

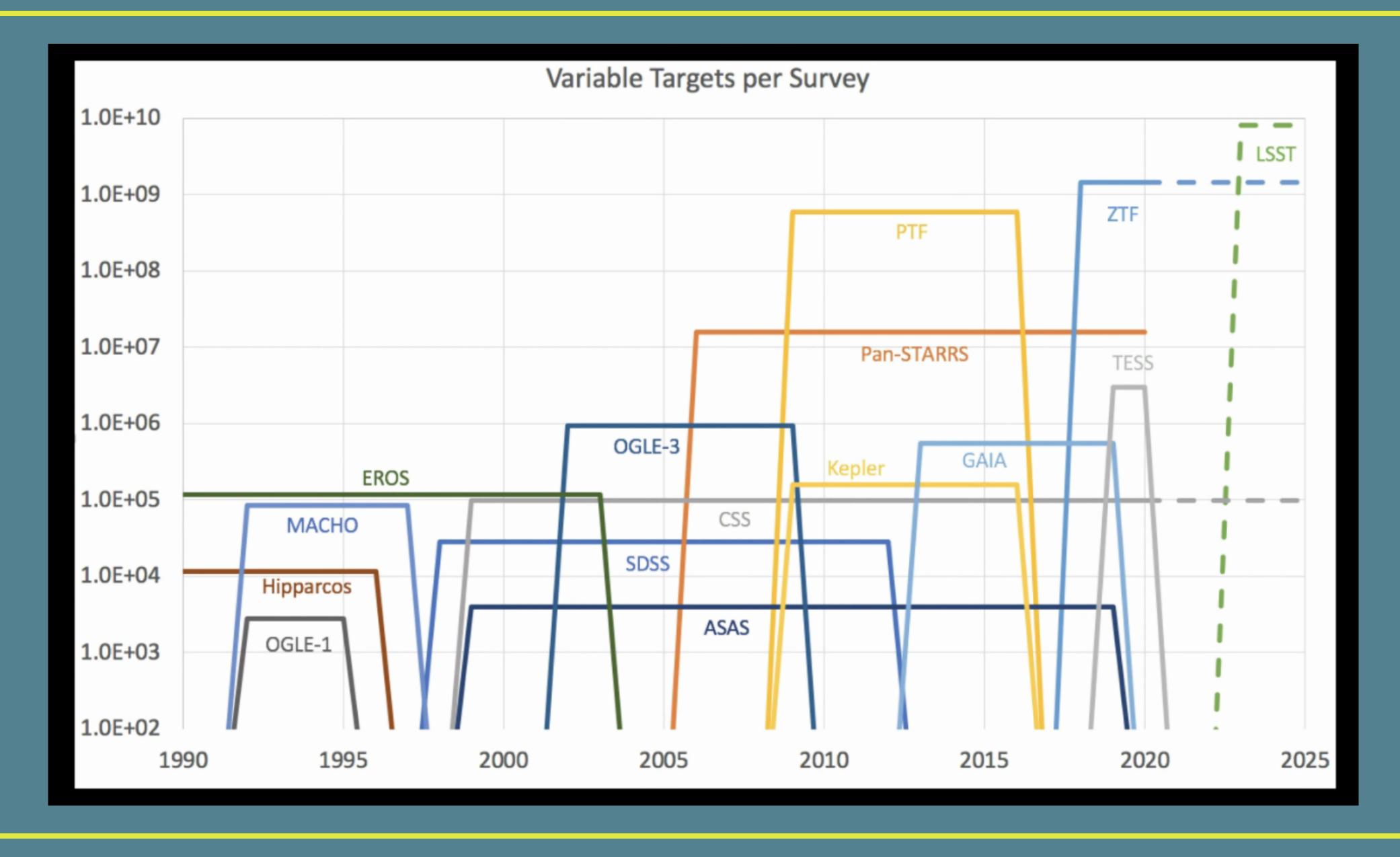
The Restless Nature of AGN: 10 years later

SPECTROSCOPIC SURVEYS FOR TIME-DOMAIN SCIENCE

Paulina Lira

28 June 2023

VARIABILITY SURVEYS



VARIABILITY SURVEYS



VARIABILITY SURVEYS

Dedicated facilities:

Zwicky Transient Factory: 1.2m, gri ~ 20.5 PS1: 1.8m

ATLAS: 50cm Schmidt telescope Kiso Schmidt: 1m

ASAS-SN: 24 x 14cm telescopes SkyMapper: 1.35m

QUEST-La Silla: 1m Schmidt -> LS4 BlackGem: 65cm telescopes

Vera Rubin LSST: 6.5m, 10 years, ugrizy ~ 25-27 over 18,000 deg²

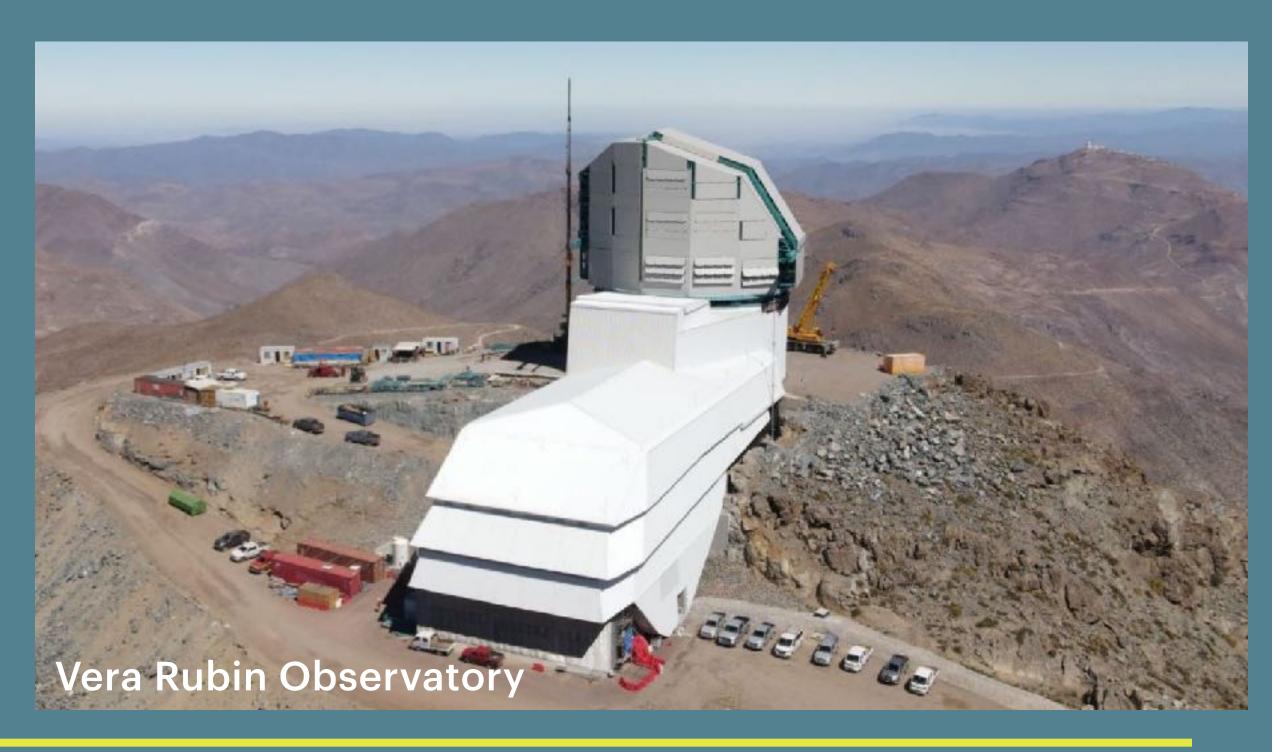
ugrizy ~ 27-29 over 38 deg² (deep drilling fields)

Time-resolved Large Imaging Surveys:

