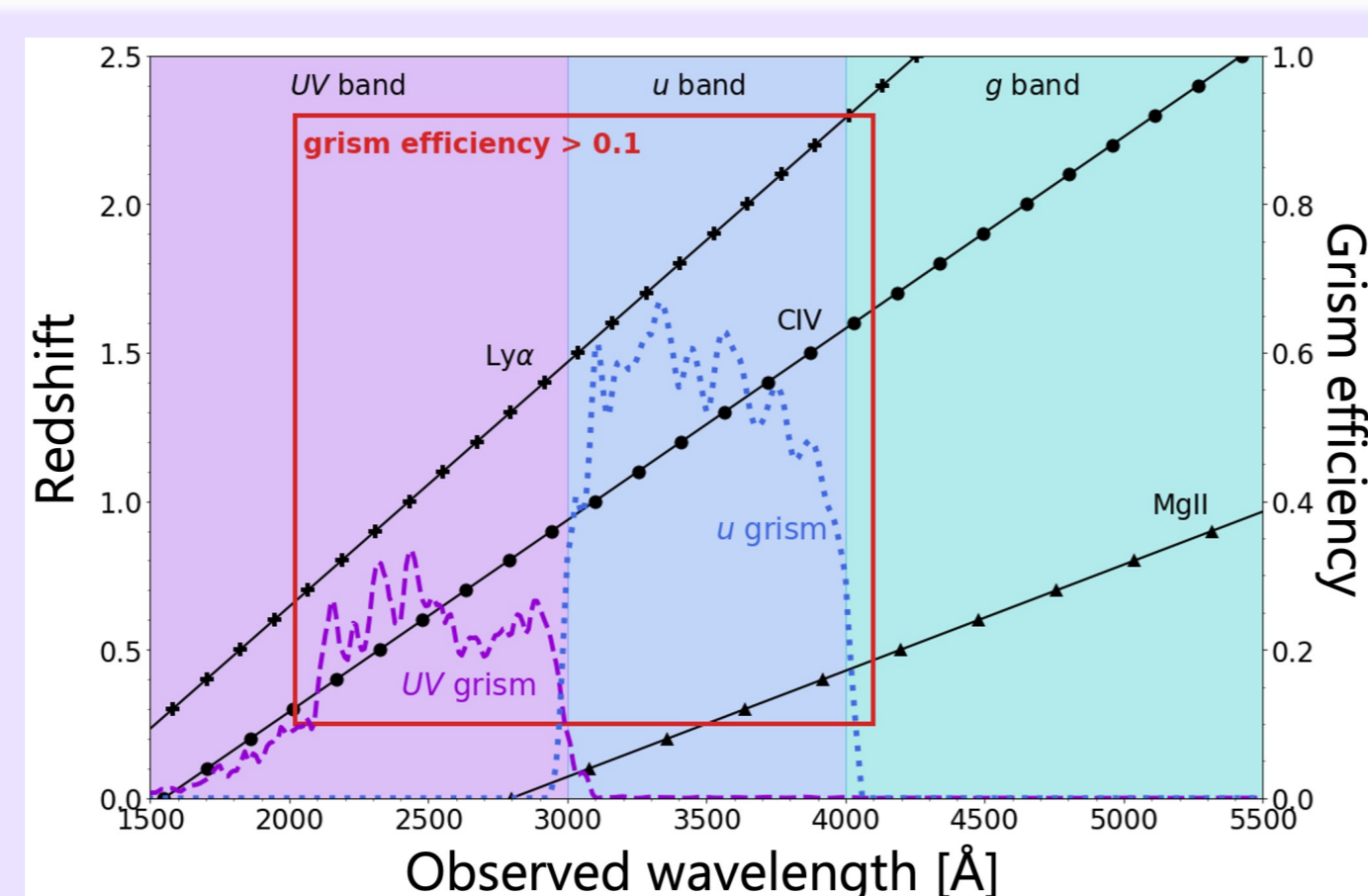
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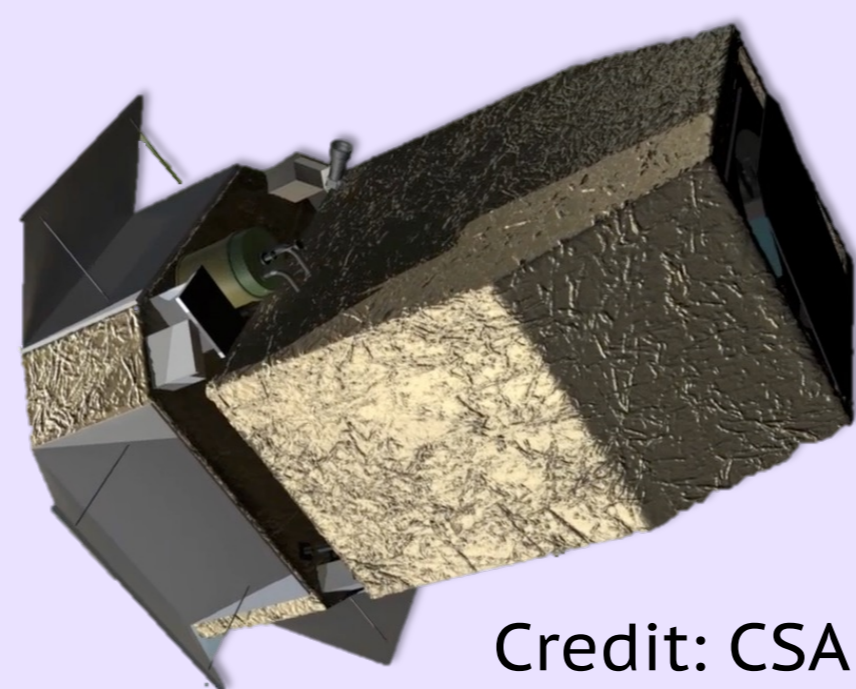
| INTRODUCTION |

