

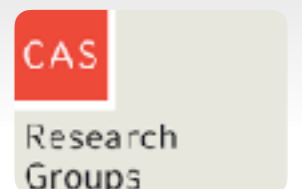
Minimizing systematics with CLONES

(Constrained Local & NEsting Environment Simulations)

Jenny Sorce
and many collaborators

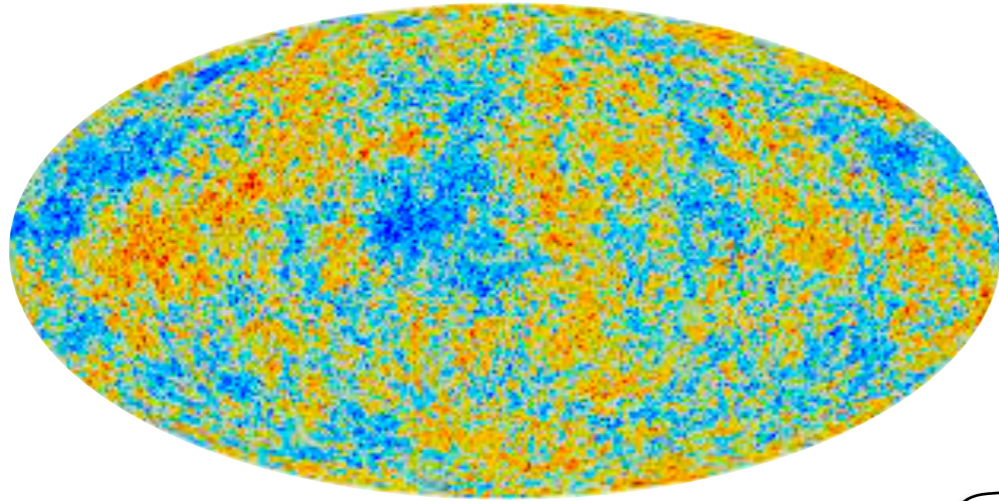
Researcher at CRIStAL, Lille & Associate Researcher at IAS, Orsay &
Guest researcher at AIP, Potsdam & CAS fellow at LMU, Munich

CosmoVerse@Lisbon - May 30th, 2023

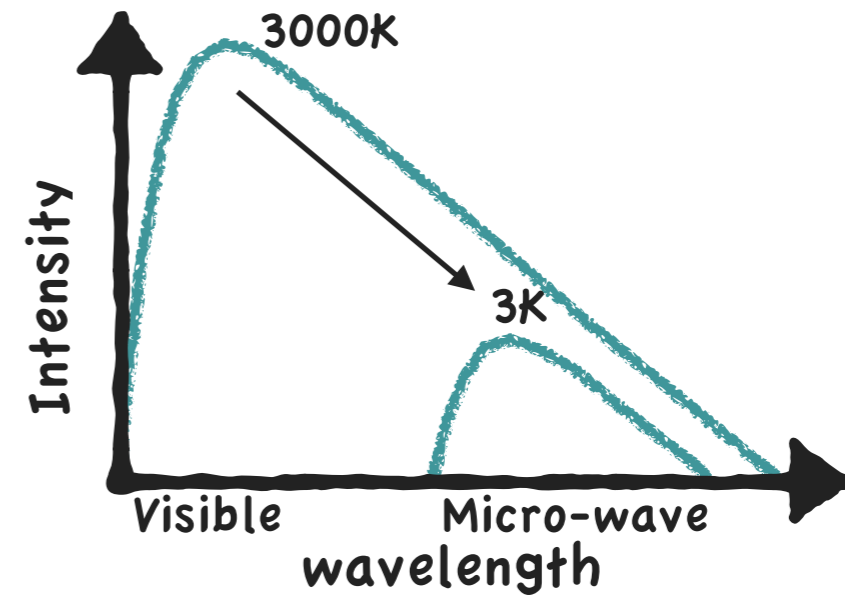


Cosmology: Λ CDM?

Cosmic Microwave Background

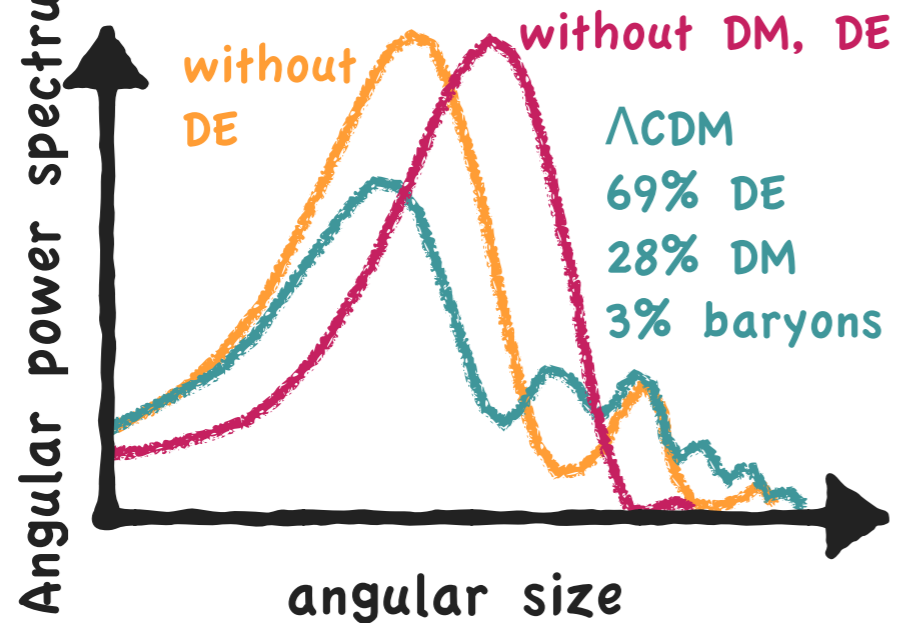


Universe expansion

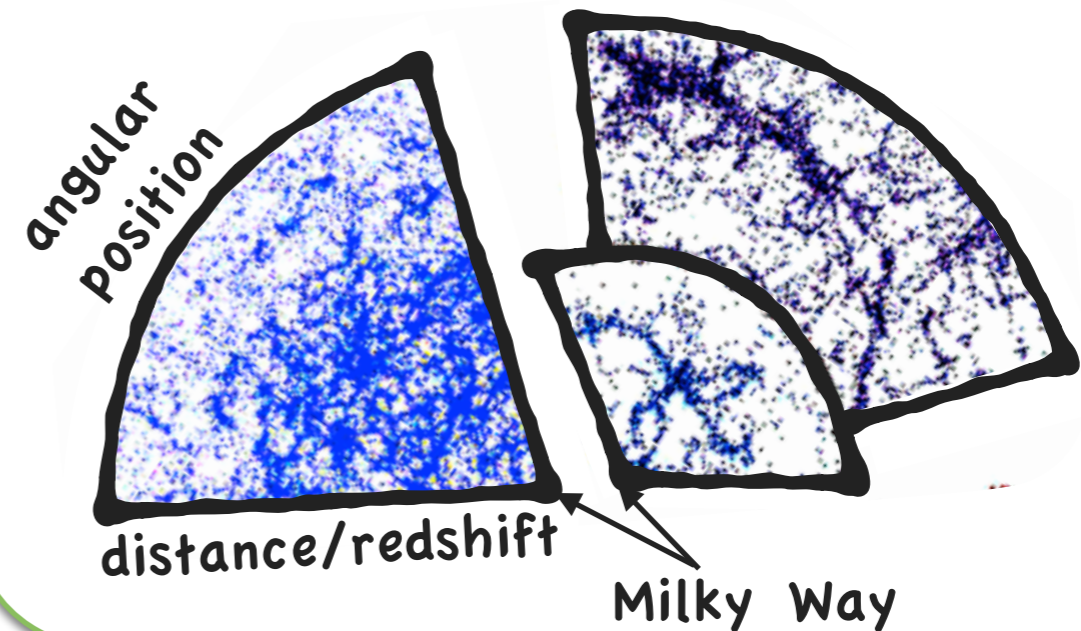


Λ CDM

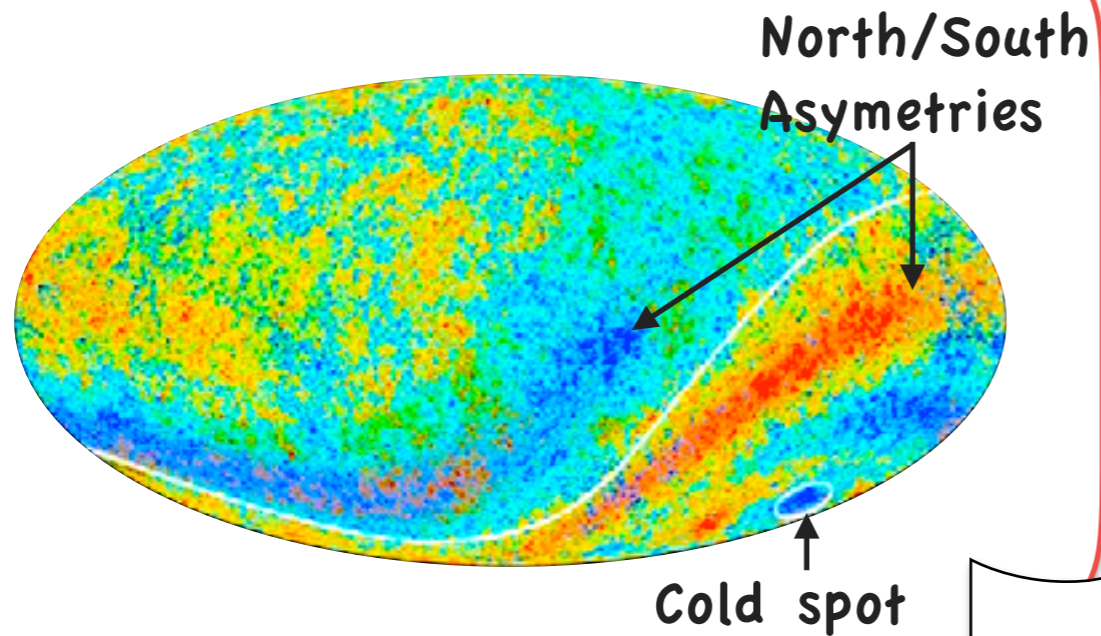
Universe content



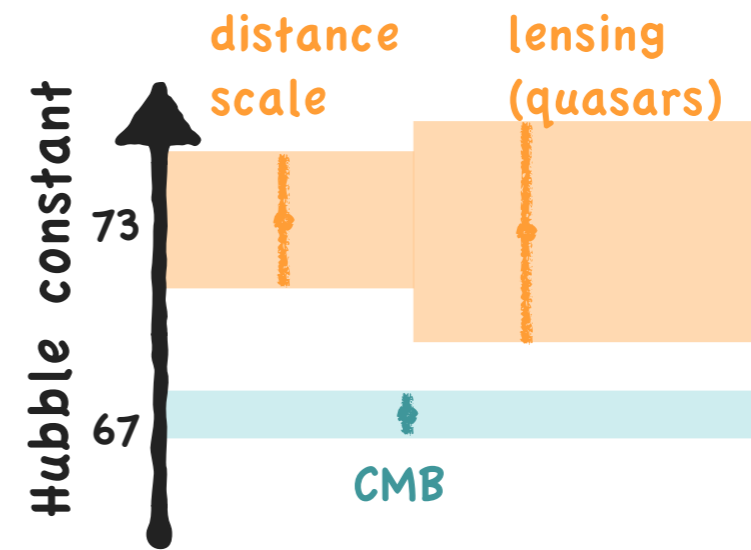
Cosmic Web and galaxies



Anomalies in the CMB

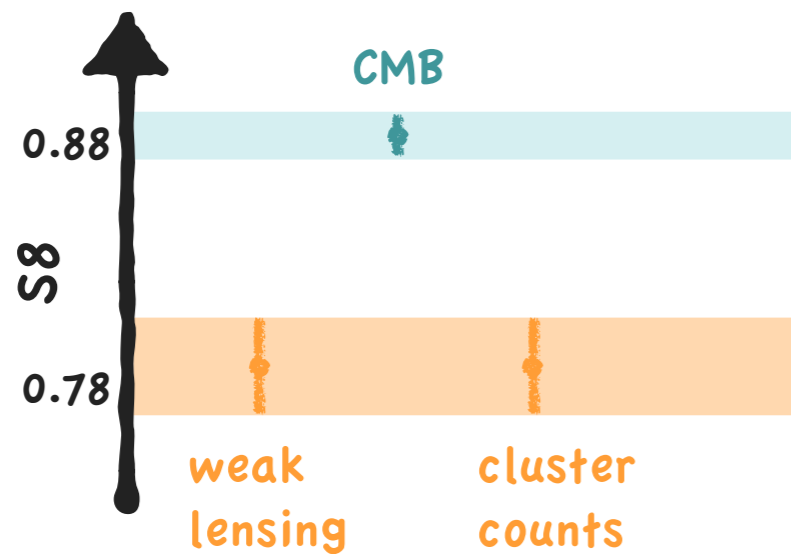


Universe expansion rate (H_0)