SDSS/S82: 2.5m APO, 3 deg²

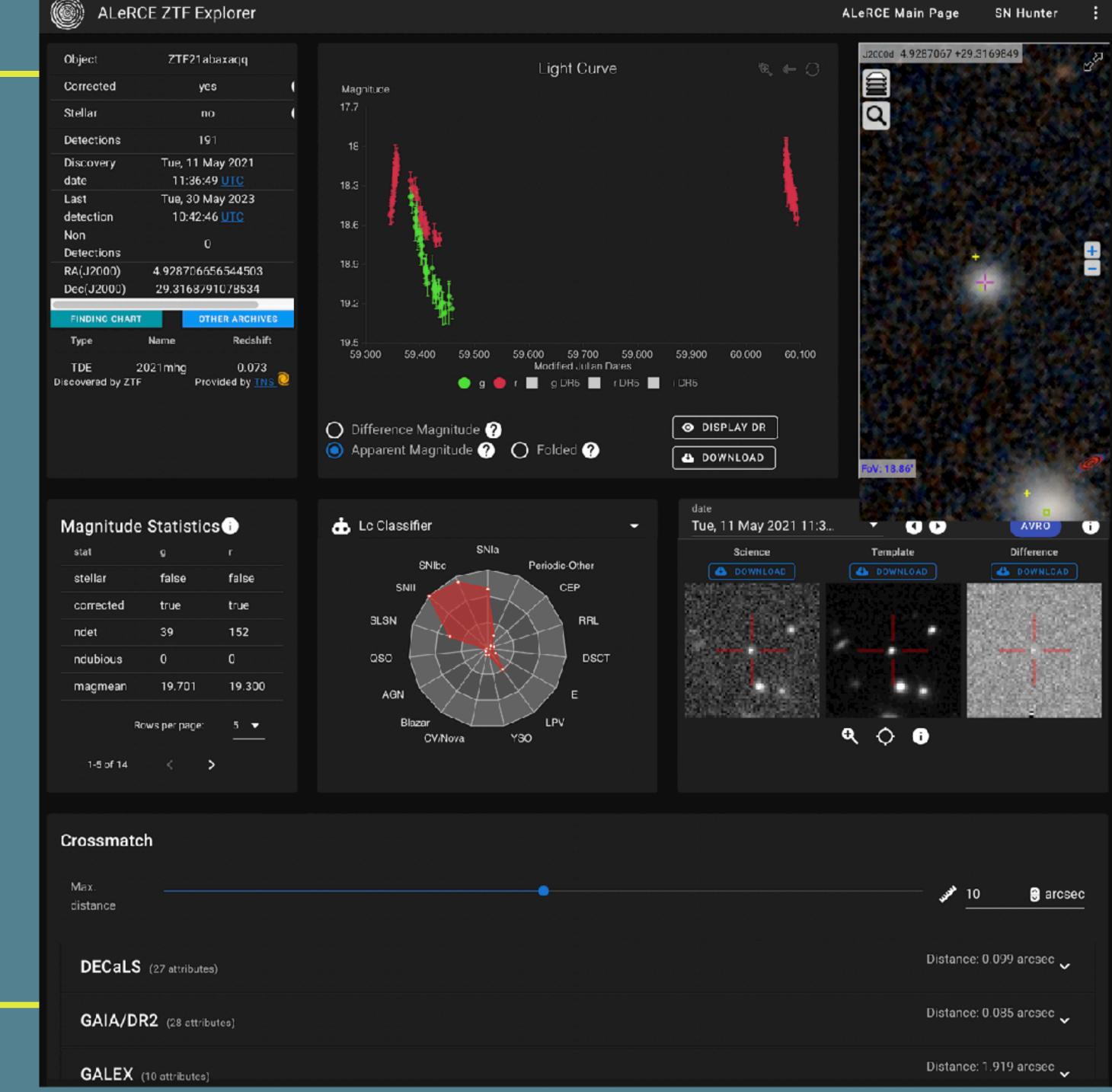
DES: 4m Blanco, 2.5 deg²

HSC-SSP: 8m Subaru, 1.7 deg²



Alert stream brokers will help with the classification bottleneck of transient events

See <u>alerce.online</u>





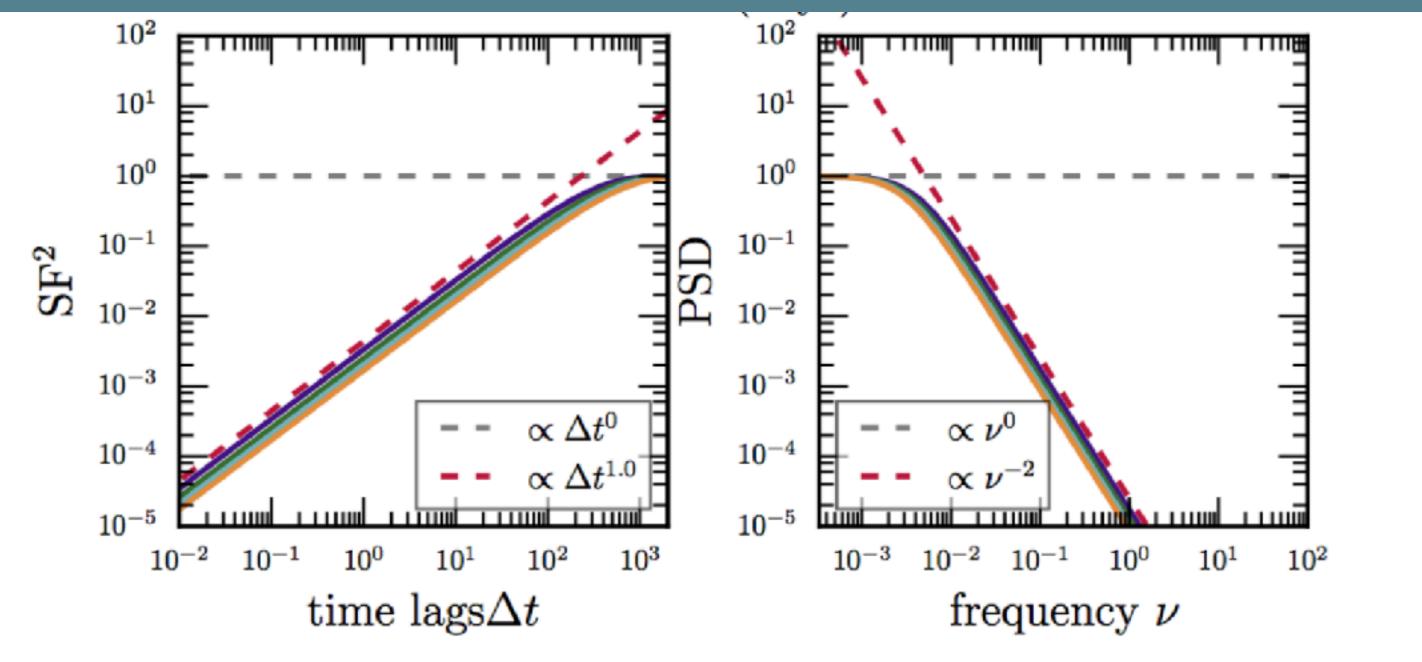
AGN rarely trigger transient-like events.....

The bread and butter of AGN variability studies: Data Releases

With the exception of CL AGN and TDEs



Power Spectrum



Moreno et al. 2019

MULTI-WAVELENGTH SYNERGY

Other massive DEEP-WIDE field imaging surveys coming online in the next years

Vera Rubin LSST (NOIRLAb): 6.5m, 10 years, $ugrizy \sim 25-27$ over 18,000 deg² $ugrizy \sim 27-29$ over 38 deg² (deep drilling fields)

Euclid Space Telescope (ESA): 1.2m, 6 year mission, riz+YJH ~ 24 over 15,000 deg², also spectra! ~ 26 over 40 deg²

Nancy Roman Space Telescope (NASA / ex WFIRST): 2.4m, YJH ~ 26-27 over 2,000 deg²







IMAGING - SPECTROSCOPY PARTNERSHIP

Natural symbiotic relation between time-domain astronomy and massively multiplexed spectrographs

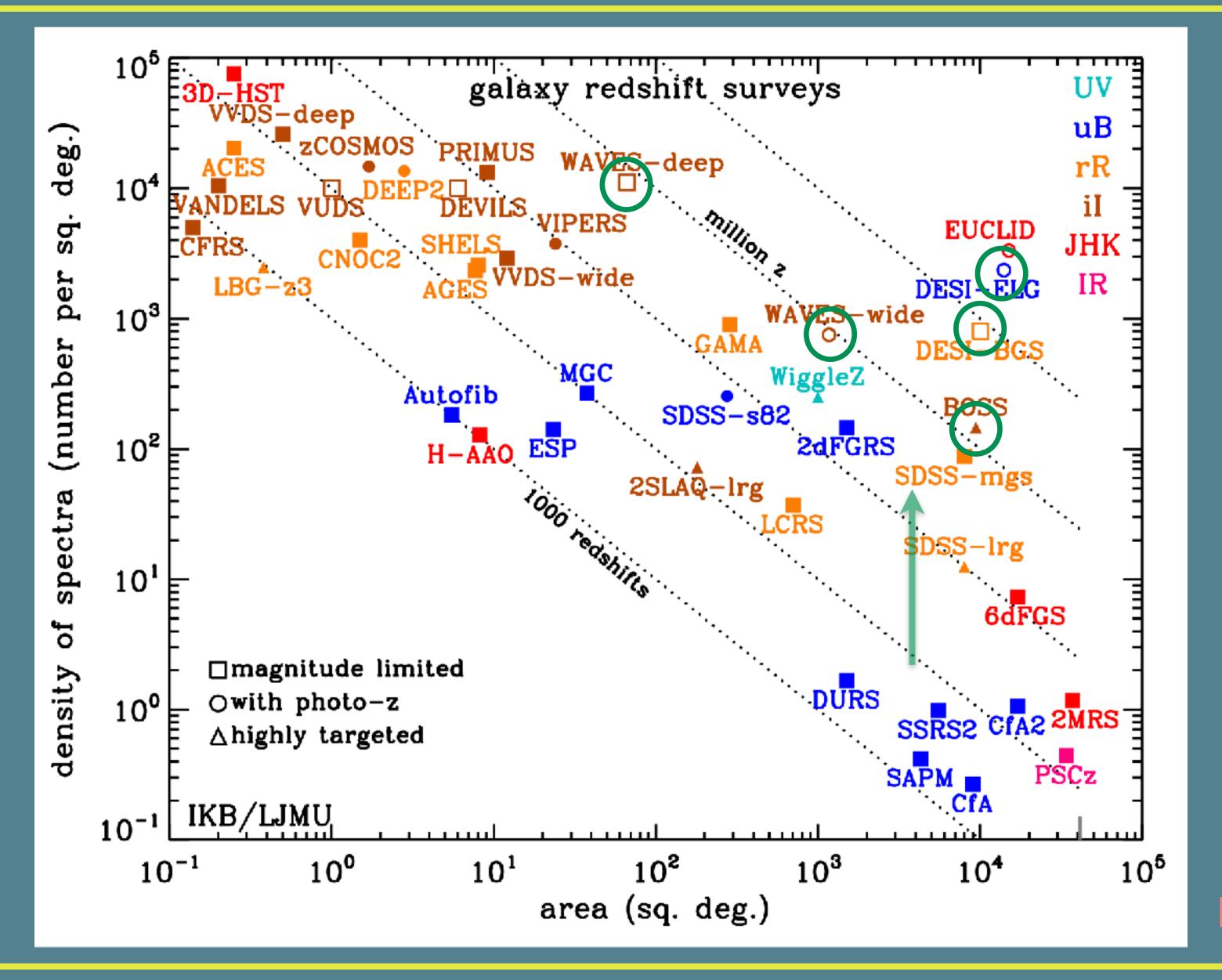
Desired Measurables:

- Redshifts
- Emission line fluxes / continuum
- Detection of broad components
- Host Galaxy characterization
- Variations in continuum / line emission profiles from repeated observations

GROUND BASED SPECTROSCOPY

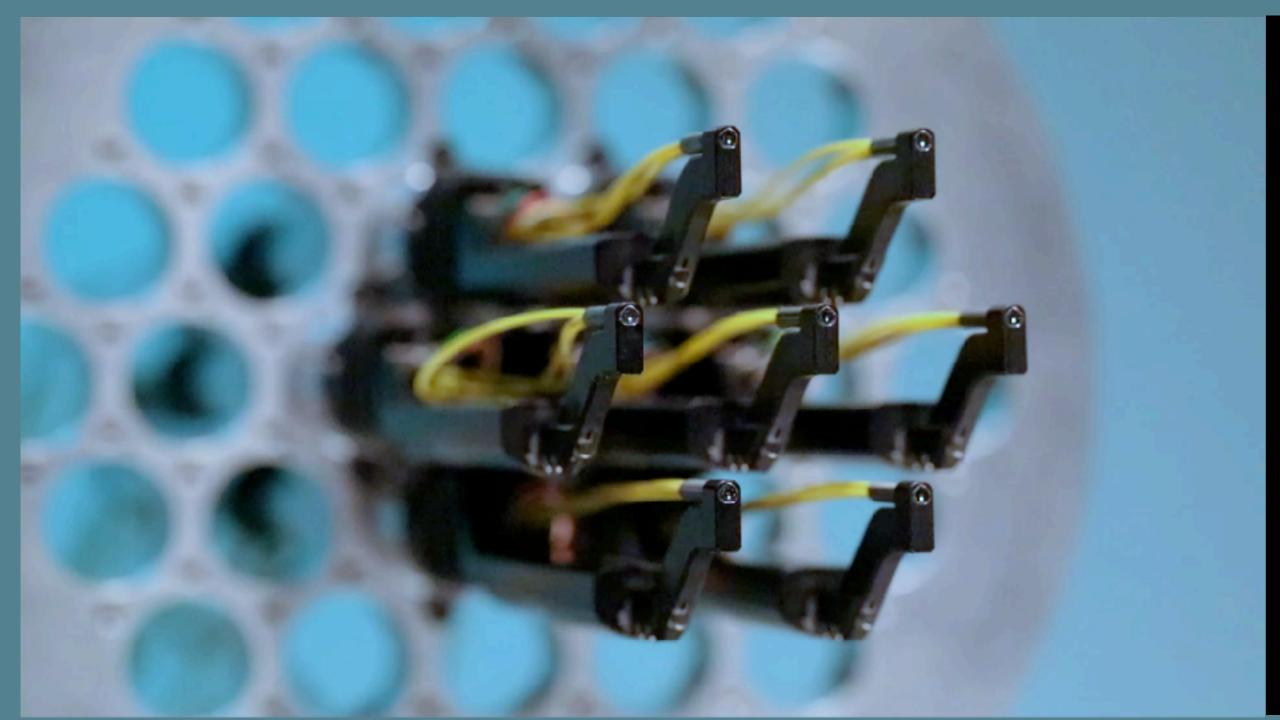
	Instrument/Telescope	Collecting Area m ²	Field of view deg ²	Multiplex	R opt/NIR/both
SDSS-V BOSS Spec APOGEE Spec	APO / du Pont ✓	4.9	3.0 / 2.1	500	> 1500
				300	22000
4m class funded	4MOST	10.7	4.00	1400	6500 20000
	Mayall 4m / DESI ✓	11.4	7.08	5000	> 2000
	WHT / Weave	13.0	3.14	1000	5000 20000
8-10m class funded	Subaru / PFS	52.8	1.25	2400	> 3000
	VLT / MOONS	52.8	0.14	500	> 4000
					9000 20000
Proposed & unfunded	Megamapper @ GMT	28.0	7.06	> 20,000	> 2000
	Keck / FOBOS	76.0	0.087	1800	3500
	MSE @ CFHT	78.5	1.52	4000	> 1000 > 20000
	ESO Spectel	113.1	4.90	15000	3000
	ESO wST	78.5	5.00	20000	> 2000 > 20000