In the present era of active galactic nuclei (AGN) science, multi-wavelength surveys are gaining ground. The use of reverberation mapping (RM; Peterson et al. 1983) as a powerful technique to extract AGN sizes from their varying fluxes has flourished over the past decades. Next-generation facilities, e.g., the *Cosmological Advanced Survey Telescope for Optical and ultraviolet (UV) Research (CASTOR; P.I.: P. Côté)* – a proposed mission led by Canada, will perform large-scale RM surveys. We are contributing to the planning of such surveys by developing an AGN Survey Simulator tool (Khatu 2022, PhD thesis). Although a stand-alone RM program with *CASTOR* is feasible, synergies with other facilities coming up in the future will allow maximizing the science outcomes from AGN RM surveys.

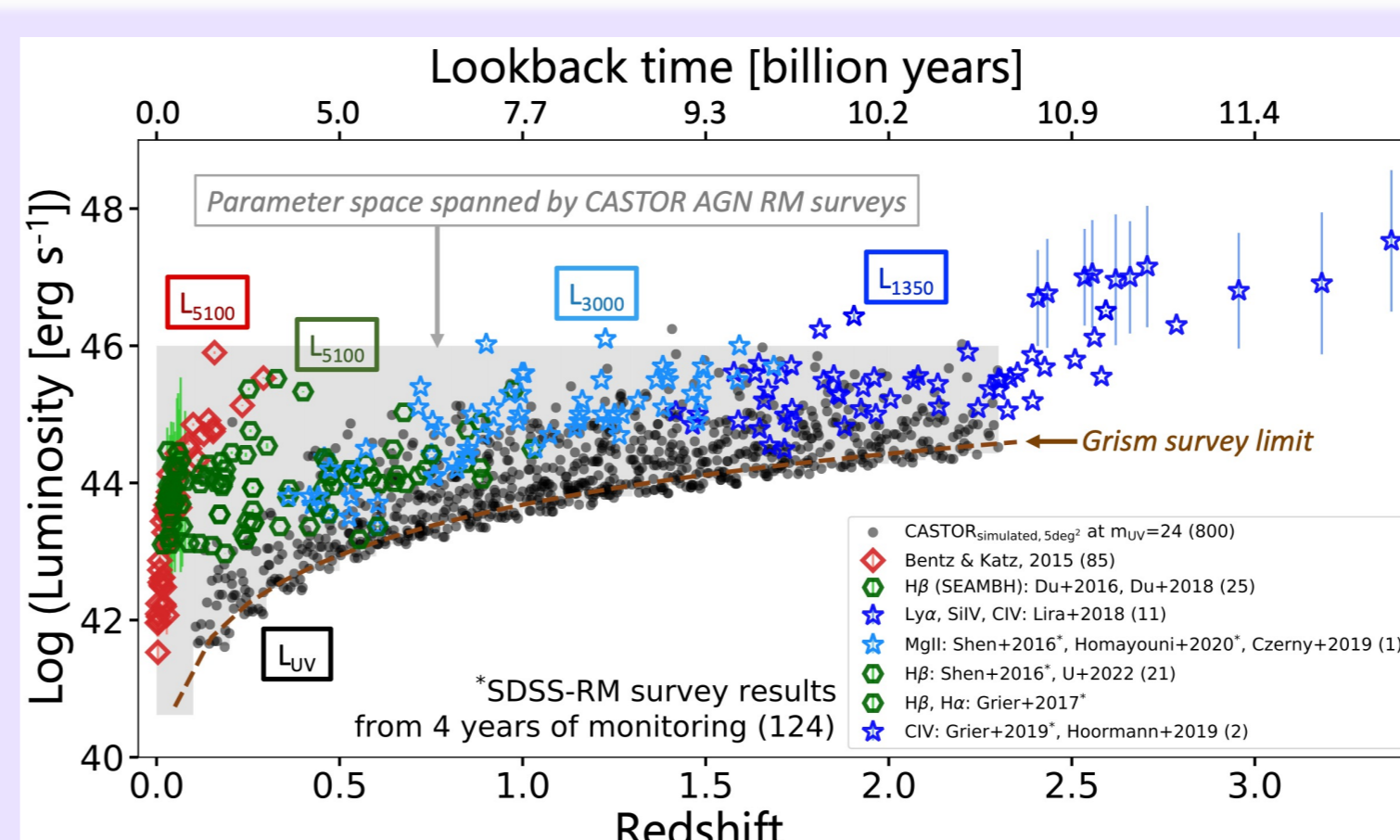
Photometry (UV, u, and Rubin/g)
+
Grism (UV and u)
+
UV Multi-Object Spectroscopy (UV and u)



CASTOR

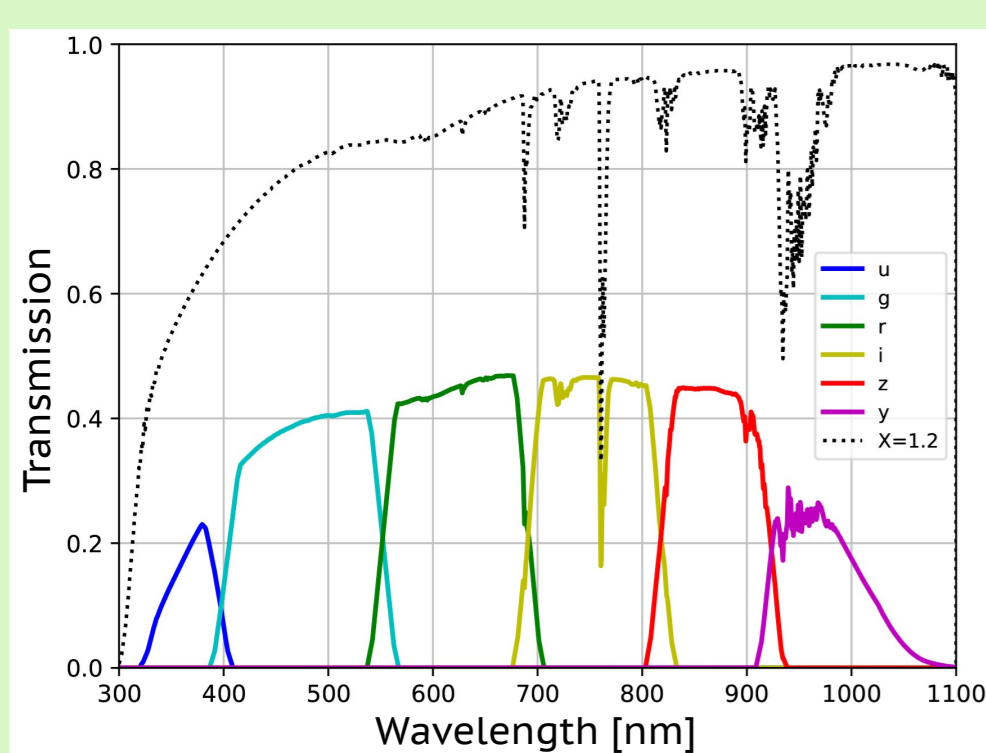


Credit: CSA



AGN Legacy Survey aims to measure black hole masses of a significantly large number of AGN in a unique phase space