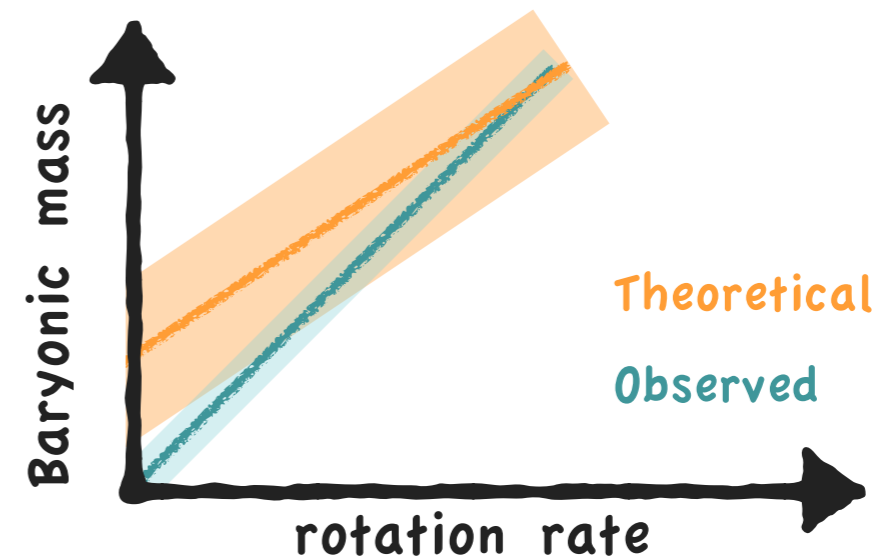


~~**Λ CDM**~~

S8 (σ_8, Ω_m)

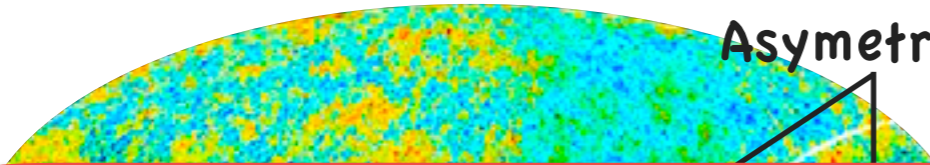


Galaxy properties



Anomalies in the CMB

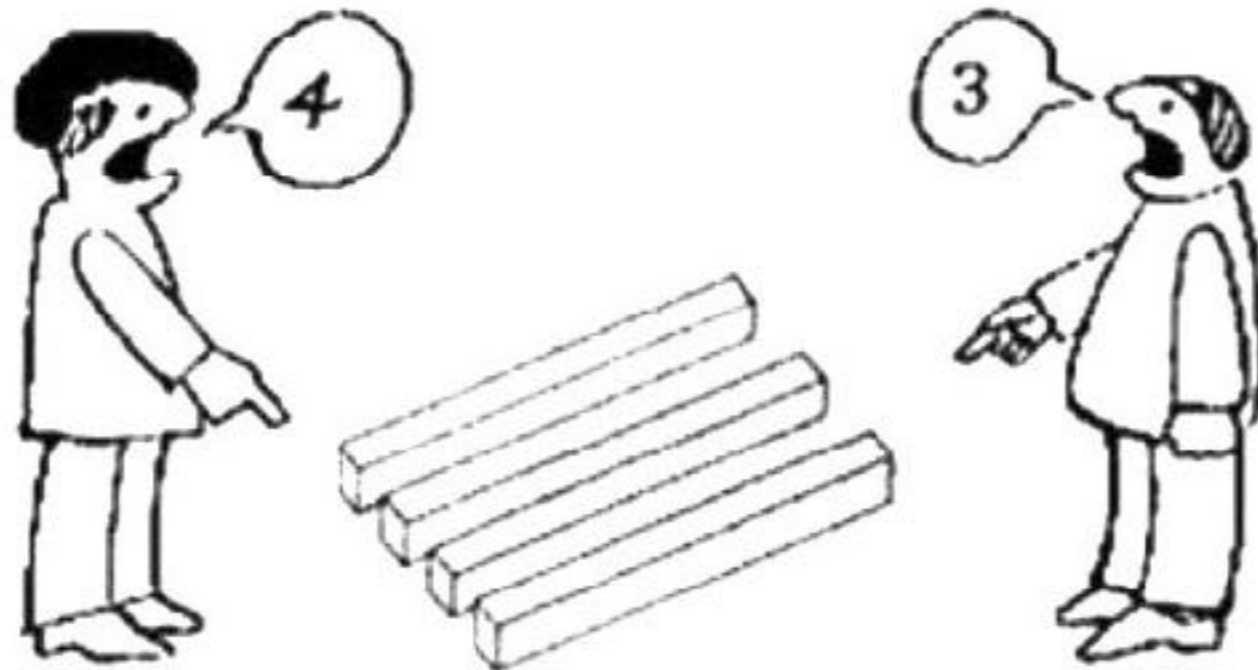
North/South
Asymmetries



Universe expansion rate (H_0)



New physics or biases/systematics?

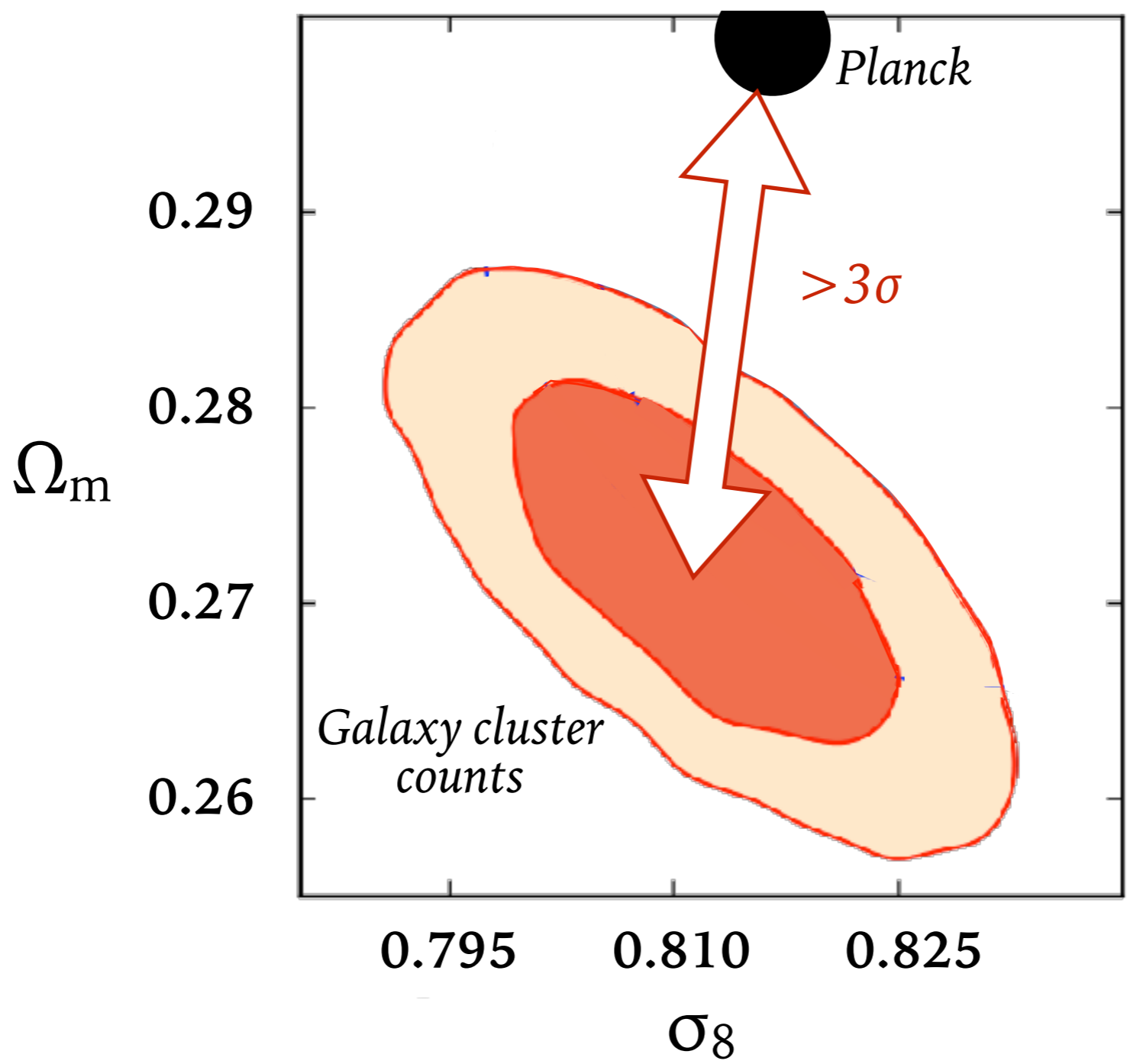


weak
lensing

cluster
counts

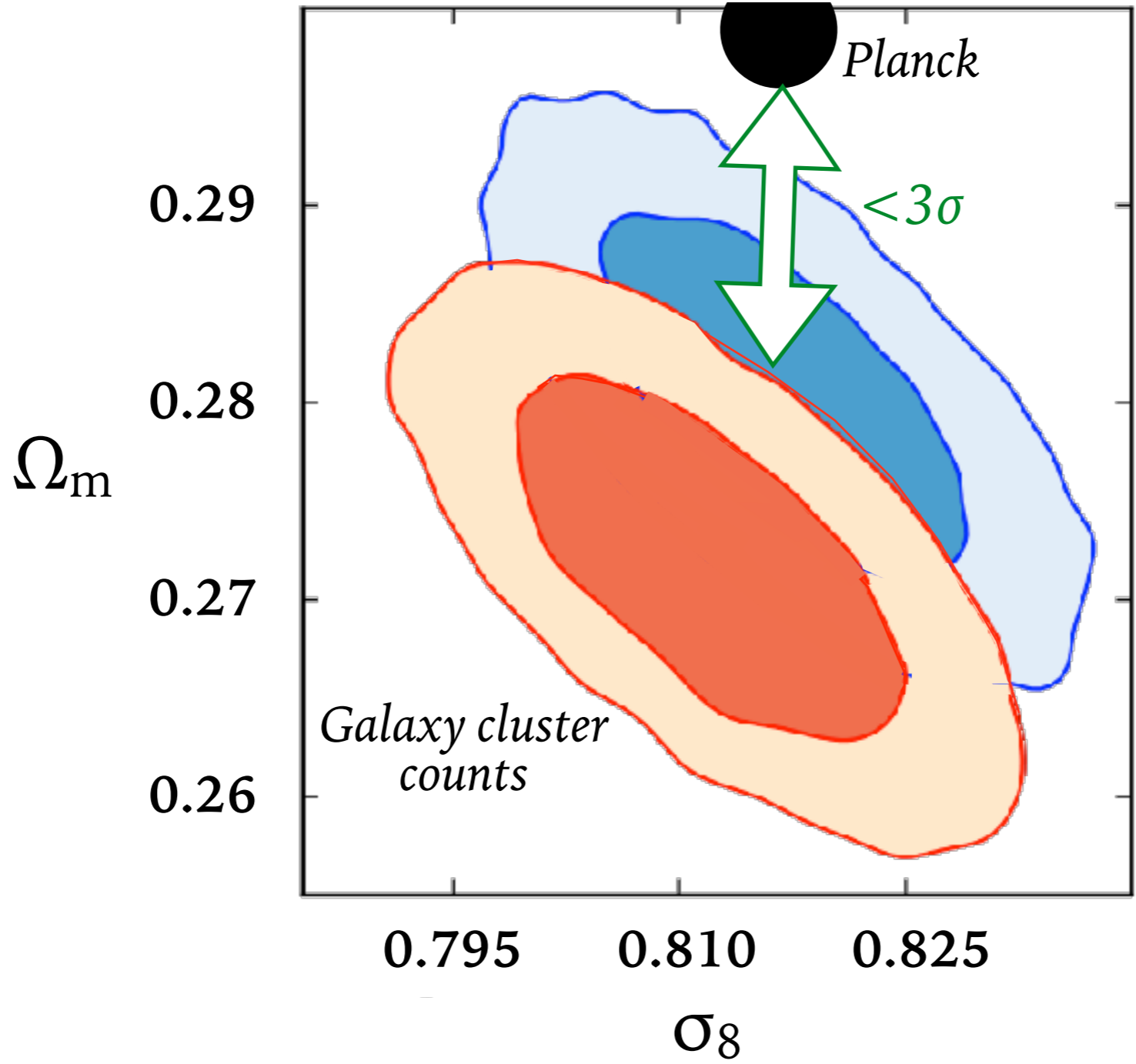


Example of S8 (σ_8, Ω_m) \blacktriangleright Galaxy cluster mass function



Planck Collaboration, Pratt+2018

Example of S8 (σ_8, Ω_m) Changing mass calibration



Planck Collaboration, Pratt+2018

Hydrostatic equilibrium :
intracluster medium

$$\frac{dP}{dr} = - \frac{G\rho M_{HE}}{r^2}$$

Spherical symmetry + no
turbulent/magnetic pressure :

$$\Rightarrow M_{HE}(r) = - \frac{rP_{th}(r)}{G\mu m_p n_e(r)} \frac{d \ln P_{th}(r)}{d \ln r}$$