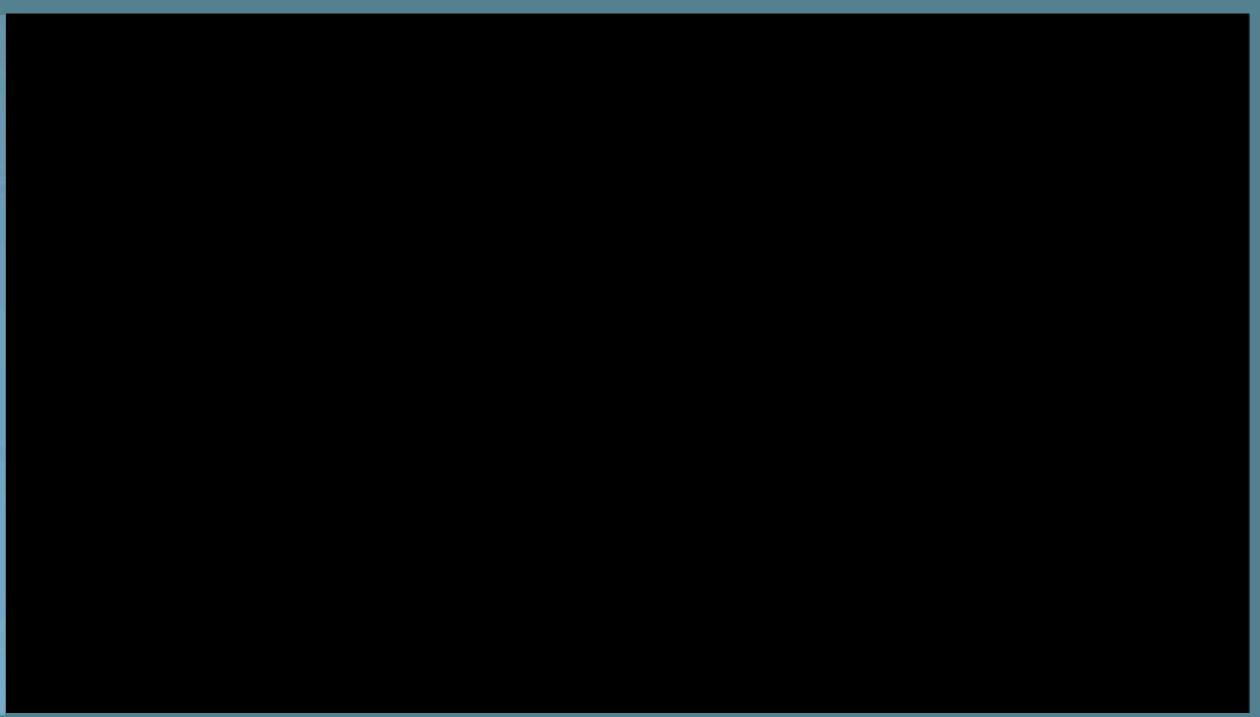
SPECTROSCOPIC SURVEYS



I. K. Baldry

SDSS-V APOGEE-2 / WEAVE



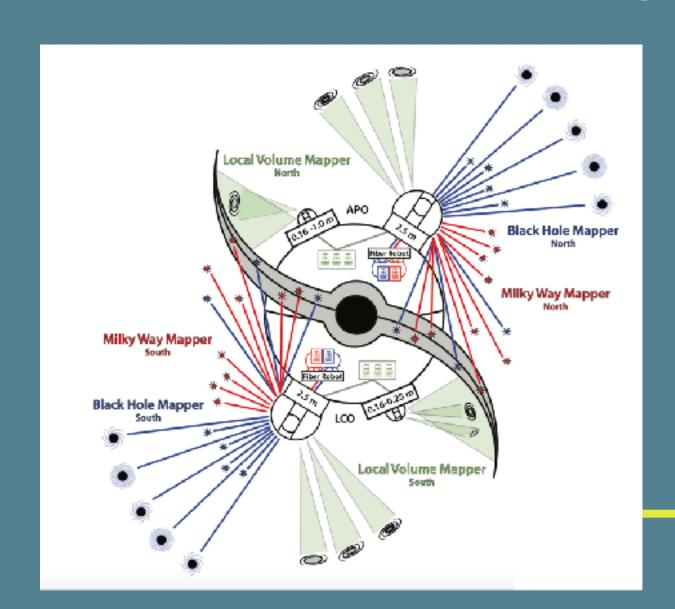


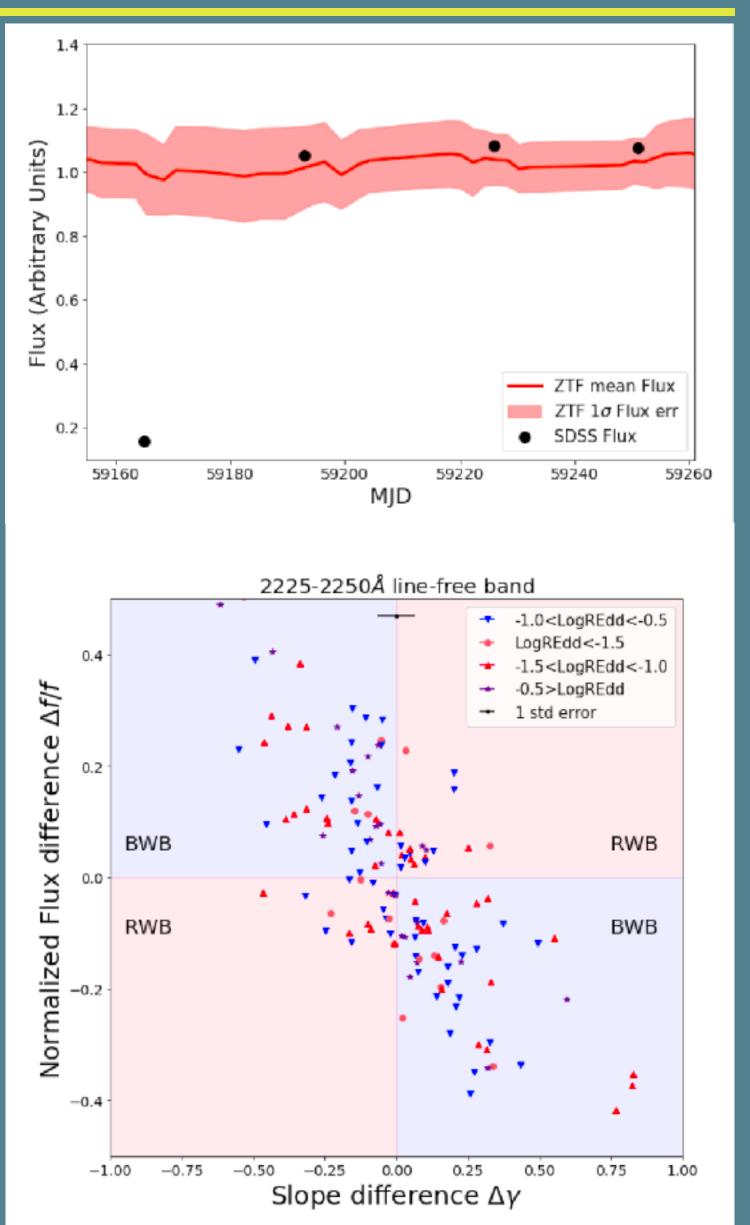
SPECTROSCOPIC SURVEY HIGHLIGHTS: SDSS-V BHM

- AQMES: All-Quasar Multi-Epoch Spectroscopy
- Reverberation Mapping (# SDSS-III RM Project)
- Spiders: SPectroscopic IDentfication of ERosita Sources
- SCS: Chandra Source Catalog

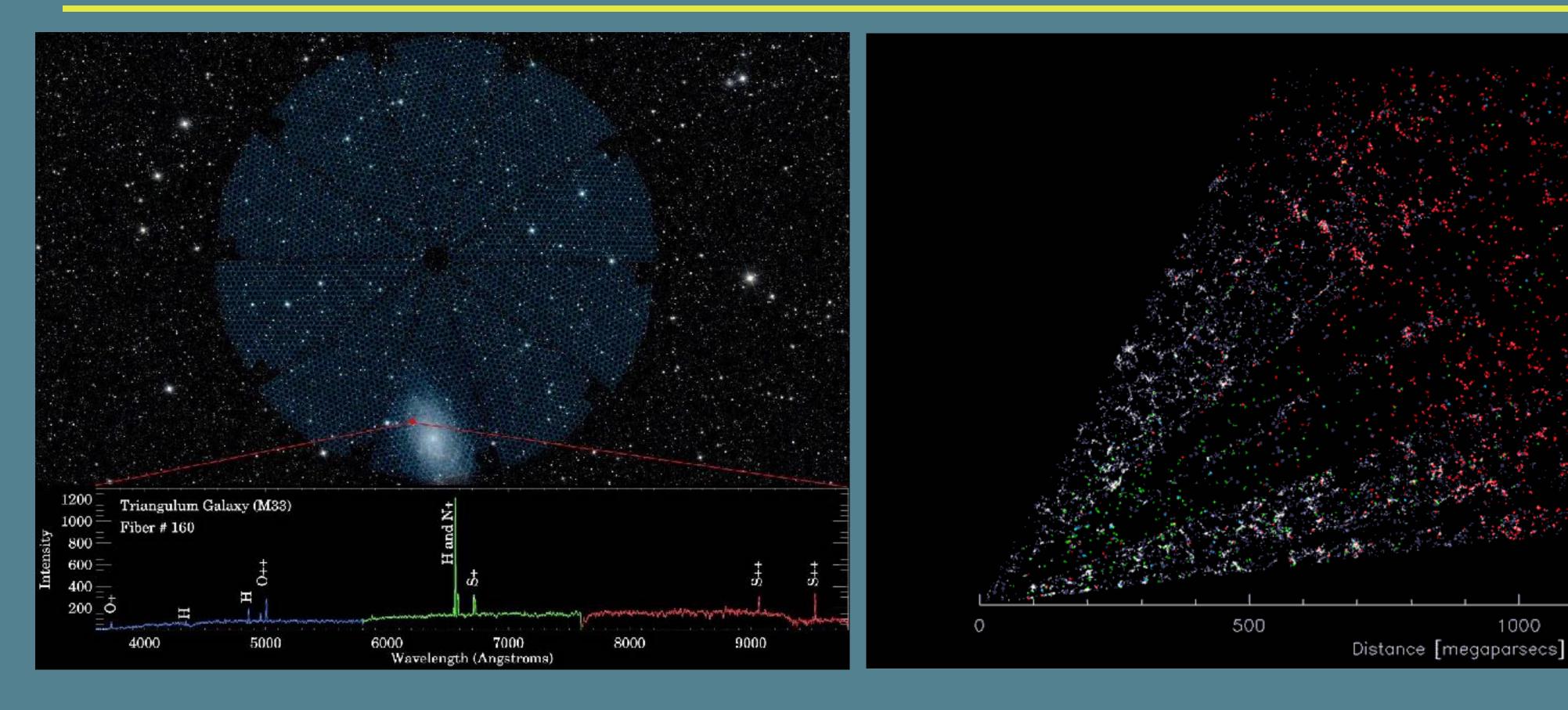
18th DRs contains SDSS-V spectra

First results: Grisha Zeltyn (see talk tomorrow)
Santiago Bernal (see poster)





SPECTROSCOPIC SURVEY HIGHLIGHTS: DESI



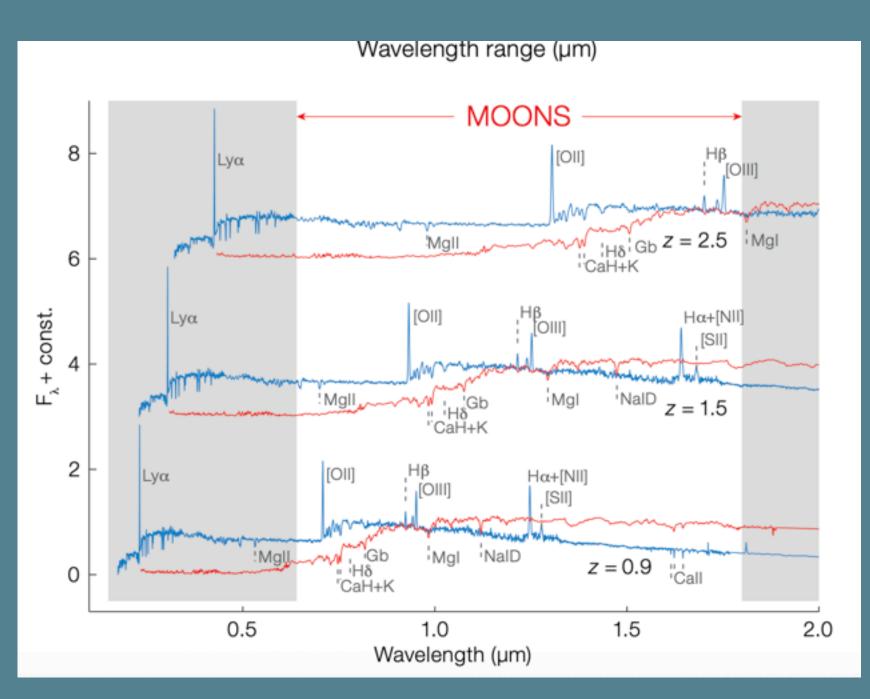
DESI Early Data Release contains 2M spectra from the Survey Validation phase of galaxies, quasars and stars