Vera C. Rubin Observatory (Rubin)



Credit: Ivezić et al. (2019)



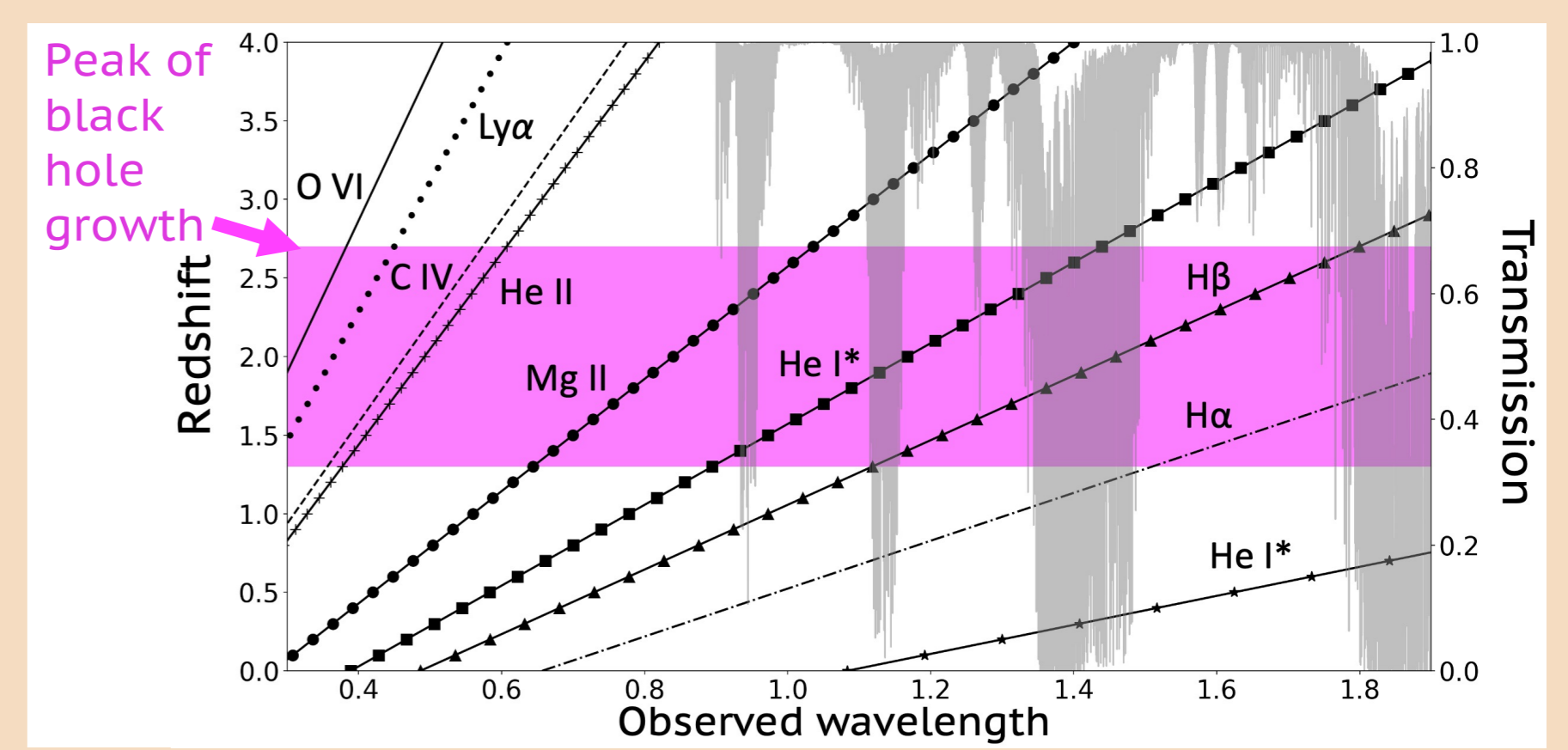
Credit: Rubin/NSF/AURA

- High cadence + long baseline for RM studies
- Science: Accretion-disk RM + construction of spectral energy distributions for AGN in Rubin Deep Drilling Fields