Gravitational potential well : DM +
Baryons

$$M_{tot} = M_{DM} + M_{gas} + M_{stars}$$



Hydrostatic mass bias

$$\longrightarrow M_{HE} = (1 - b)M_{tot}$$

Hydrostatic equilibrium :
intracluster medium

$$\frac{dP}{dr} = - \frac{G\rho M_{HE}}{r^2}$$

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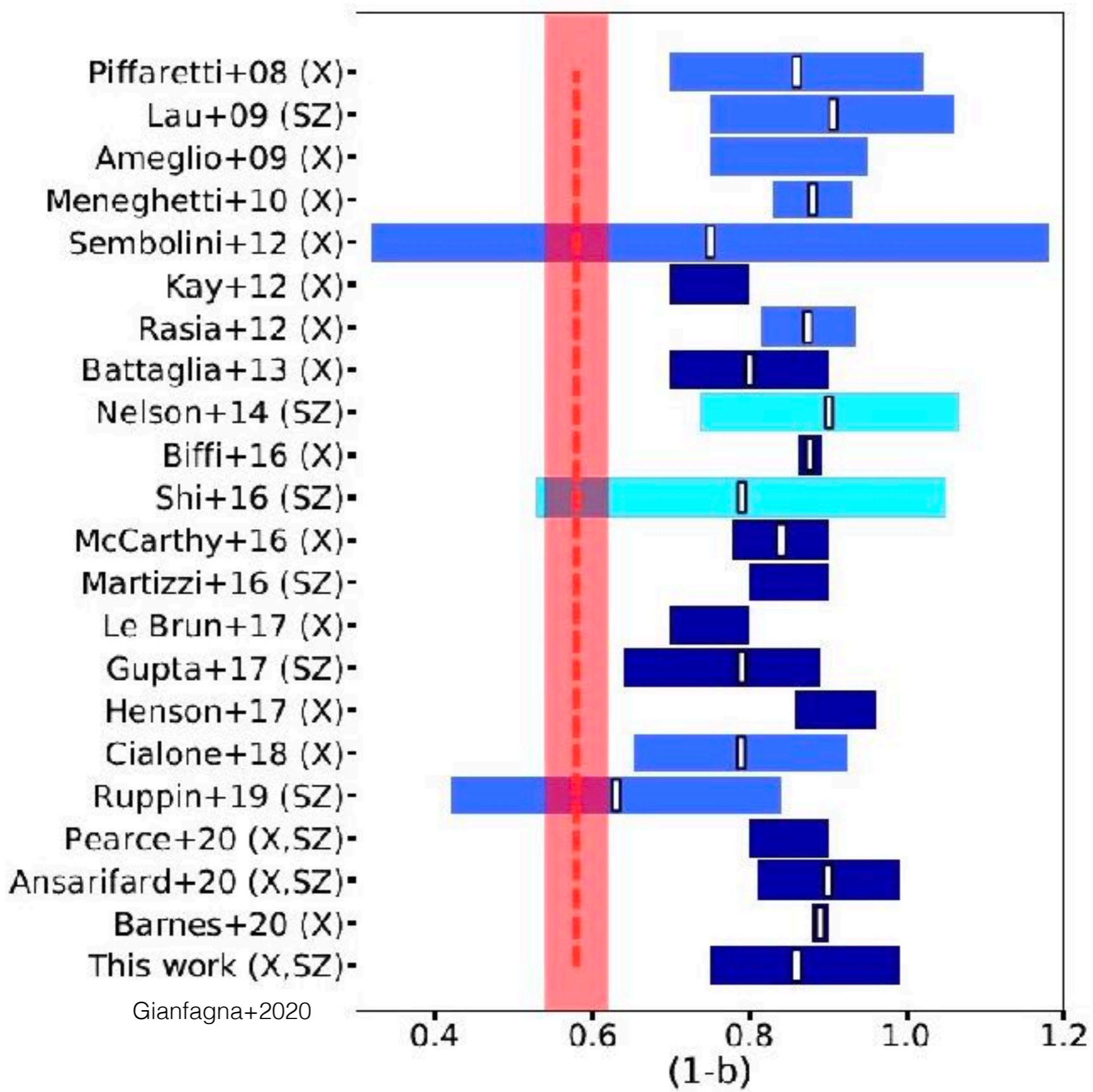


Hydrostatic mass bias

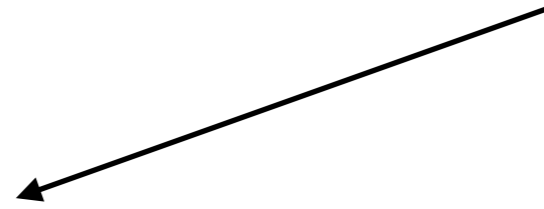
$$\longrightarrow M_{HE} = (1 - b) M_{tot}$$

→ From cosmological simulations

Example of S8 (σ_8, Ω_m) ▶ Huge disparity



$$S8 (\sigma_8, \Omega_m) = X \pm \sigma_{\text{measure}} \pm \sigma_{\text{systematics}}$$

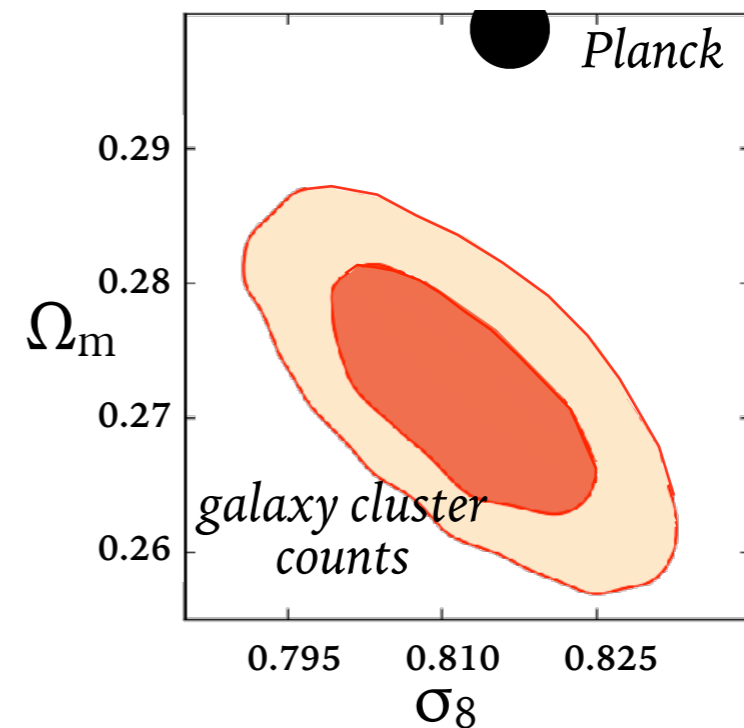
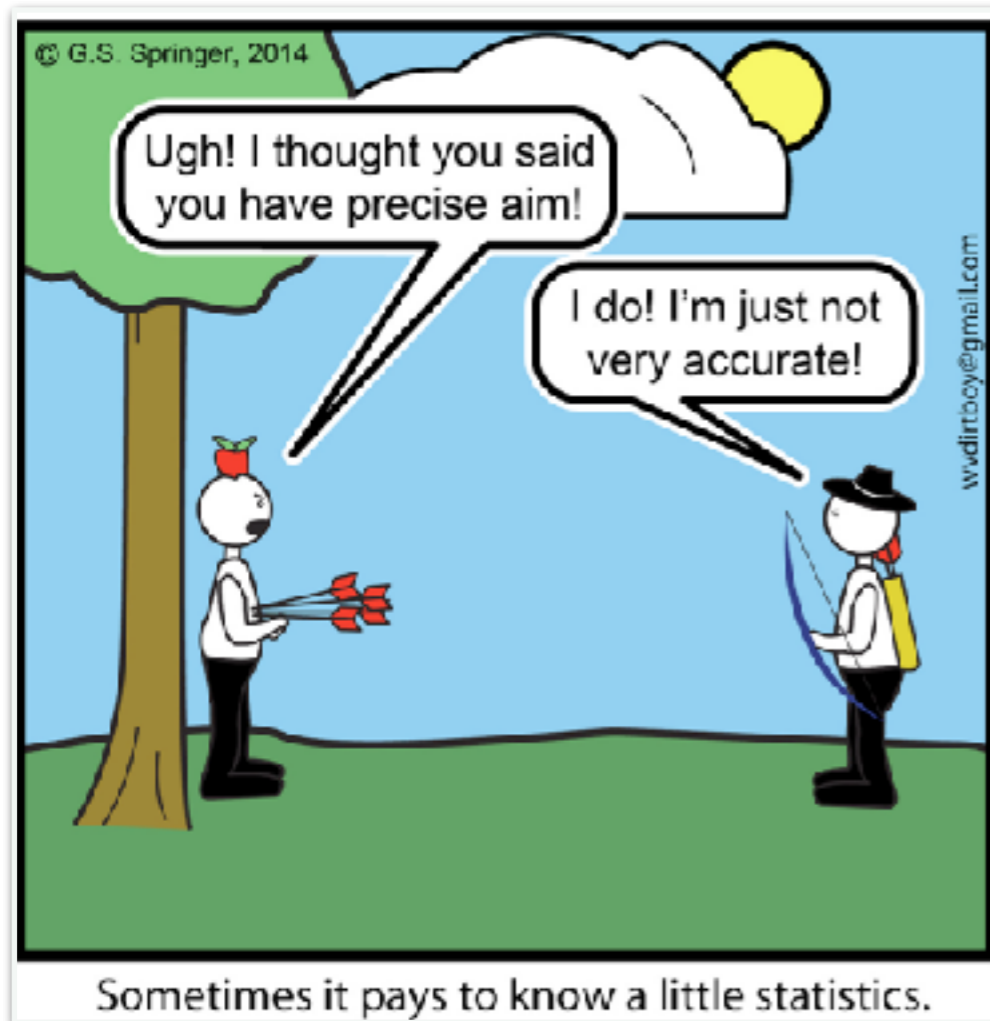


- nb measurements
- instruments/tools sensitivity
= precision

$$S8 (\sigma_8, \Omega_m) = X \pm \sigma_{\text{measure}} \pm \sigma_{\text{systematics}}$$