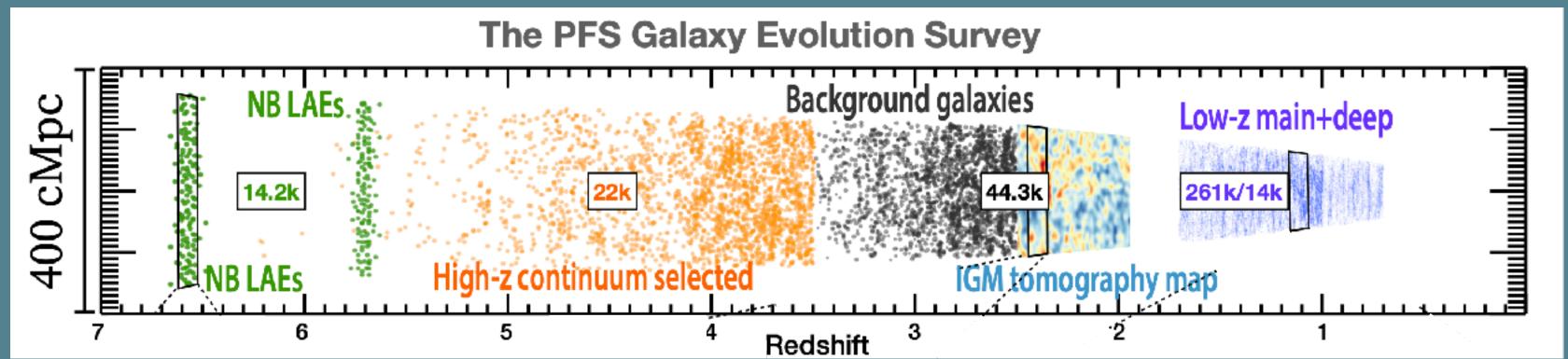
DESI Press Release 2022

DESI targets have been selected from the Legacy Surveys (DECaLS, BASS and MzLS)

SPECTROSCOPIC SURVEY HIGHLIGHTS: MOONS & PFS

8m-class: MOONS (VLT) / PFS (Subaru)

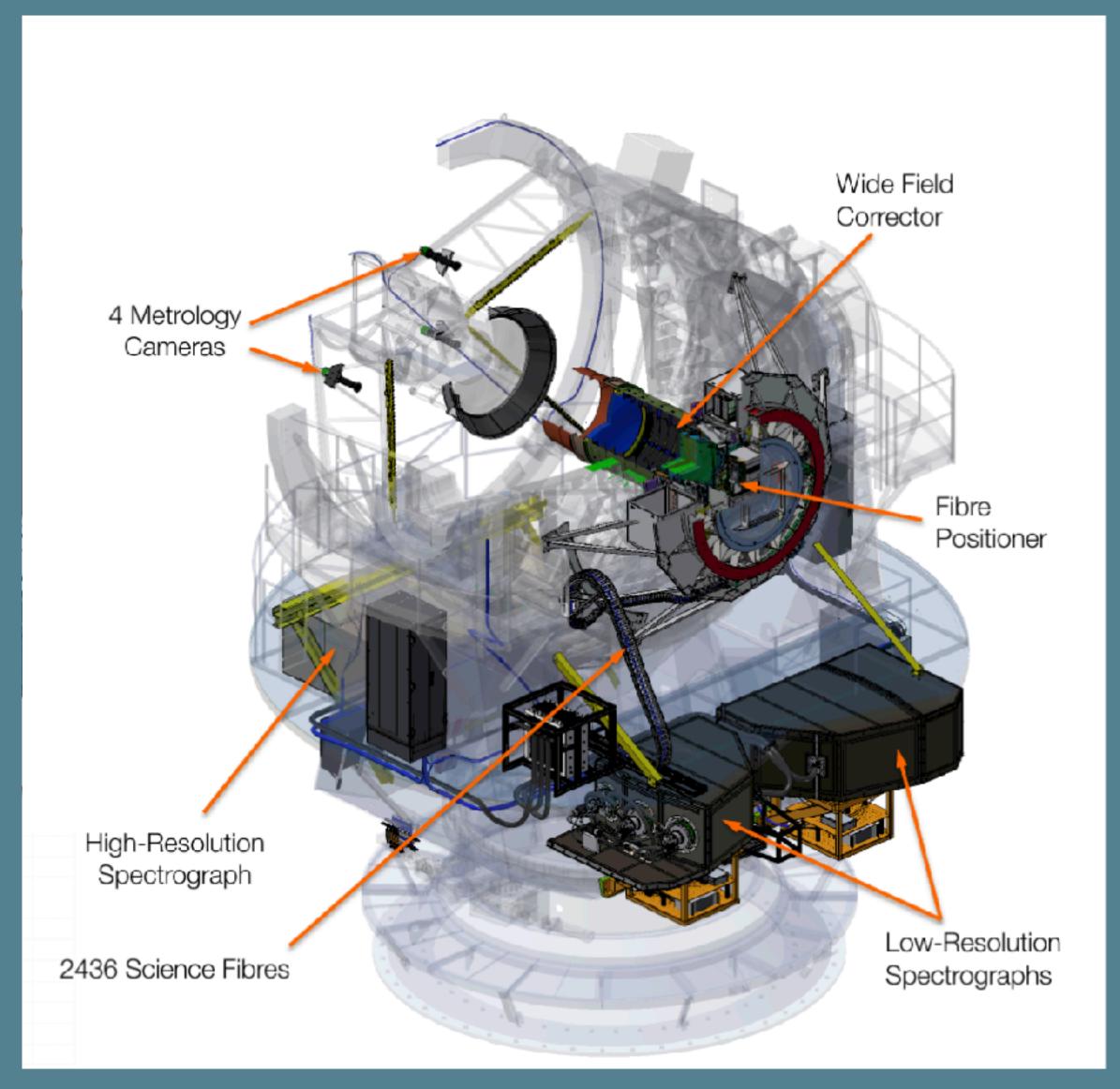




Greene+2022

Maiolino+2020

THE 4M MULTI-OBJECT-SPECTROSCOPIC-TELESCOPE (4MOST)



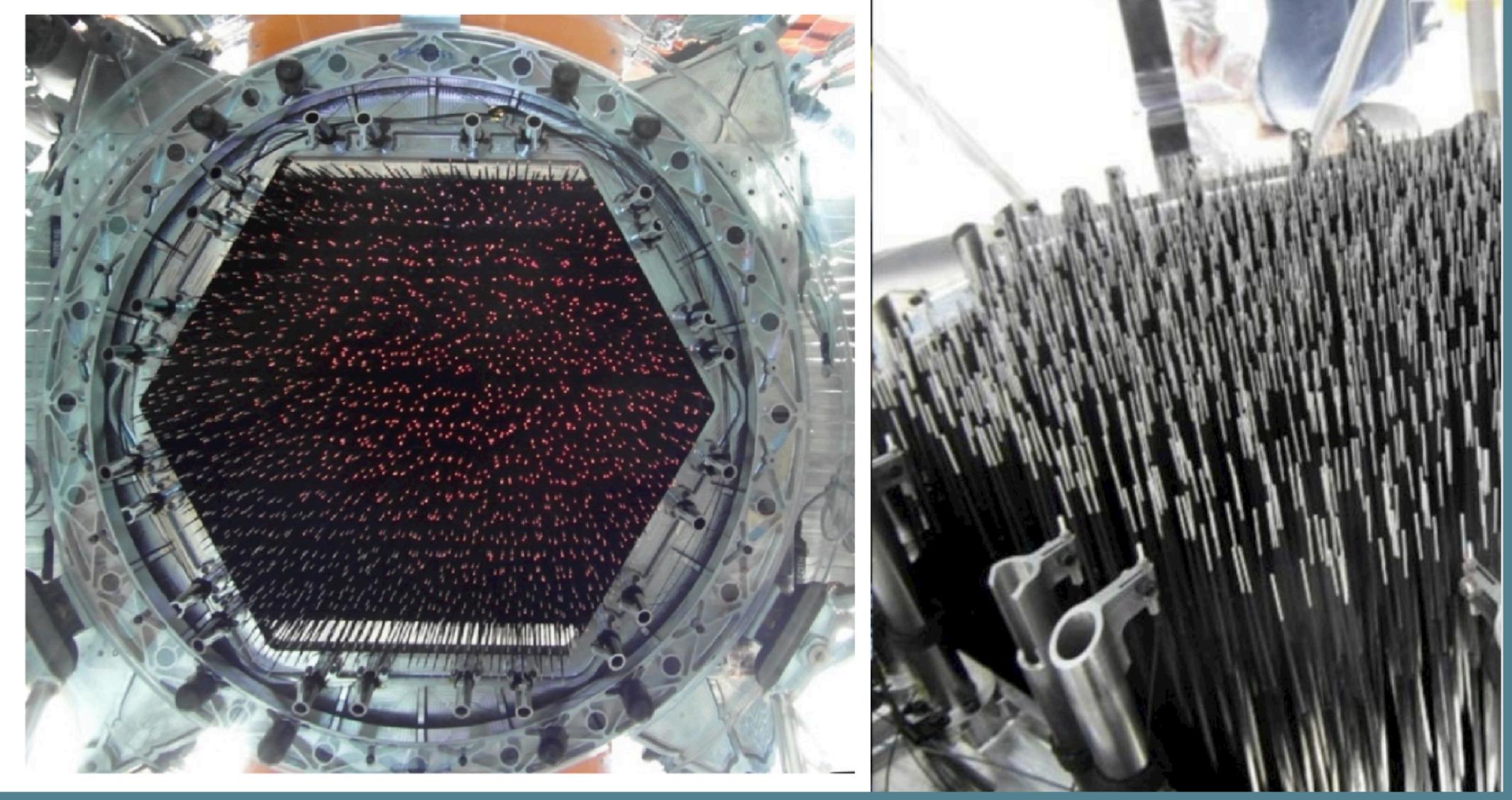
Low Resolution Spectrographs (2x)	Passband: $370\text{-}950\text{nm}$ R > $\lambda \times 10$ for $400\text{nm} < 500nm$ R > 6000 for $500\text{nm} < 885nm$ Velocity Accuracy < 1 km/s 1624 fibres in total
High Resolution Spectrograph	Passbands: 392.6-435.5, 516-573, 610-679 nm <i>R</i> > 18,500 Velocity Accuracy < 1km/s 812 fibres