Maunakea Spectroscopic Explorer (MSE)



Credit: MSE



- Probe high redshift, high luminosity AGN: H β , MgII
- Science: Accretion-disk and BLR RM + follow-up high-resolution spectroscopy of *CASTOR* targets

Fig. 1. AGN Survey Synergies

| AGN SURVEY SIMULATOR (USE CASE: CASTOR) |

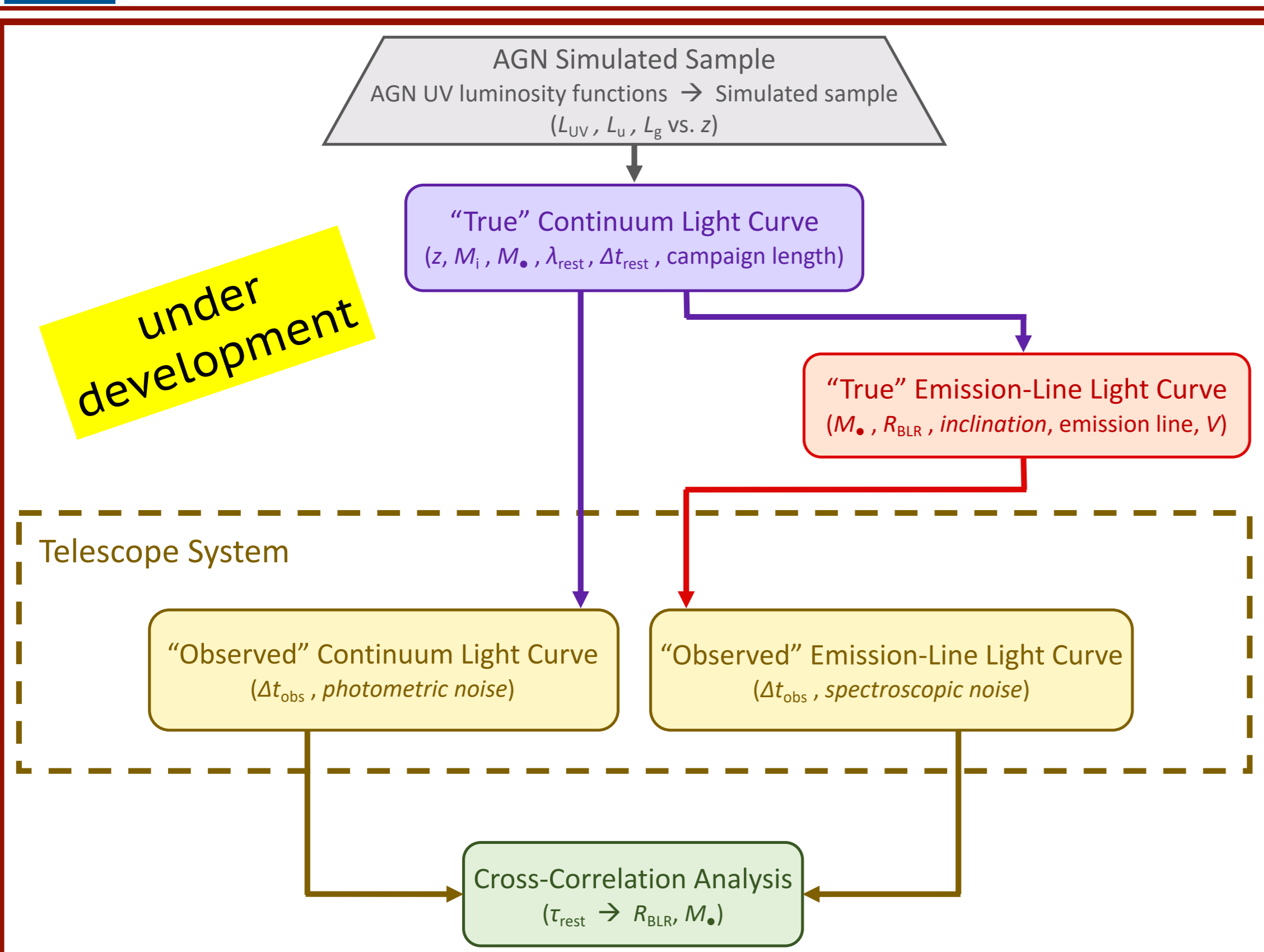


Fig. 2. AGN Survey Simulator design with modular framework. (Language: Python)

Fig. 3. Illustration 1 (Khatu 2022, PhD thesis) – Higher redshift, higher luminosity (larger symbols) AGN need longer campaigns. In a nominal, 1-year period, *CASTOR* will probe AGN with $L_{AGN} < 10^{46.0}$ erg s⁻¹.