- nb measurements
- instruments/tools sensitivity
= precision

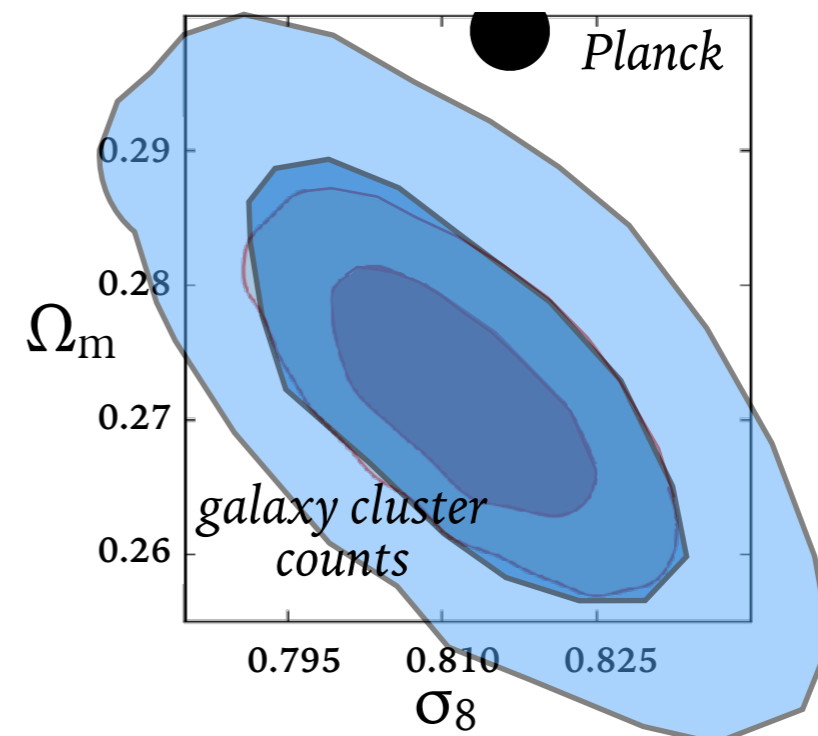
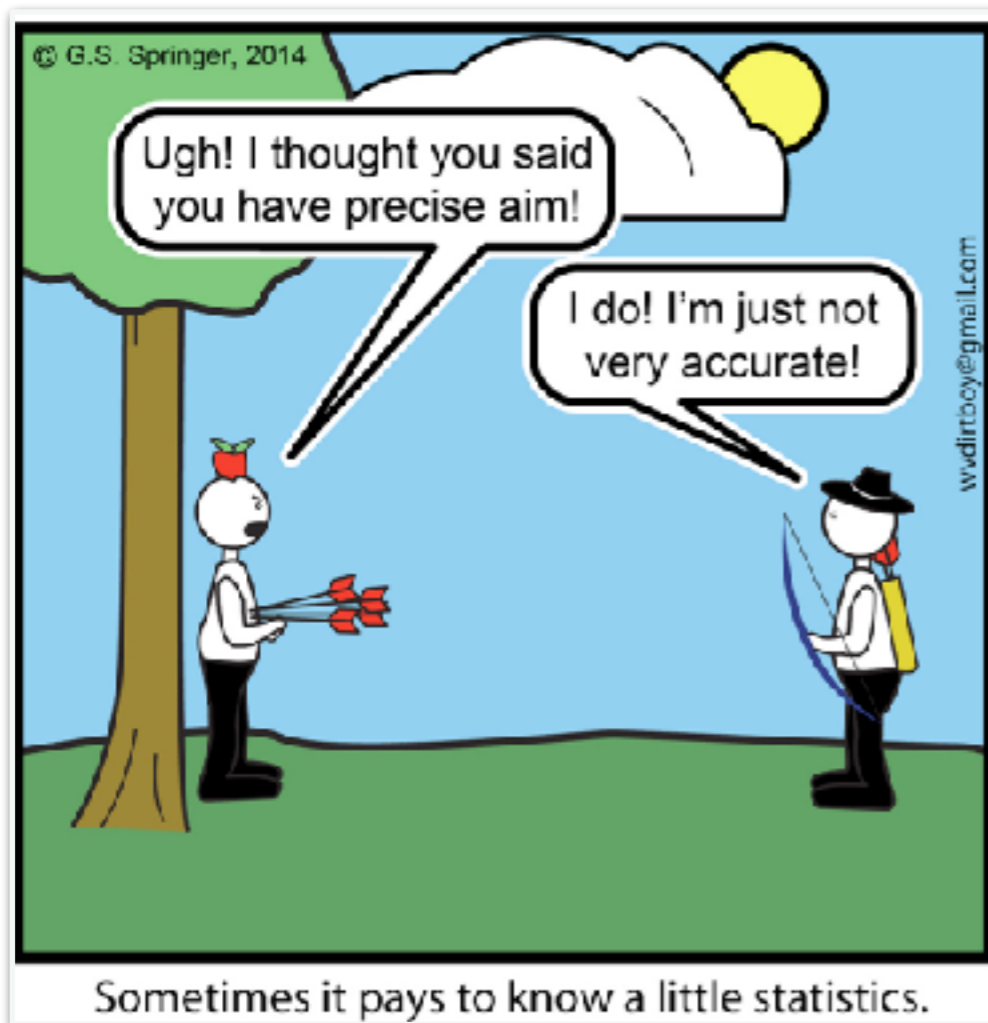
Standard cosmological simulations can give the total uncertainty but cannot reduce the systematics



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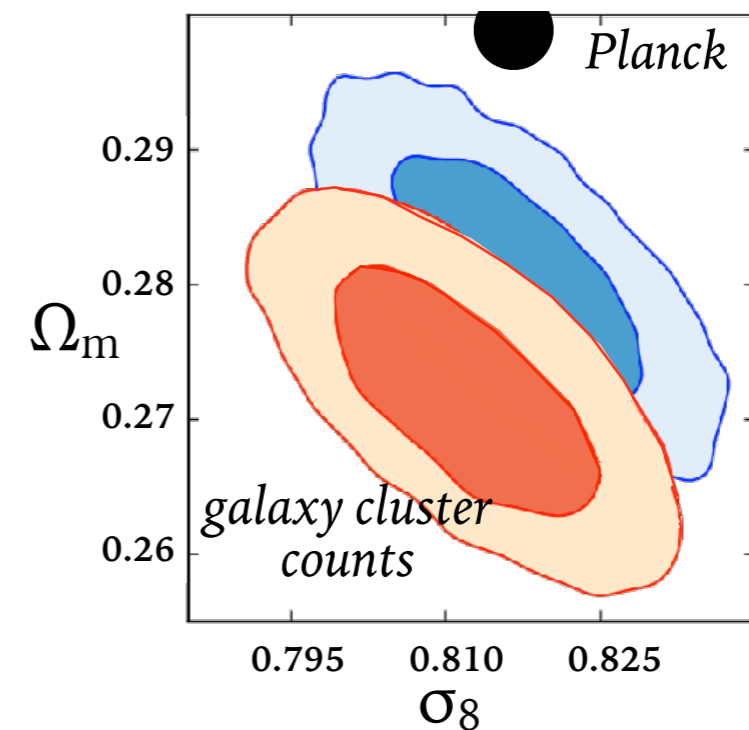
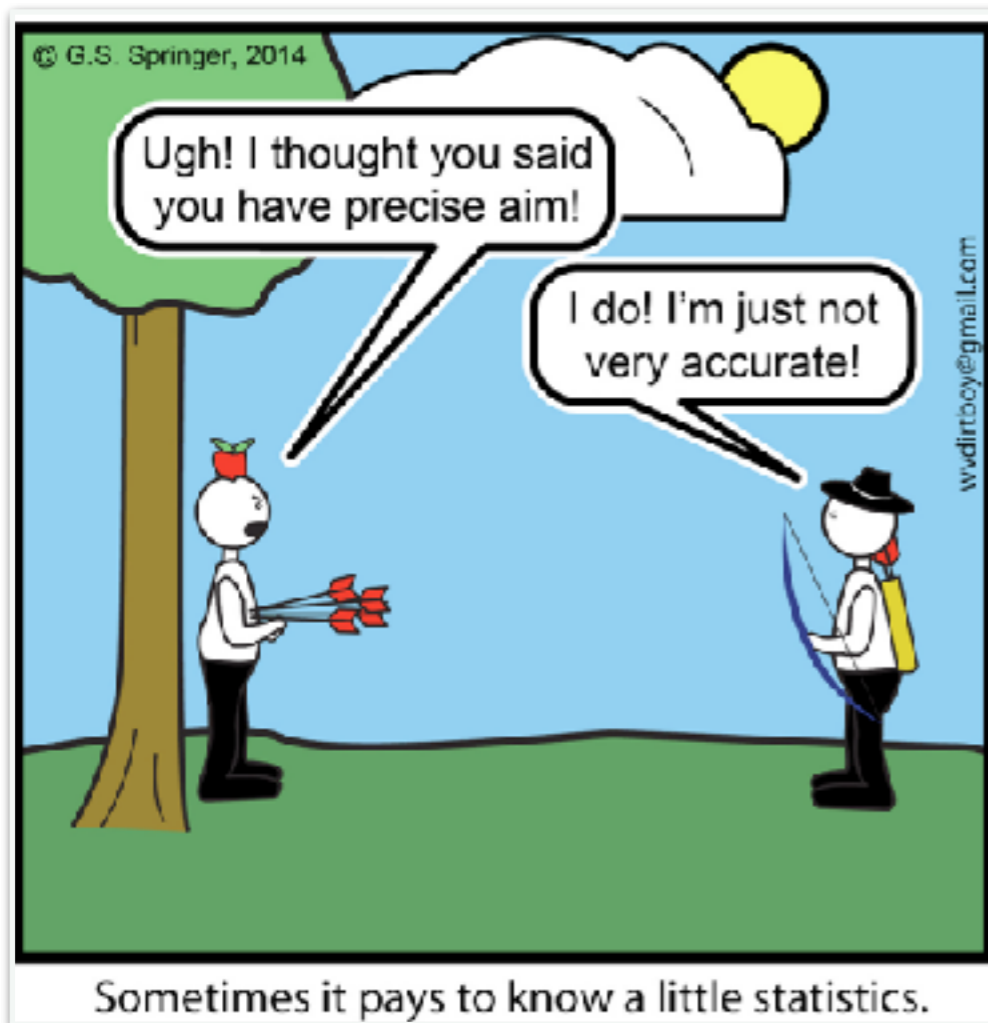
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$$S8 (\sigma_8, \Omega_m) = X \pm \sigma_{\text{measure}} \pm \sigma_{\text{systematics}}$$

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Standard cosmological simulations can give the total uncertainty but cannot reduce the systematics

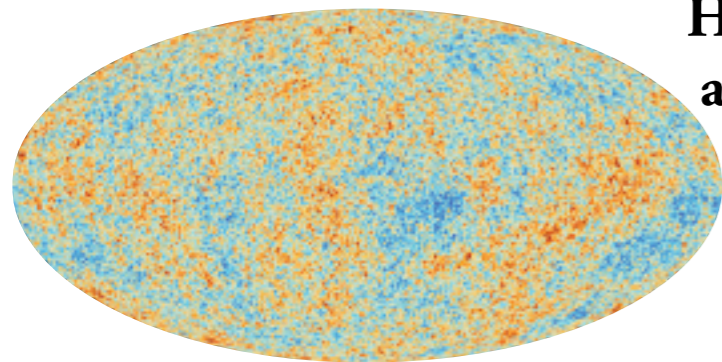


Constrained cosmological simulations can help reduce biases

Standard cosmological simulations

Initial conditions (ICs)

Part of the Universe at
13.7 light-Gyr
Photons received today
have been emitted when it
was ~380 000 yrs. old



Homogeneous
and Isotropic
Universe

→

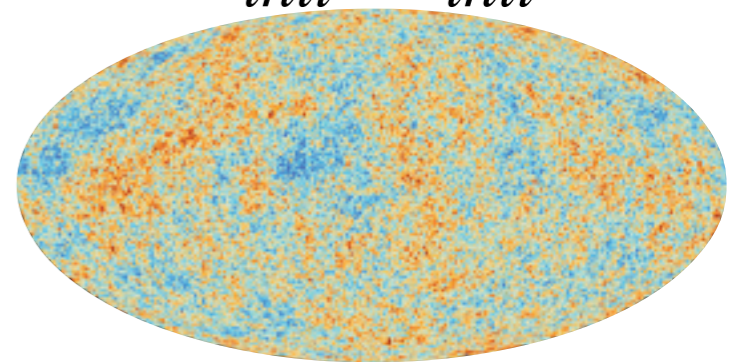
$$P(k)$$

Gaussian
initial density
field

→

$$\delta(\mathbf{k}) = \sqrt{P(\mathbf{k})} \cdot \omega(\mathbf{k})$$

→

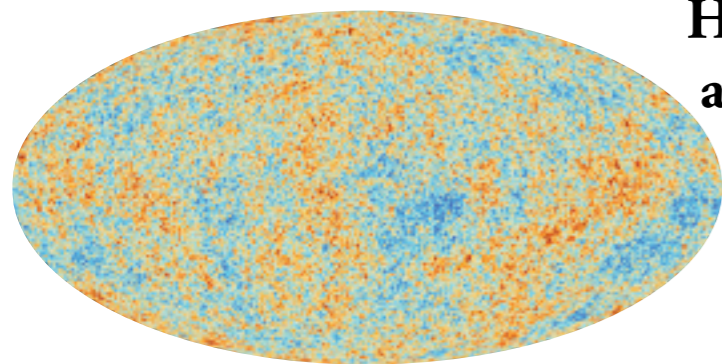


initial conditions of
a random patch of
the Universe
 $\{\delta_{init}, v_{init}\}$