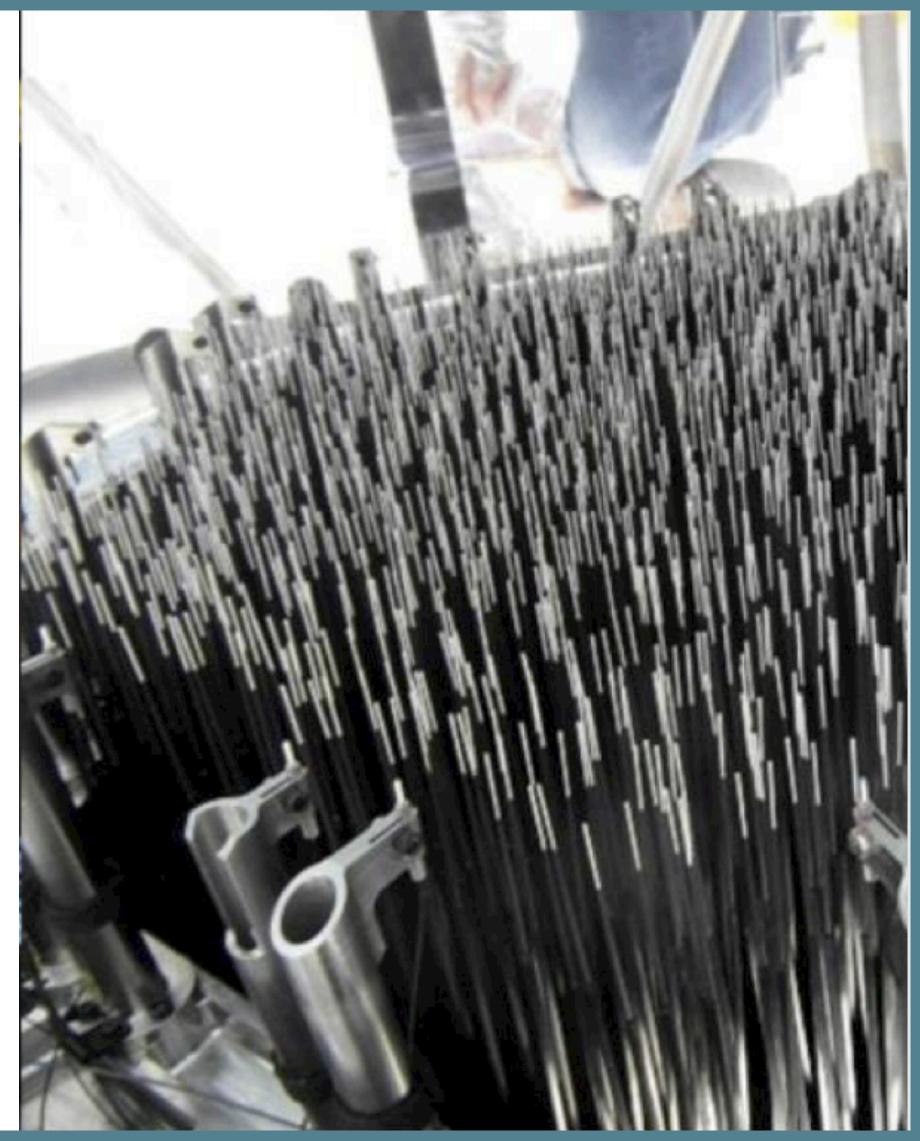
In a 5-year survey 4MOST observations will:

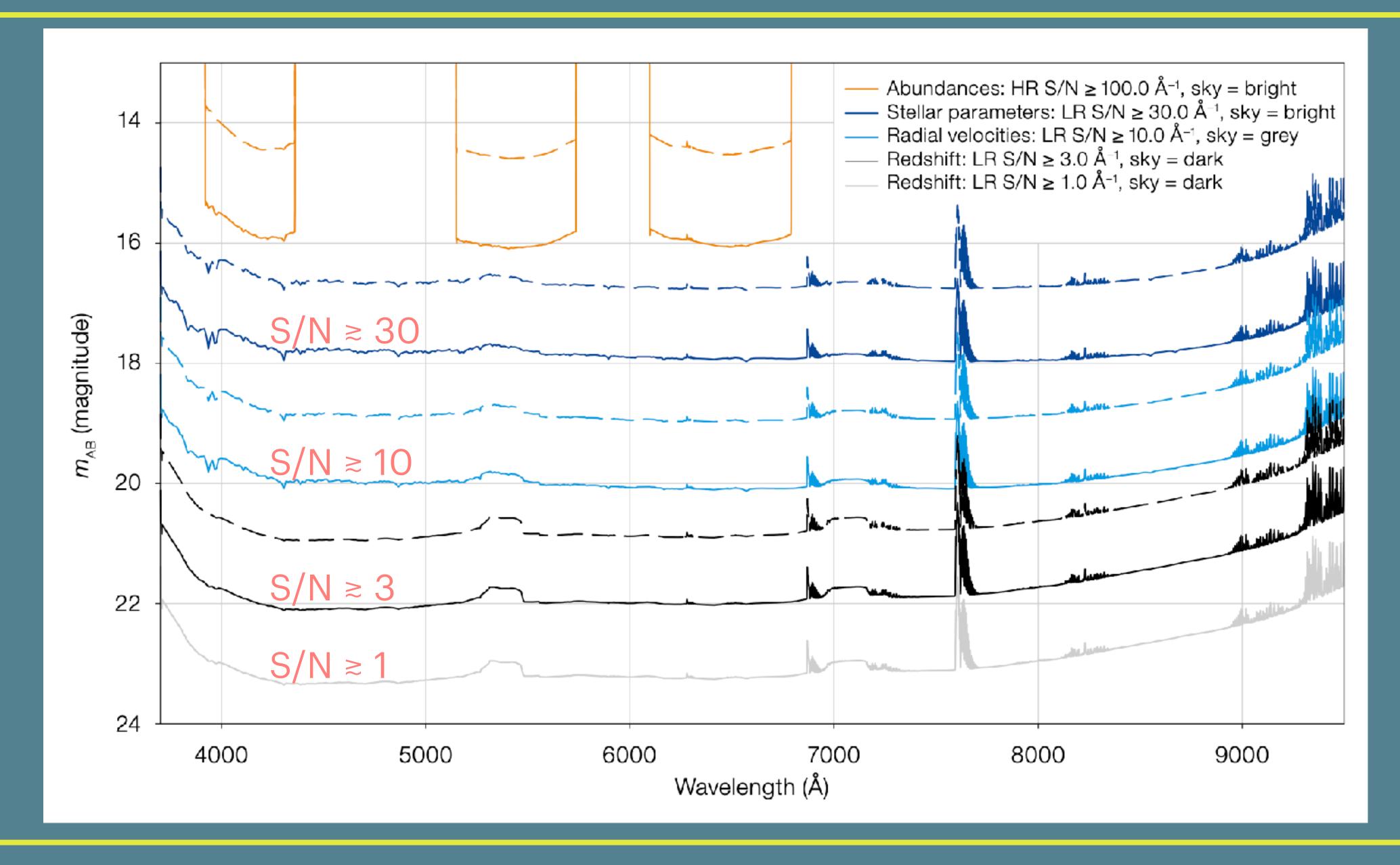
- Cover at least twice an area of =16,000 degree² (goal >20,000) $(\sim 4\pi/2)$
- Obtain >15 million (goal >25 million) spectra at resolution R~5000
- Obtain >1 million (goal >2 million) spectra at resolution R~20,000.