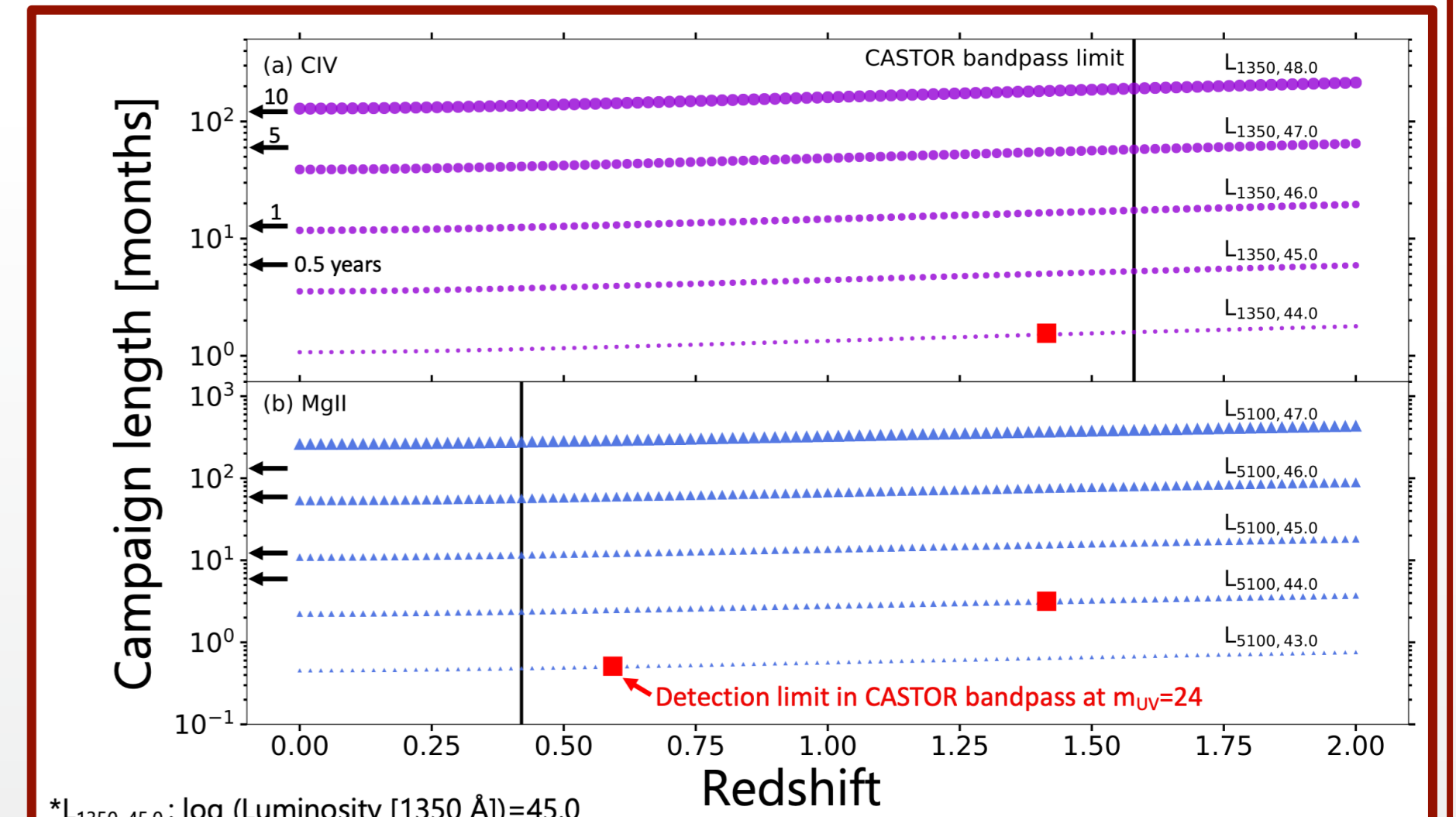
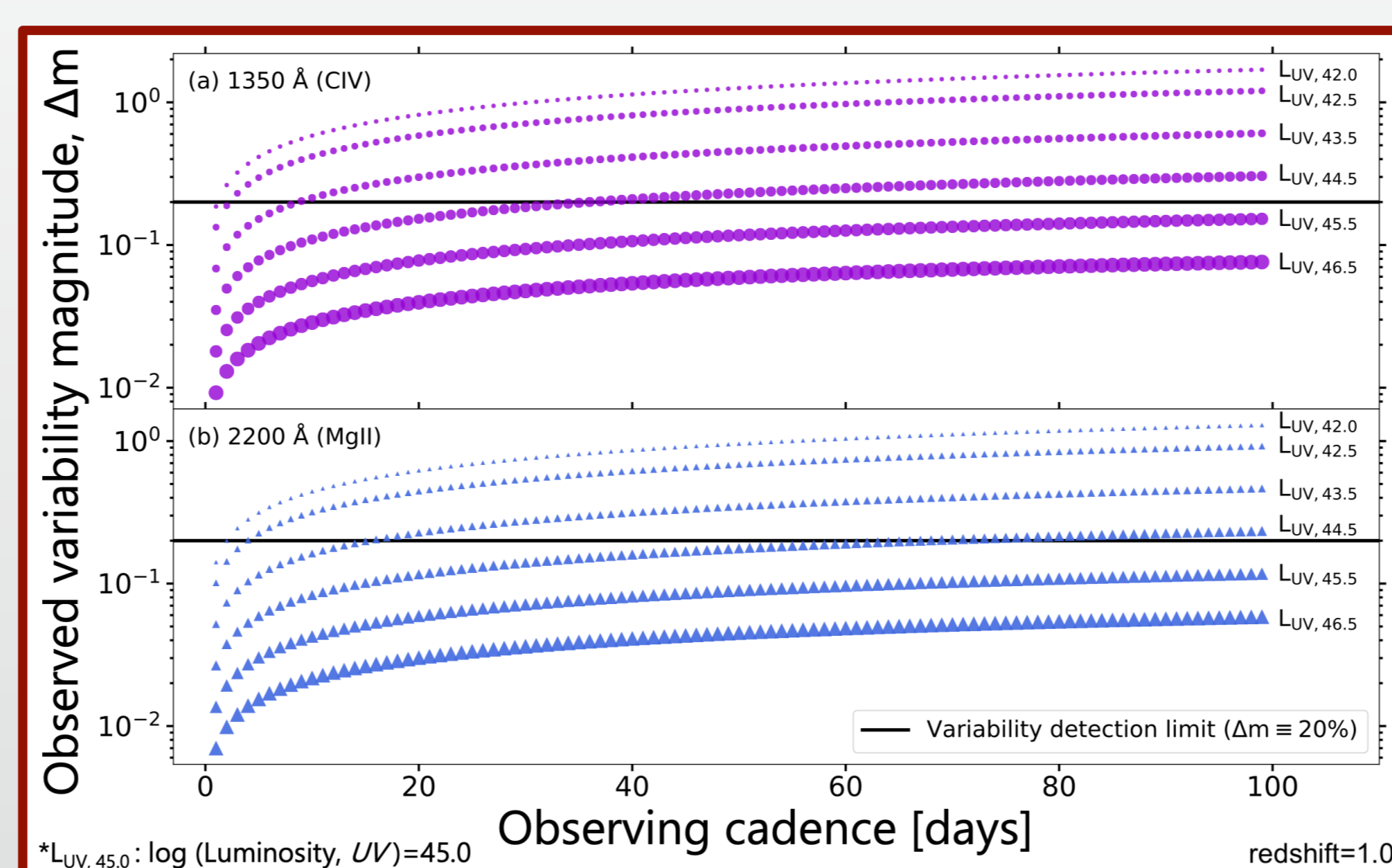


Fig. 4. Illustration 2 (Khatu 2022, PhD thesis) -- For more luminous AGN, we need longer cadences to detect sufficient variability ($\Delta m \geq 20\%$) in these objects.

| FUTURE WORK |

- Testing forward modelling techniques for grism spectral analysis in the low-count regime.
- Adapting the simulator for MSE survey planning

| ACKNOWLEDGEMENTS |

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