Standard cosmological simulations

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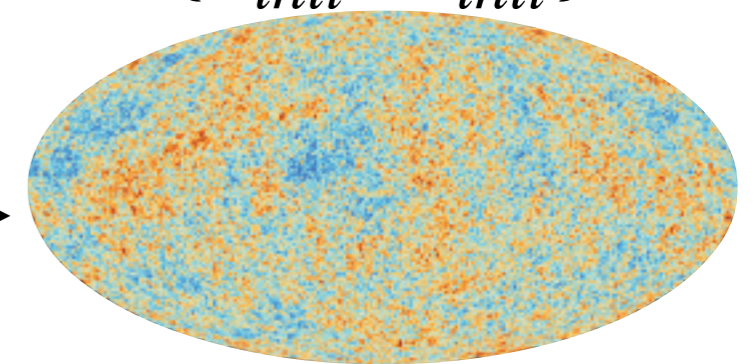
$$P(k)$$

Gaussian
initial density
field

→

$$\delta(\mathbf{k}) = \sqrt{P(\mathbf{k})} \omega(\mathbf{k})$$

→



initial conditions of
a random patch of
the Universe

$$\{\delta_{init}, v_{init}\}$$

Linear perturbation
theory (Euler+
Continuity+Poisson)

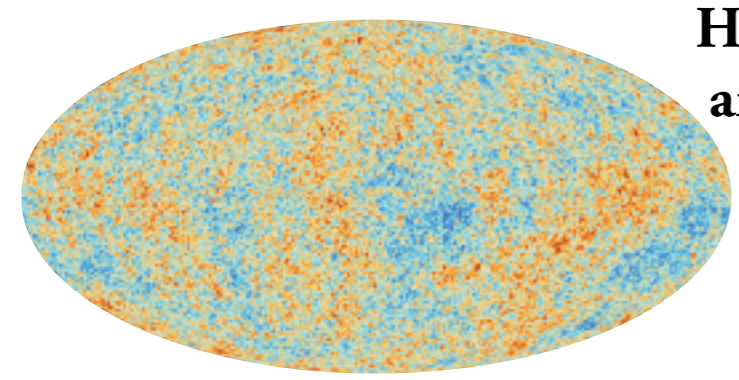
Why only δ ? → $\nabla \cdot v = -\dot{a}f\delta$

NB: only divergent (no tidal) but periodic boundaries

Standard cosmological simulations

Initial conditions (ICs)

Part of the Universe at
13.7 light-Gyr
Photons received today
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Homogeneous
and Isotropic
Universe

→

$$P(k)$$

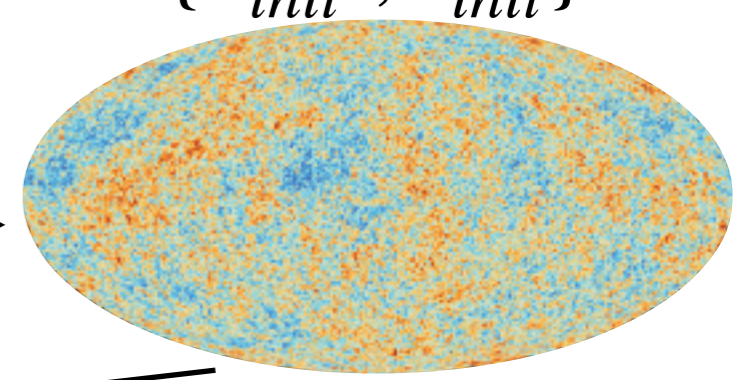
Gaussian
initial density
field

→

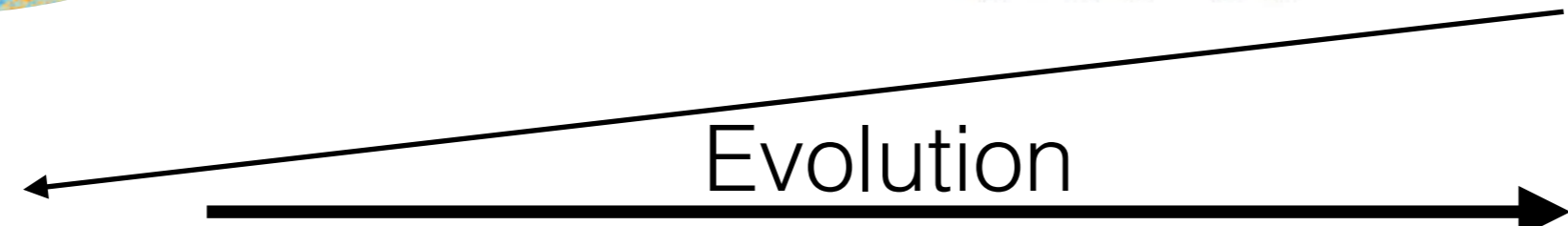
$$\delta(\mathbf{k}) = \sqrt{P(\mathbf{k})} \cdot \omega(\mathbf{k})$$

→

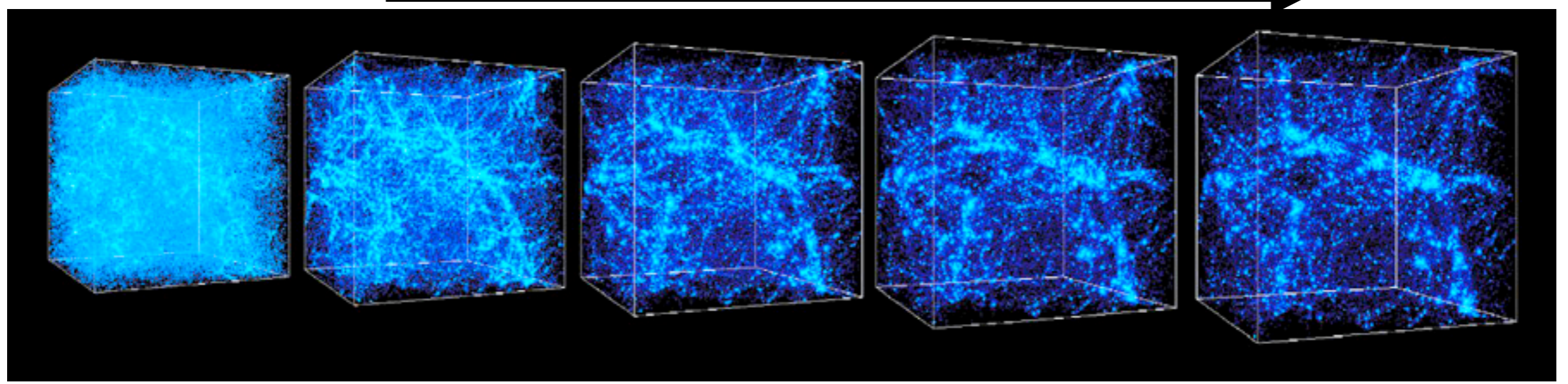
initial conditions of
a random patch of
the Universe
 $\{\delta_{init}, v_{init}\}$



Linear perturbation
theory + "kick"

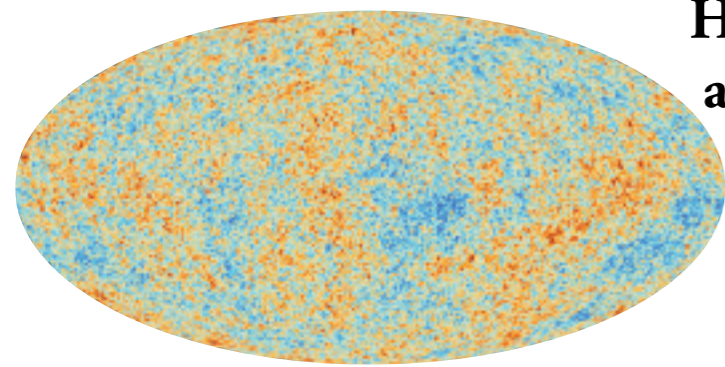


Evolution



Constrained cosmological simulations

Part of the Universe at
13.7 light-Gyr
Photons received today
have been emitted when it
was $\sim 380\,000$ yrs. old



Homogeneous
and Isotropic
Universe

→

$$P(k)$$

Gaussian
initial density
field

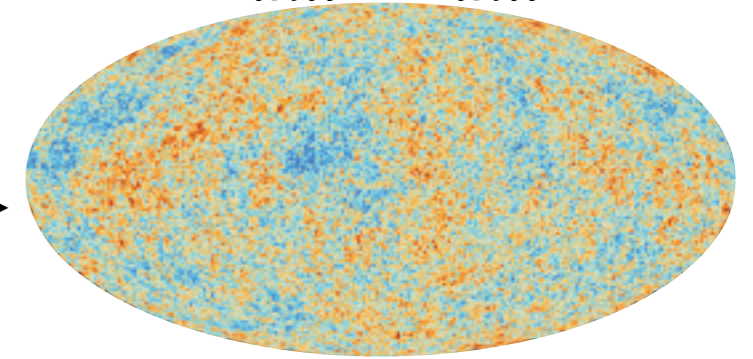
→

$$\delta(\mathbf{k}) = \sqrt{P(\mathbf{k})} \cdot \omega(\mathbf{k})$$

?

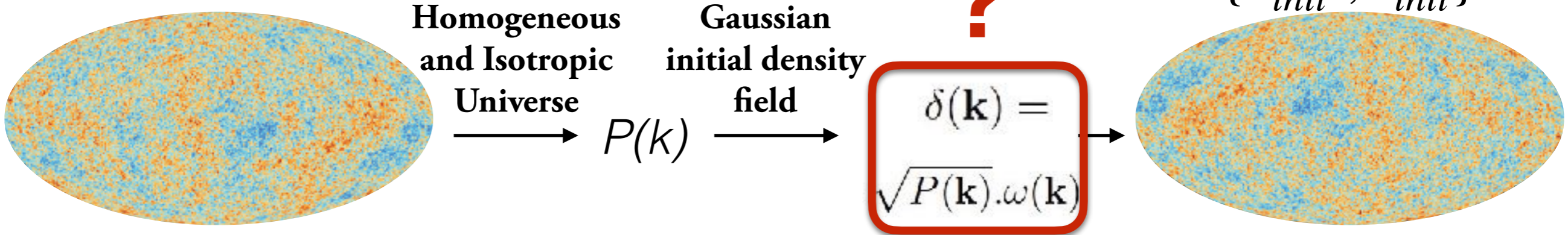
initial conditions of
the local Universe

$$\{\delta_{init}, v_{init}\}$$



Constrained cosmological simulations

Part of the Universe at
13.7 light-Gyr
Photons received today
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Type of constraints

Redshift

Λ CDM

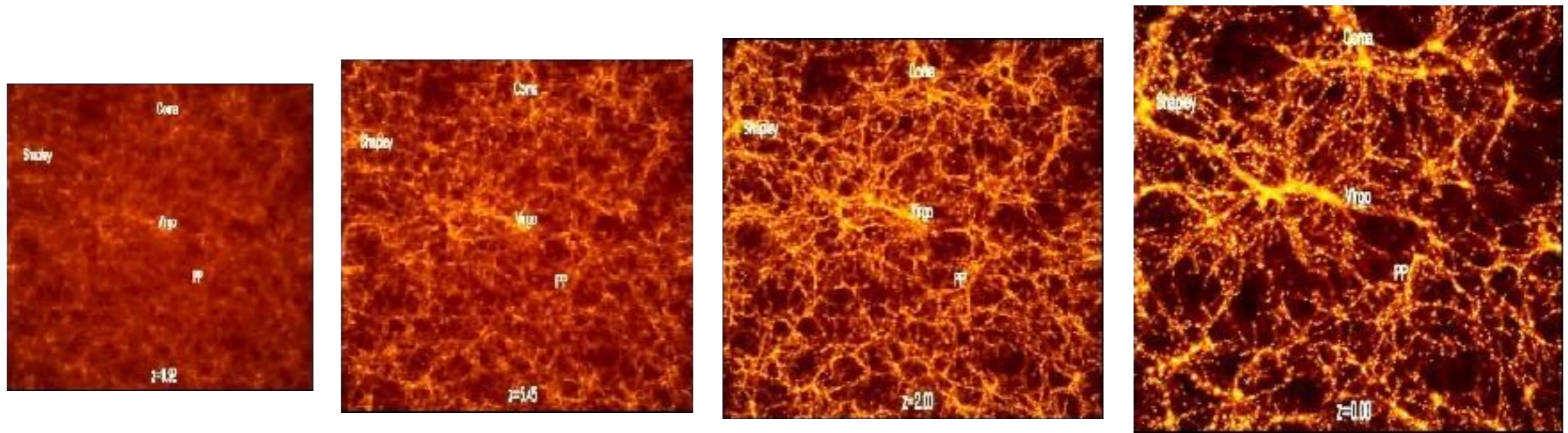
Peculiar velocity

Λ CDM

NB: both with pros and cons!