VISTA Telescope @ Paranal

Messenger Nº175







Consortium Surveys (70%)

Milky Way Halo LR Survey

Milky Way Halo HR Survey

Milky Way Disk and Bulge LR Survey

Milky Way Disk and Bulge HR Survey

Galaxy Clusters Survey

AGN Survey

Galaxy Evolution Survey (WAVES)

Cosmology Redshift Survey

Magellanic Clouds Survey

Irwin (IoA), Helmi (RuG)

Christlieb (ZAH)

Chiappini, Minchev, Starkenburg (AIP)

Bensby (LU), Bergemann (MPIA)

Finoguenov (MPE)

Merloni (MPE)

Driver (UWA), Liske (UHH)

Kitaura (AIP), Richard (CRAL), Kneib (EPFL)

Cioni (AIP)

Time-Domain Extragalactic Survey (TIDES) Sullivan (Southampton)

Community Surveys (30%)

- 1 Toloza, O. et al. The White Dwarf Binary Survey (WDB)
- Sacco, G. G. et al. The 4MOST Survey of Young Stars (4SYS)
- 3 Ibata, R. et al. 4MOST Gaia RR Lyrae Survey (4GRoundS)
- 4 Lucatello, S. et al. Stellar Clusters in 4MOST
- Pawlak, M. et al. Spectroscopic Discovery of Binaries with Dormant Black Holes
- Skúladóttir, Á. et al. The 4MOST Survey of Dwarf Galaxies and their Stellar Streams (4DWARFS)
- lovino, A. et al. Stellar Population Survey Using 4MOST (4MOST-StePS)
- Buncan, K. et al. Optical, Radio Continuum and HI Deep Spectroscopic Survey (ORCHIDSS)
- Gruen, D. et al. 4MOST Complete Calibration of the Colour-Redshift Relation (4C3R2)
- Haines, C. et al. CHANCES: A CHileAN Cluster galaxy Evolution Survey
- Bauer, F. E. et al. Chilean AGN/Galaxy Extragalactic Survey (ChANGES)
- 12 Krogager, J.-K. et al. The 4MOST–Gaia Purely Astrometric Quasar Survey (4G-PAQS)
- Peroux, C. et al. Transform our Understanding of the Baryon Cycle with High-Resolution Quasar Spectroscopy (ByCycle)
- 14 Taylor, E. N. et al. The 4MOST Hemisphere Survey of the Nearby Universe (4HS)
- Collett, E. T. et al. The 4MOST Strong Lensing Spectroscopic Legacy Survey (4SLSLS)

Messenger Nº190

4MOST Community Survey time: ~ 2 Mhrs of low-resolution spectroscopy for a representative sample of variability and SED selected AGN

Variability selection:

Currently ZTF + La Silla-Quest (LS4)

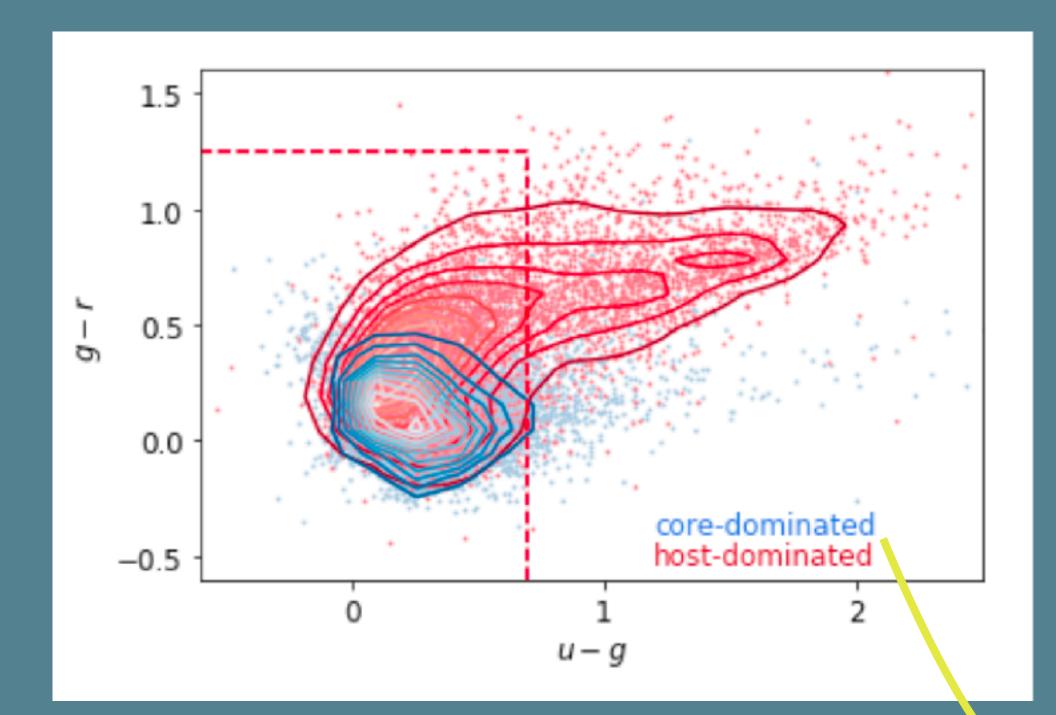
In the future: Rubin LSST survey

SED selection:

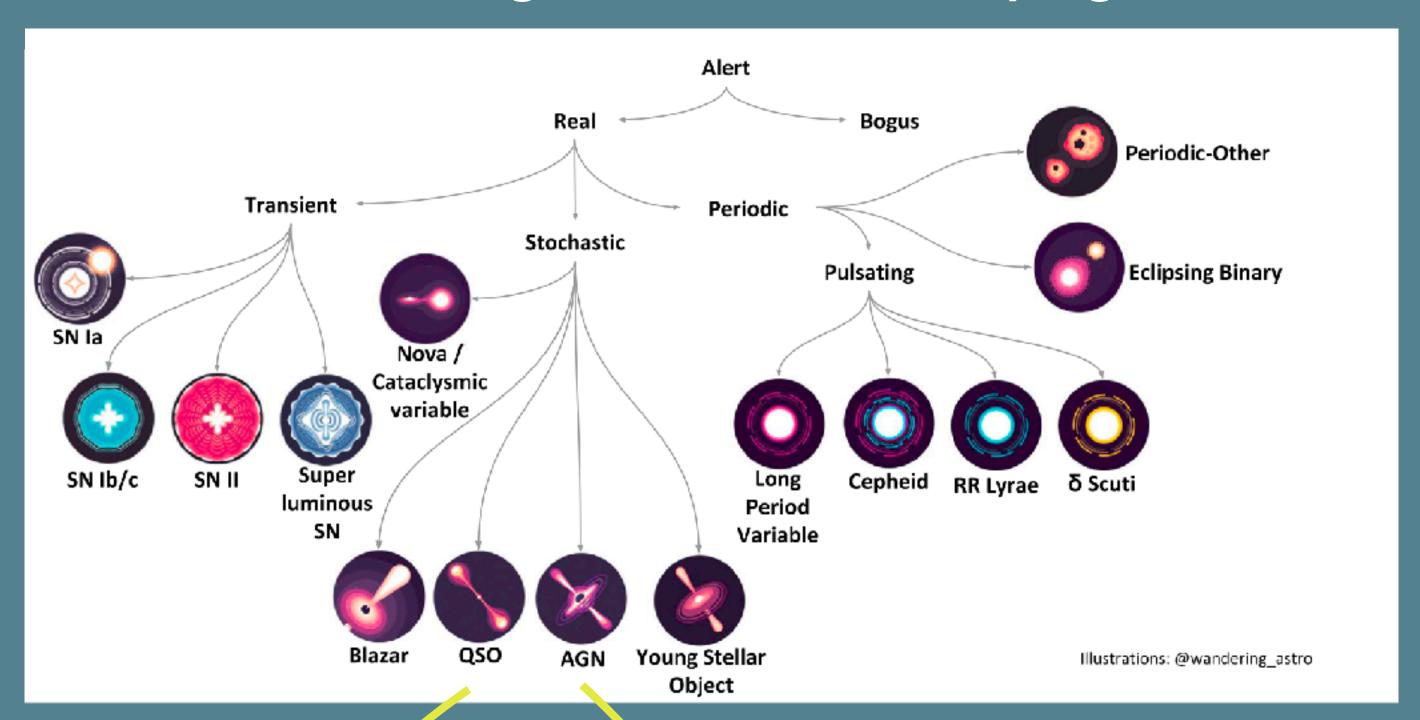
Optical-NIR-MIR SED modeling to look for warm dust

+ Changing Look / State AGN, TDEs, Lensing, Intervening QALs + Repeat Visits

Variability selection:



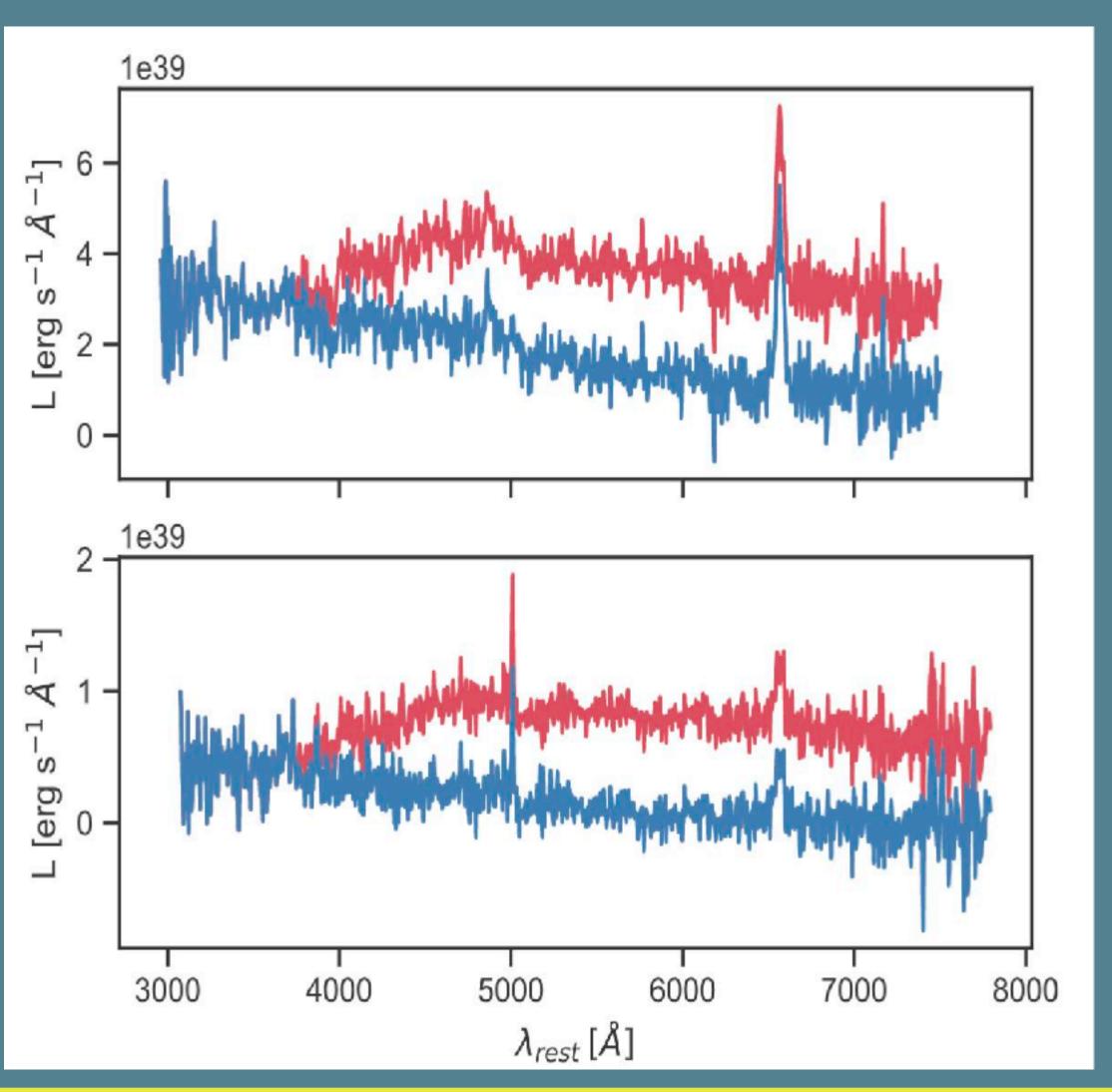
Random forest algorithm to classify light curves



core-dominated host-dominated

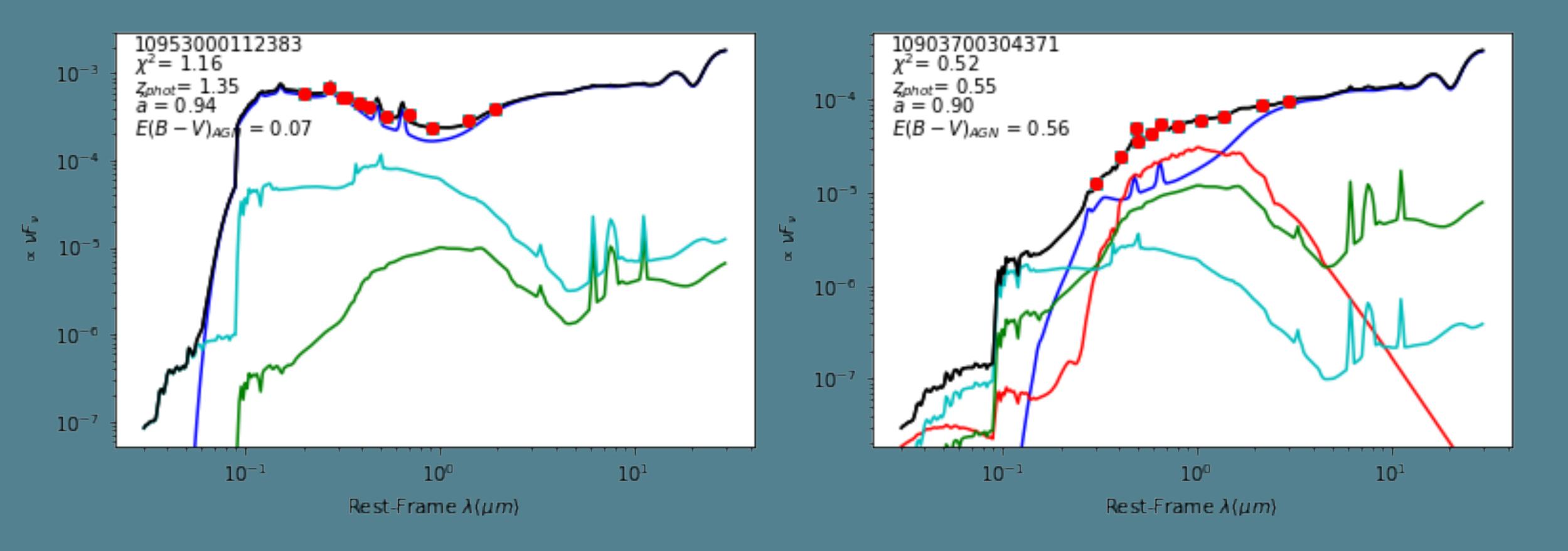
Sánchez-Sáez+2021

Variability selection:

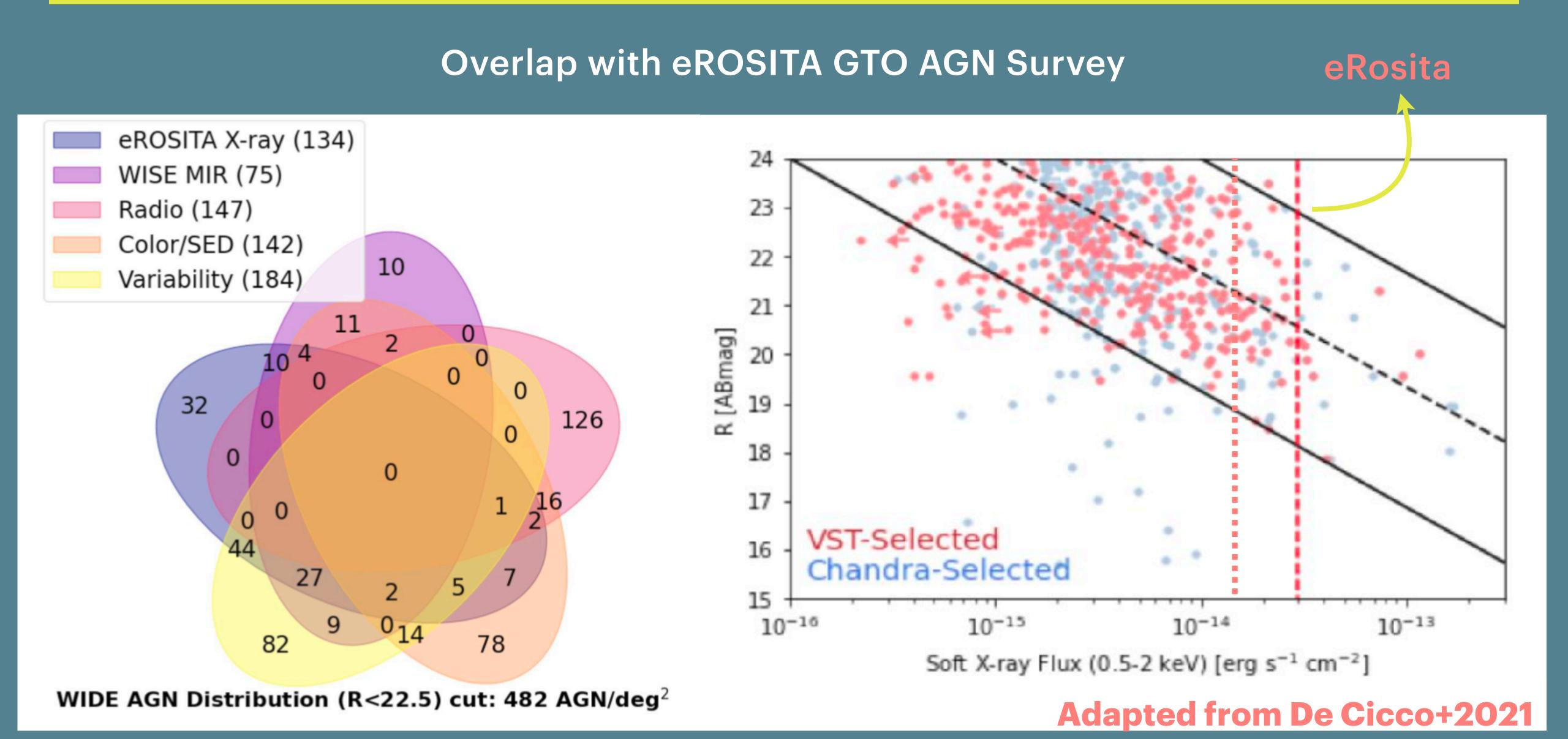


Sánchez-Sáez+2019

SED selection:



DATA: Delve (DECam grizy) - VHS/Viking (VISTA YJHKs) - catWISE (W1 W2)



Main Deliverables:

BH accretion rate densities, evolution and host synergies for moderately accreting AGN ($10^{-4} < L/L$ Edd $< 10^{-1}$) that comprises ~ 50–80% of the estimated total mass accretion onto BHs in type 1 and mildly obscured AGN that strongly complements other 4MOST AGN samples.