Constrained cosmological simulations ▶ e.g. CLONES



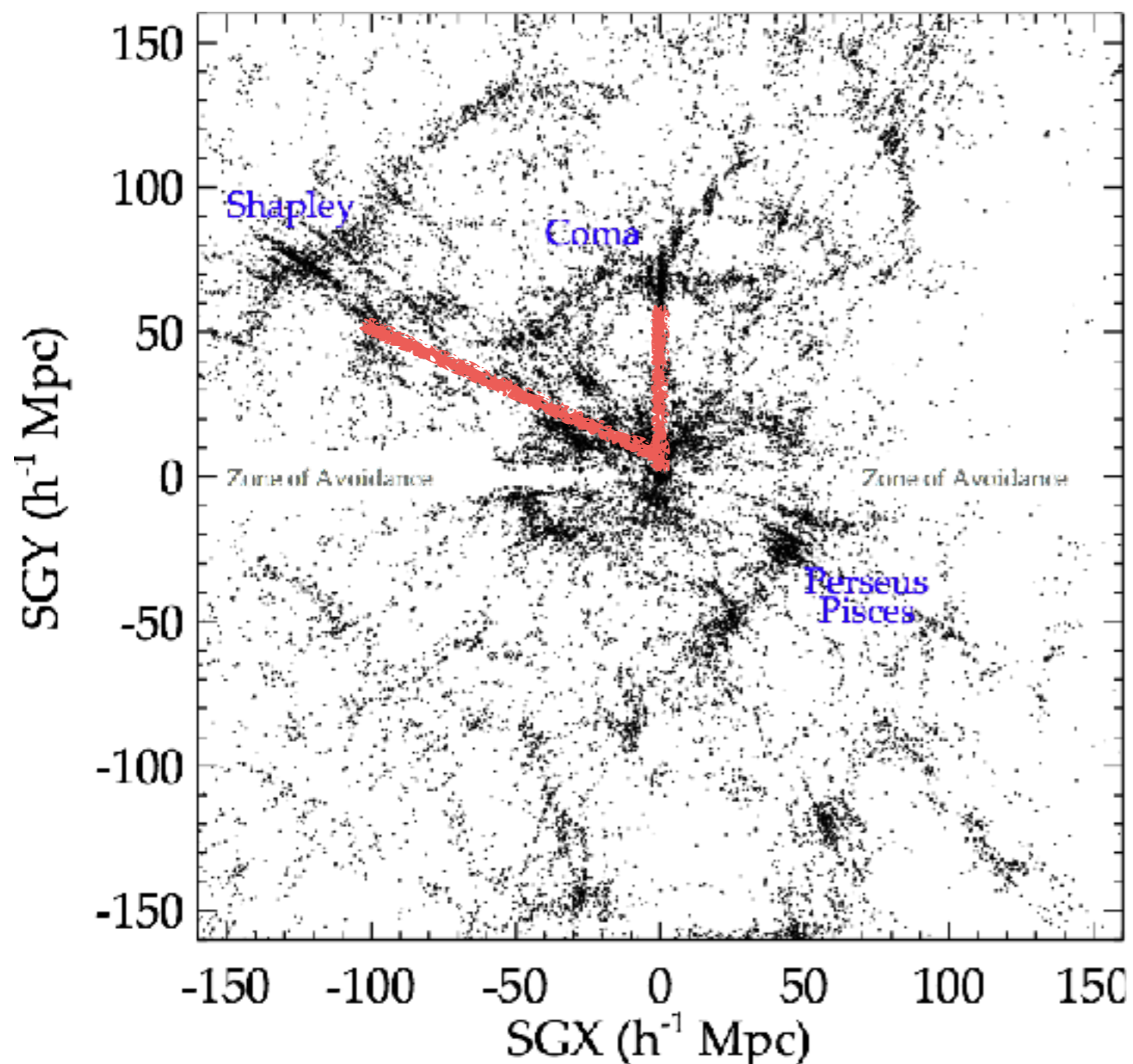
Evolution

Sorce+2016

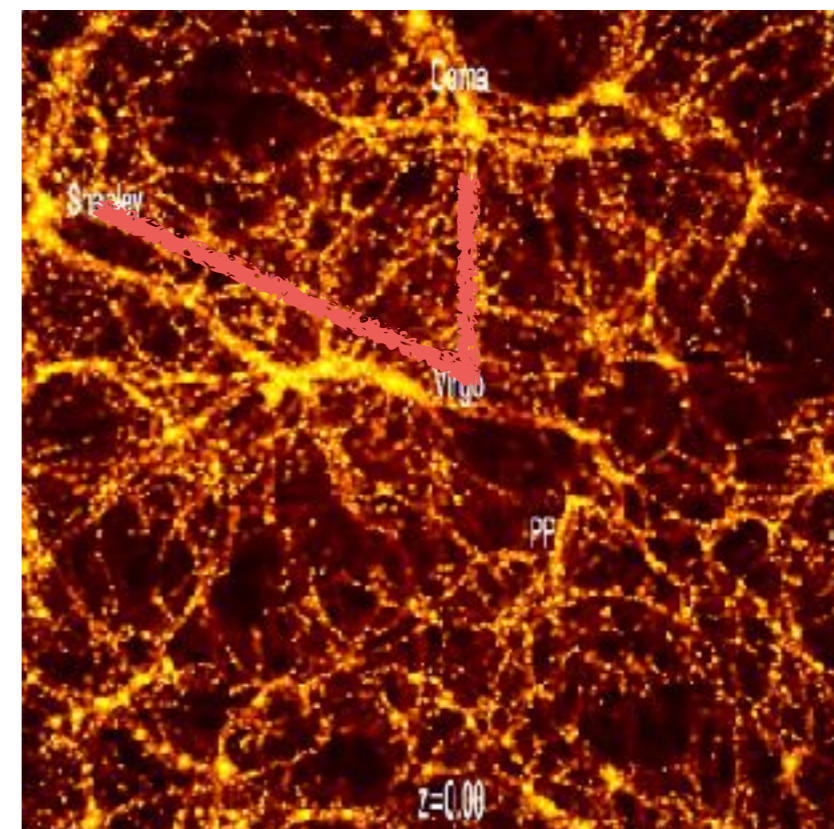
Sorce2018



CLONES = Constrained LOcal & Nesting Environment Simulations

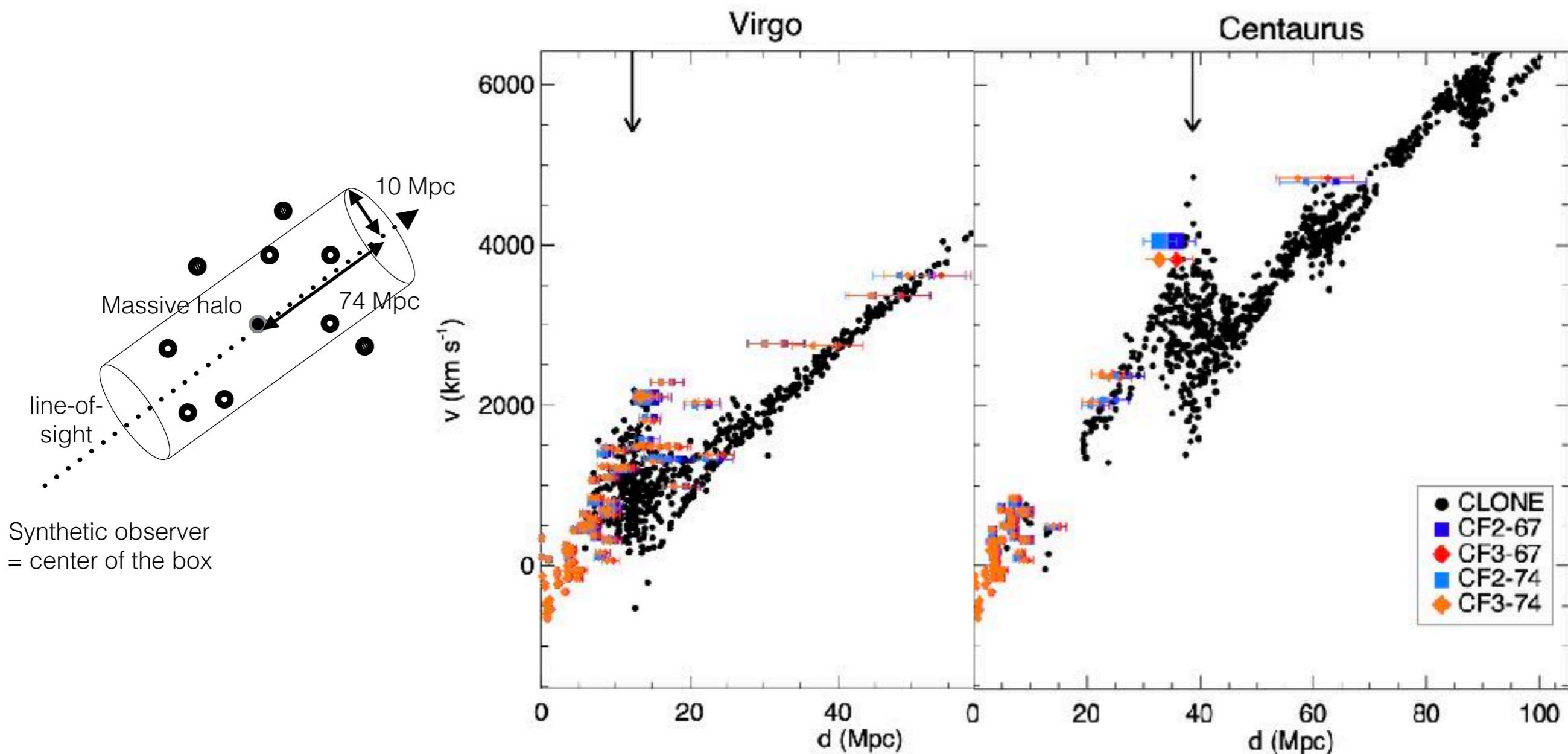


Note the fingers of gods



500 Mpc/h, 1024^3 particles,
DM only, Planck cosmology

Velocity wave signatures in the Hubble diagram



500 Mpc/h, 2048³ particles, DM only, Planck cosmology



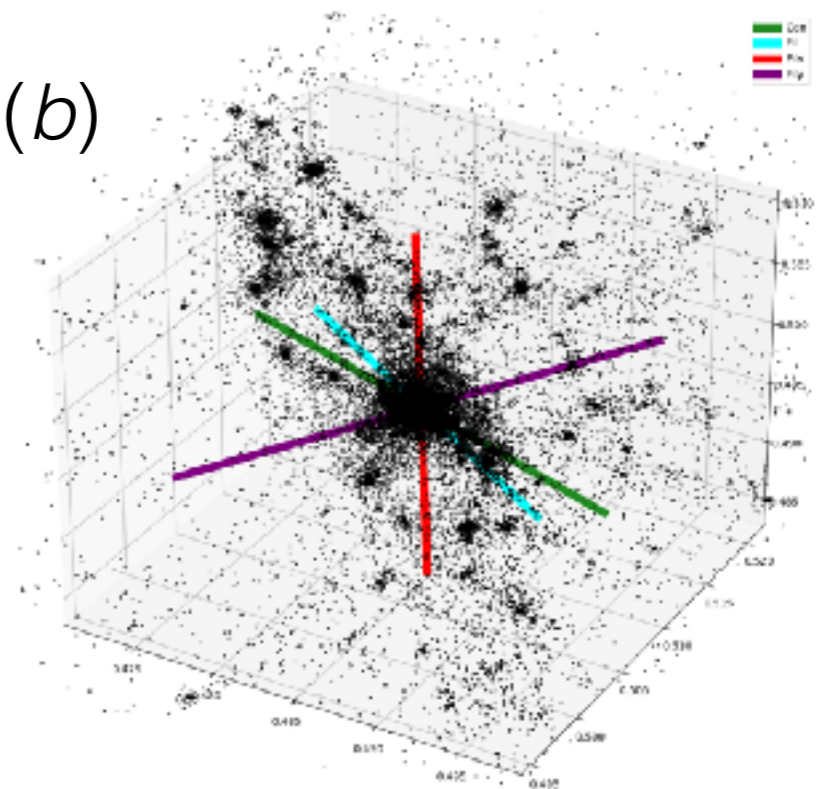
Théo Lebeau

Project: using a CLONE of the local Volume that contains replicas of local clusters to study the impact of

- the dynamical state of the cluster (substructures, morphology)
- the local Environment (connectivity)
- the formation history (accretion from filaments, merging)

on the hydrostatic mass bias (b)

-> *Example of the projection effects on the hydrostatic mass bias in the case of the Virgo cluster*

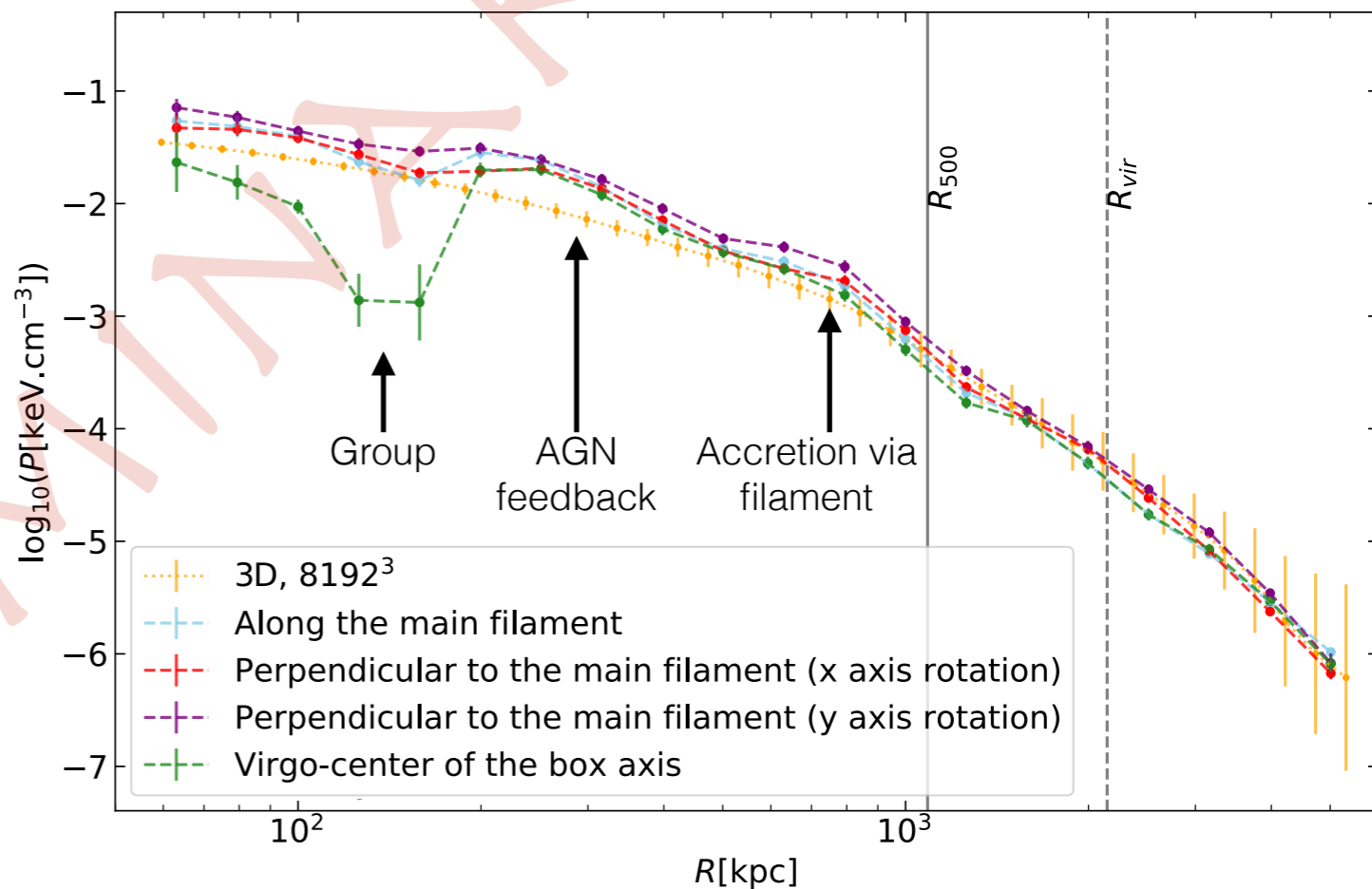
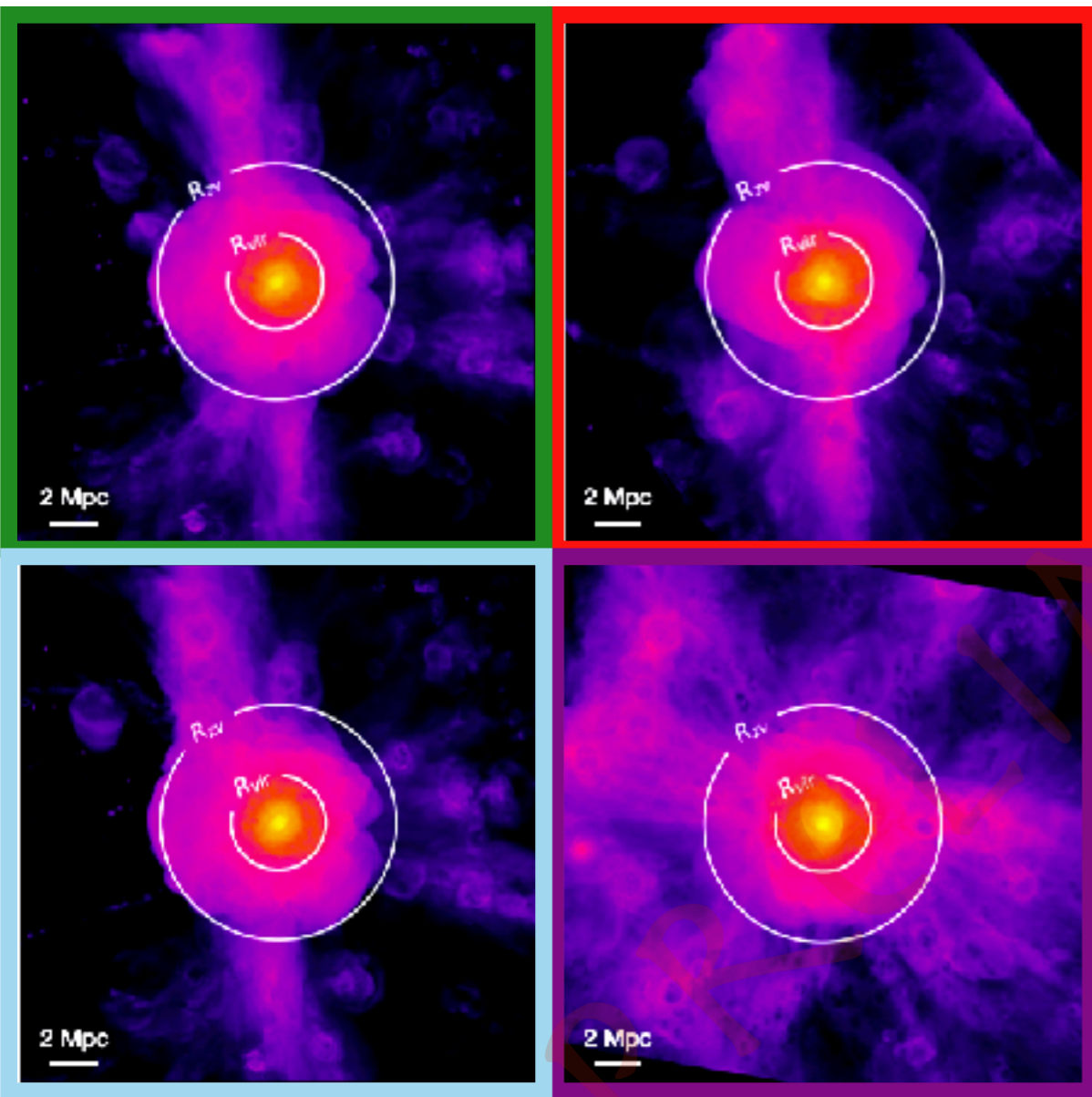


CLONES: hydrostatic mass bias & projection effects



Théo Lebeau

Example of the Virgo galaxy cluster



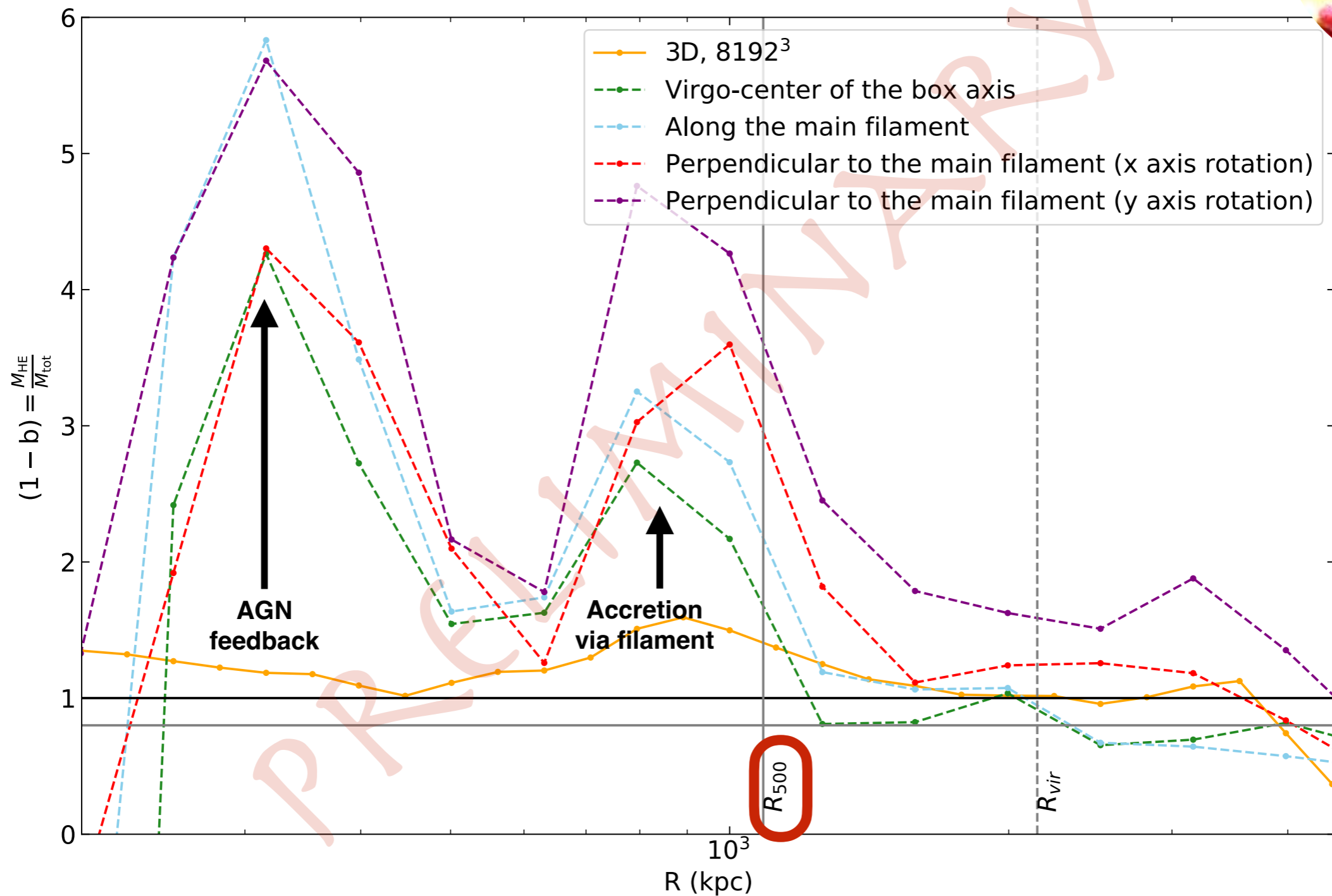
$$M_{HE}(r) = - \frac{r P_{th}(r)}{G \mu m_p n_e(r)} \frac{d \ln P_{th}(r)}{d \ln r} = (1 - b) M_{tot}$$

Lebeau+in prep.

CLONES: hydrostatic mass bias & projection effects



Théo Lebeau

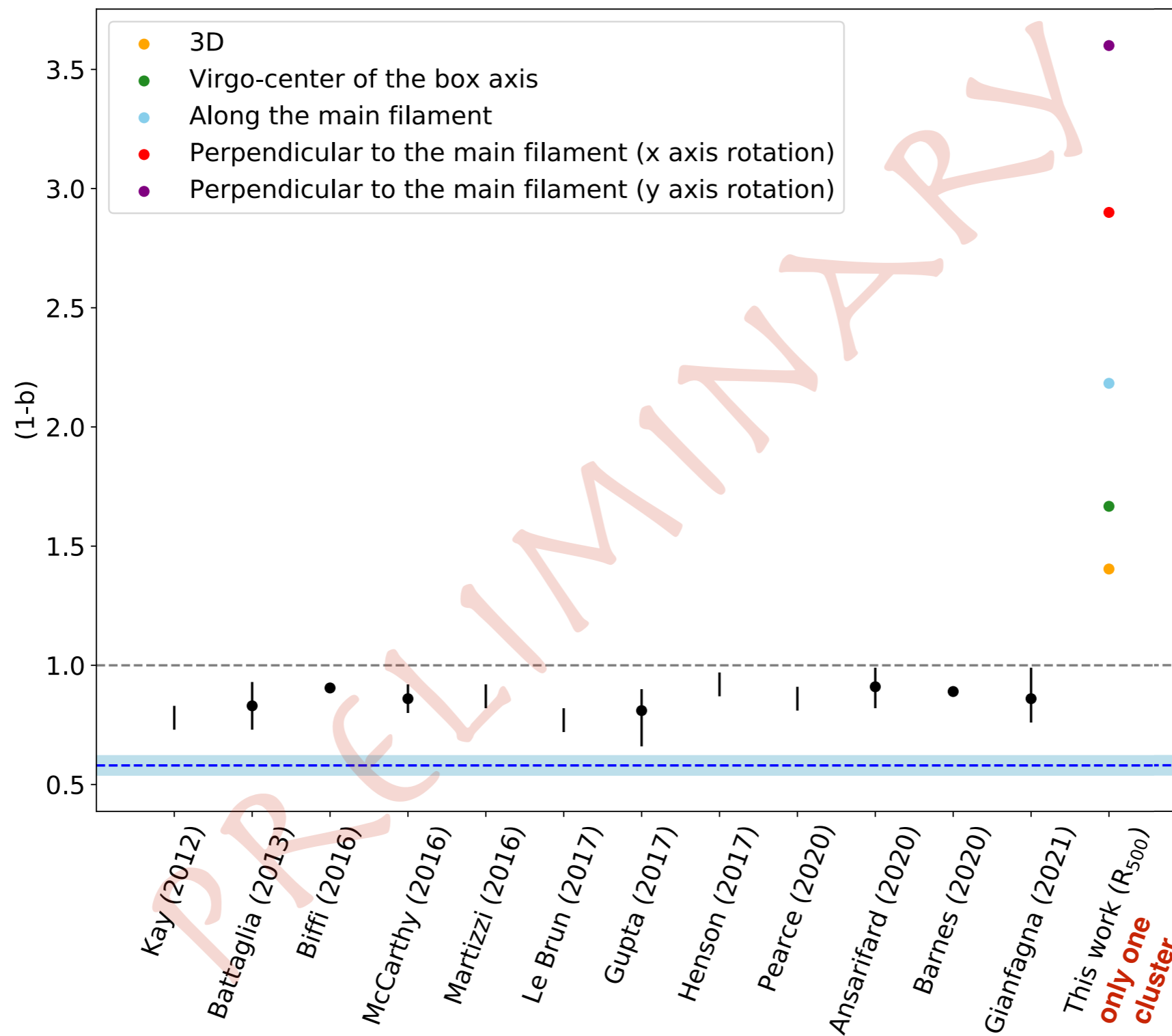


Lebeau+in prep.

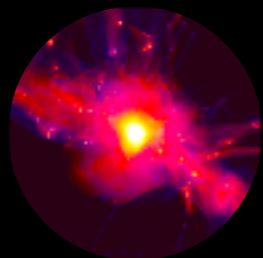
CLONES: hydrostatic mass bias & projection effects



Théo Lebeau

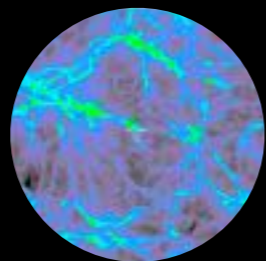


Lebeau+in prep.



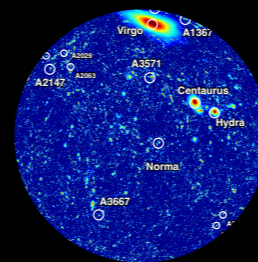
Virgo Cluster

(Sorce+2016, 2019, 2021, in prep., Olchanski & Sorce 2018, Lebeau+ in prep.)



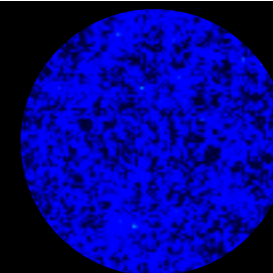
Cosmic Rays in the local Universe

(Hackstein+2018, Boess+in prep.)



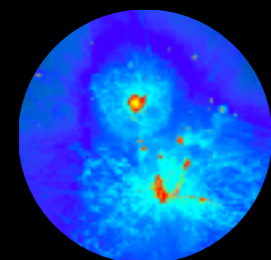
SLOW: local web

(Dolag, Sorce+2023)



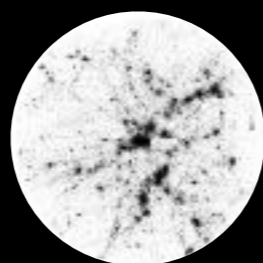
LOCALIZATION: local cluster signatures

(Sorce, Aghanim, Lebeau, Jung, Dolag)



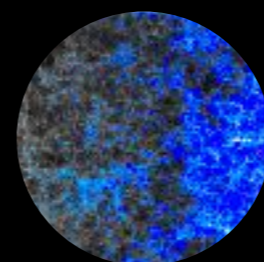
HESTIA: Local Group

Carlesi, Sorce+2016, Carlesi+2016, 2017, Libeskind+2020, Damle+2022, Newton+2022; Luis+2022, Dupuy+2022, Arora+2022, Khoperskov+2022a,b,c



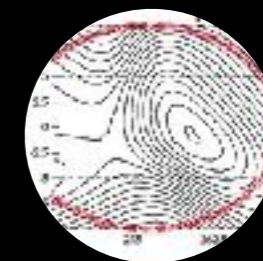
Coma connectivity

(Malavasi, Sorce, Dolag, Aghanim submitted)



CoDa: Reionization of the local Universe

(Ocvirk+2020, Lewis+2020, Gronke+2021, Sorce+2022, Lewis+2022, Park+2022)

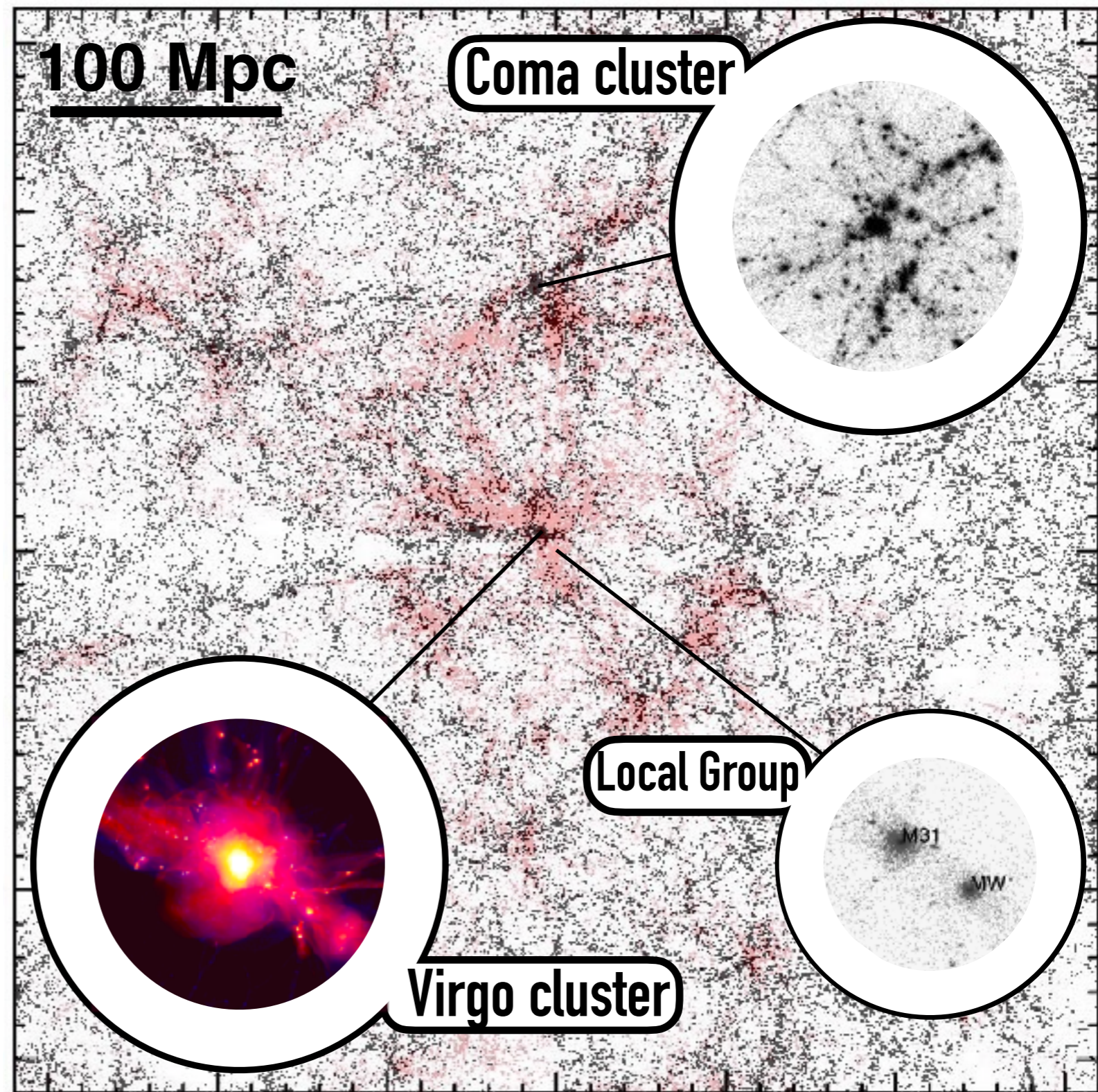


Zone of Avoidance

(Sorce+2017)

Conclusion

- **Standard** cosmological simulations give only the full uncertainty
- **Constrained** cosmological simulations can permit **reducing biases/systematics**
- **CLONES are constrained** cosmological simulations valid down to the cluster scales with induced smaller scales
- CLONES are **widely used** and **maybe you are the next users!**



**Thank you, Merci, Grazie,
Gracias, Danke,
Mahalo, 谢谢, ありがとう,
הודת, Obrigada, Dank u,
Tak, Cảm ơn, Dziękuję, 감사합니다
Kiitos, Aitäh, diolch, dankewol,
ಧನ್ಯವಾದಗಳು, ...***

* Missing your 'thanks' spelling? It means I did not get the chance to learn how to say it so far

